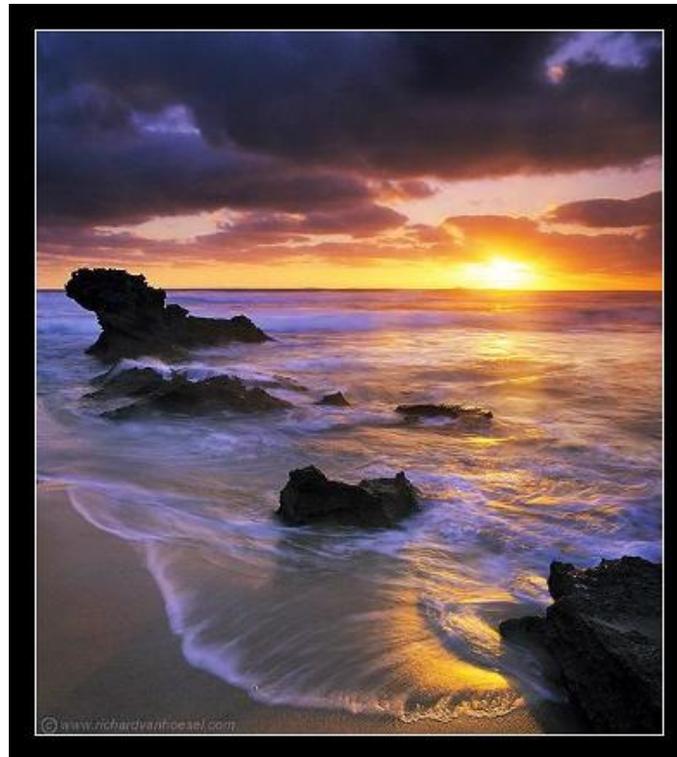


CS6670: Computer Vision

Noah Snavely

Lecture 20: Light, color, and reflectance

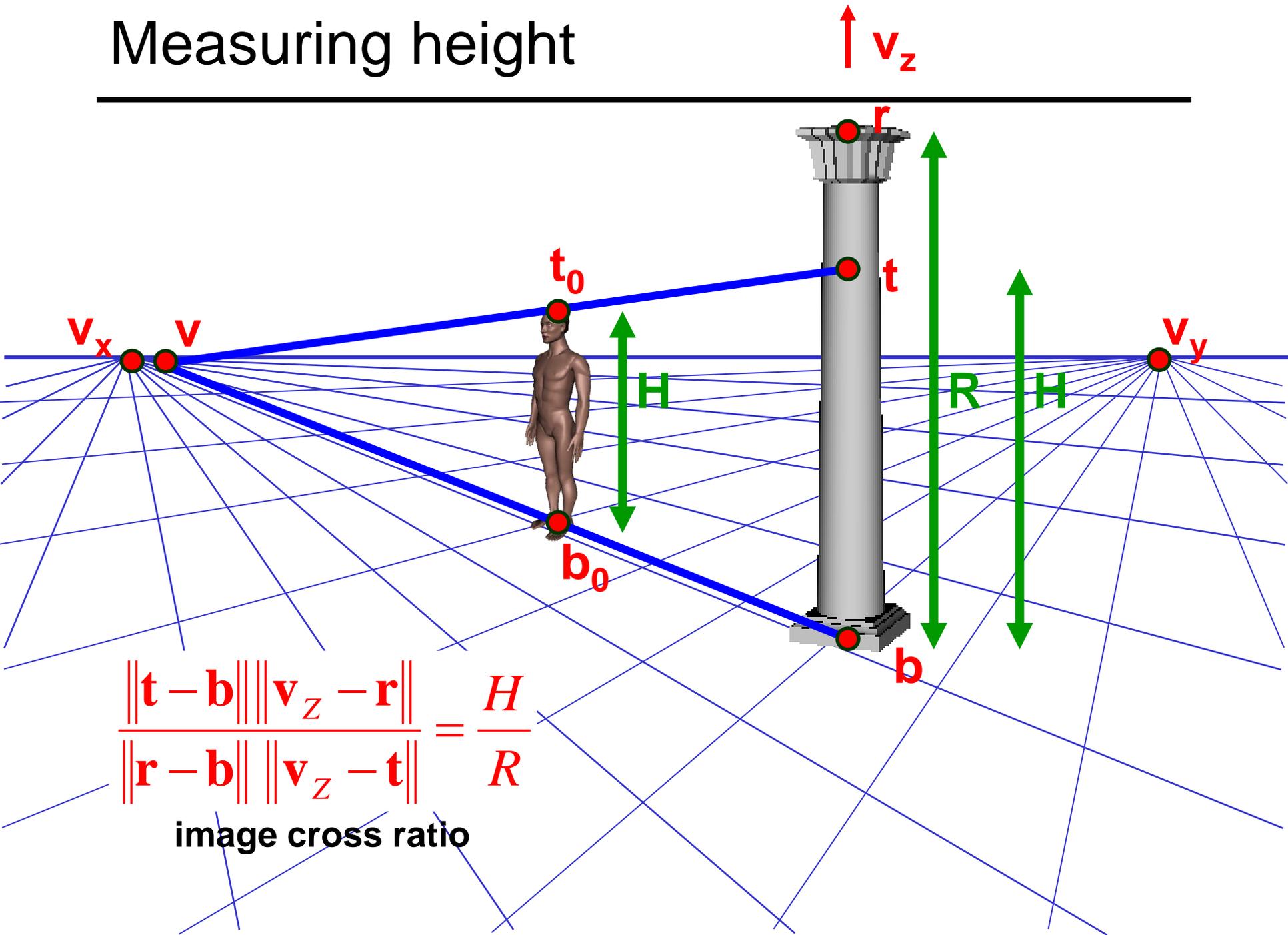


Announcements

- Final projects
 - Everyone should have received feedback by now
 - Midterm reports due November 24
 - Final presentations tentatively scheduled for the final exam period:

Wed, December 16, 7:00 PM - 9:30 PM

Measuring height



$$\frac{\|t - b\| \|v_z - r\|}{\|r - b\| \|v_z - t\|} = \frac{H}{R}$$

image cross ratio

Camera calibration

Goal: estimate the camera parameters

- Version 1: solve for projection matrix

$$\mathbf{x} = \begin{bmatrix} wx \\ wy \\ w \end{bmatrix} = \begin{bmatrix} * & * & * & * \\ * & * & * & * \\ * & * & * & * \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \\ 1 \end{bmatrix} = \mathbf{P}\mathbf{X}$$

- Version 2: solve for camera parameters separately
 - intrinsics (focal length, principle point, pixel size)
 - extrinsics (rotation angles, translation)
 - radial distortion

Vanishing points and projection matrix

$$\mathbf{\Pi} = \begin{bmatrix} * & * & * & * \\ * & * & * & * \\ * & * & * & * \end{bmatrix} = \begin{bmatrix} \boldsymbol{\pi}_1 & \boldsymbol{\pi}_2 & \boldsymbol{\pi}_3 & \boldsymbol{\pi}_4 \end{bmatrix}$$

- $\boldsymbol{\pi}_1 = \mathbf{\Pi} \begin{bmatrix} 1 & 0 & 0 & 0 \end{bmatrix}^T = \mathbf{v}_X$ (X vanishing point)
- similarly, $\boldsymbol{\pi}_2 = \mathbf{v}_Y$, $\boldsymbol{\pi}_3 = \mathbf{v}_Z$
- $\boldsymbol{\pi}_4 = \mathbf{\Pi} \begin{bmatrix} 0 & 0 & 0 & 1 \end{bmatrix}^T =$ projection of world origin

$$\mathbf{\Pi} = \begin{bmatrix} \mathbf{v}_X & \mathbf{v}_Y & \mathbf{v}_Z & \mathbf{0} \end{bmatrix}$$

Not So Fast! We only know \mathbf{v} 's up to a scale factor

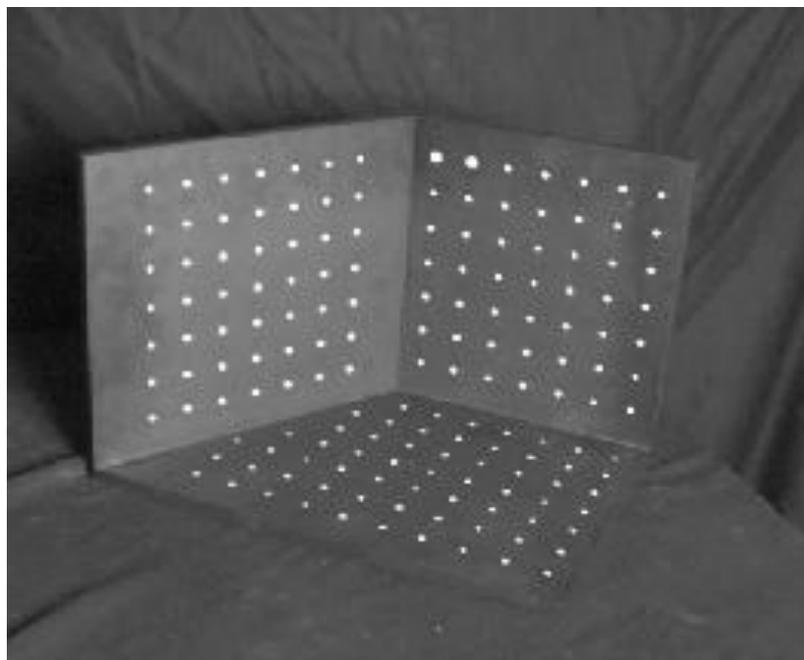
$$\mathbf{\Pi} = \begin{bmatrix} a \mathbf{v}_X & b \mathbf{v}_Y & c \mathbf{v}_Z & \mathbf{0} \end{bmatrix}$$

- Can fully specify by providing 3 reference points

Calibration using a reference object

Place a known object in the scene

- identify correspondence between image and scene
- compute mapping from scene to image



Issues

- must know geometry very accurately
- must know 3D->2D correspondence

Chromaglyphs



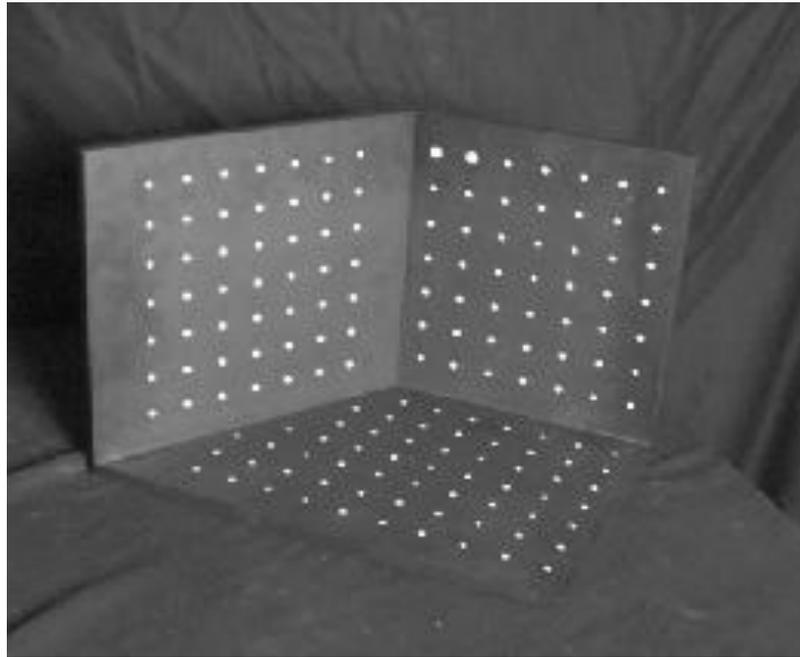
Courtesy of Bruce Culbertson, HP Labs

http://www.hpl.hp.com/personal/Bruce_Culbertson/ibr98/chromagl.htm

Estimating the projection matrix

Place a known object in the scene

- identify correspondence between image and scene
- compute mapping from scene to image



$$\begin{bmatrix} u_i \\ v_i \\ 1 \end{bmatrix} \stackrel{\mathbb{R}}{=} \begin{bmatrix} m_{00} & m_{01} & m_{02} & m_{03} \\ m_{10} & m_{11} & m_{12} & m_{13} \\ m_{20} & m_{21} & m_{22} & m_{23} \end{bmatrix} \begin{bmatrix} X_i \\ Y_i \\ Z_i \\ 1 \end{bmatrix}$$

Direct linear calibration

$$\begin{bmatrix} u_i \\ v_i \\ 1 \end{bmatrix} \cong \begin{bmatrix} m_{00} & m_{01} & m_{02} & m_{03} \\ m_{10} & m_{11} & m_{12} & m_{13} \\ m_{20} & m_{21} & m_{22} & m_{23} \end{bmatrix} \begin{bmatrix} X_i \\ Y_i \\ Z_i \\ 1 \end{bmatrix}$$

$$u_i = \frac{m_{00}X_i + m_{01}Y_i + m_{02}Z_i + m_{03}}{m_{20}X_i + m_{21}Y_i + m_{22}Z_i + m_{23}}$$

$$v_i = \frac{m_{10}X_i + m_{11}Y_i + m_{12}Z_i + m_{13}}{m_{20}X_i + m_{21}Y_i + m_{22}Z_i + m_{23}}$$

$$u_i(m_{20}X_i + m_{21}Y_i + m_{22}Z_i + m_{23}) = m_{00}X_i + m_{01}Y_i + m_{02}Z_i + m_{03}$$

$$v_i(m_{20}X_i + m_{21}Y_i + m_{22}Z_i + m_{23}) = m_{10}X_i + m_{11}Y_i + m_{12}Z_i + m_{13}$$

$$\begin{bmatrix} X_i & Y_i & Z_i & 1 & 0 & 0 & 0 & 0 & -u_iX_i & -u_iY_i & -u_iZ_i & -u_i \\ 0 & 0 & 0 & 0 & X_i & Y_i & Z_i & 1 & -v_iX_i & -v_iY_i & -v_iZ_i & -v_i \end{bmatrix} \begin{bmatrix} m_{00} \\ m_{01} \\ m_{02} \\ m_{03} \\ m_{10} \\ m_{11} \\ m_{12} \\ m_{13} \\ m_{20} \\ m_{21} \\ m_{22} \\ m_{23} \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \end{bmatrix}$$

Direct linear calibration

$$\begin{bmatrix} X_1 & Y_1 & Z_1 & 1 & 0 & 0 & 0 & 0 & -u_1 X_1 & -u_1 Y_1 & -u_1 Z_1 & -u_1 \\ 0 & 0 & 0 & 0 & X_1 & Y_1 & Z_1 & 1 & -v_1 X_1 & -v_1 Y_1 & -v_1 Z_1 & -v_1 \\ & & & & & & & \vdots & & & & \\ X_n & Y_n & Z_n & 1 & 0 & 0 & 0 & 0 & -u_n X_n & -u_n Y_n & -u_n Z_n & -u_n \\ 0 & 0 & 0 & 0 & X_n & Y_n & Z_n & 1 & -v_n X_n & -v_n Y_n & -v_n Z_n & -v_n \end{bmatrix} \begin{bmatrix} m_{00} \\ m_{01} \\ m_{02} \\ m_{03} \\ m_{10} \\ m_{11} \\ m_{12} \\ m_{13} \\ m_{20} \\ m_{21} \\ m_{22} \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ \vdots \\ 0 \\ 0 \end{bmatrix}$$

Can solve for m_{ij} by linear least squares

- use eigenvector trick that we used for homographies

Direct linear calibration

Advantage:

- Very simple to formulate and solve

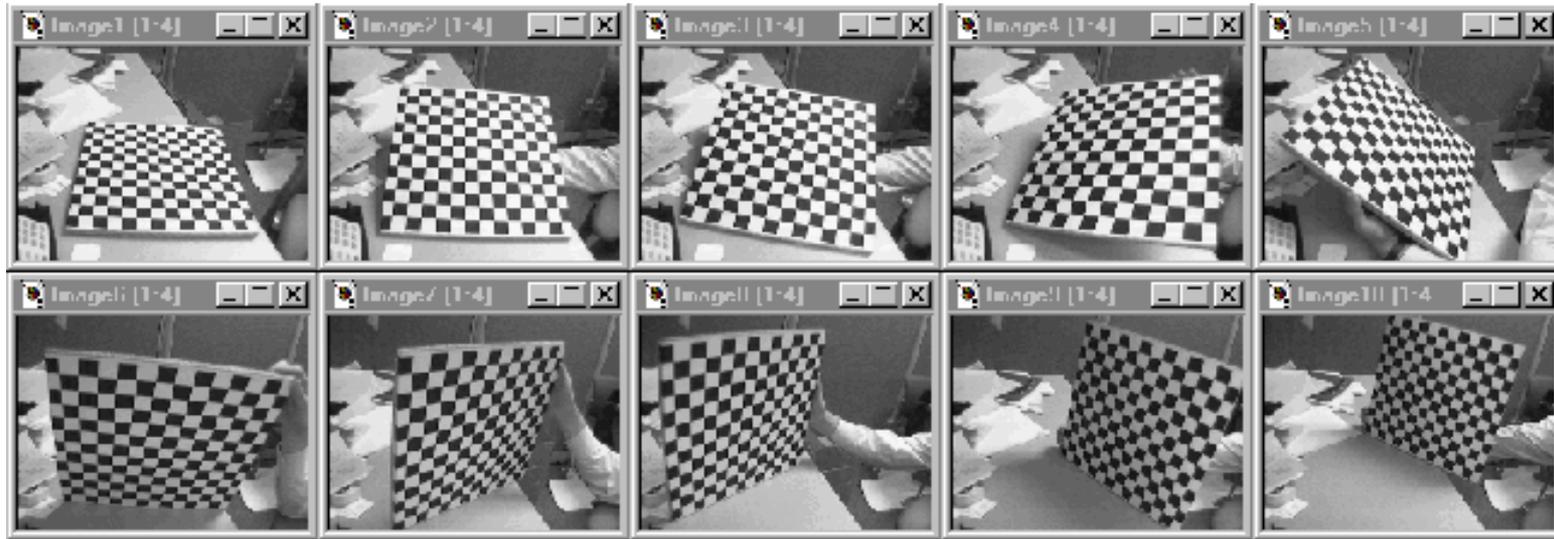
Disadvantages:

- Doesn't tell you the camera parameters
- Doesn't model radial distortion
- Hard to impose constraints (e.g., known focal length)
- Doesn't minimize the right error function

For these reasons, *nonlinear methods* are preferred

- Define error function E between projected 3D points and image positions
 - E is nonlinear function of intrinsics, extrinsics, radial distortion
- Minimize E using nonlinear optimization techniques
 - e.g., variants of Newton's method (e.g., Levenberg Marquart)

Alternative: multi-plane calibration

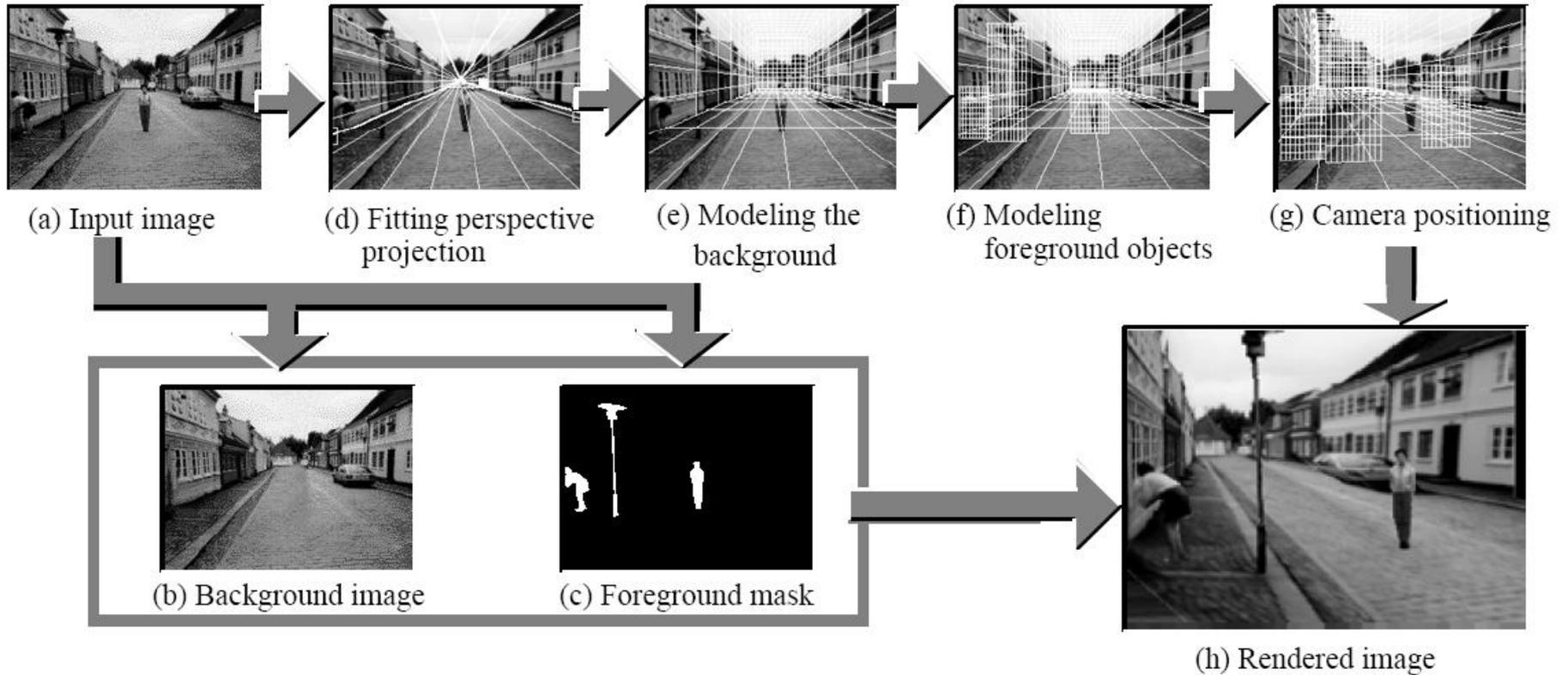


Images courtesy Jean-Yves Bouguet, Intel Corp.

Advantage

- Only requires a plane
- Don't have to know positions/orientations
- Good code available online!
 - Intel's OpenCV library: <http://www.intel.com/research/mrl/research/opencv/>
 - Matlab version by Jean-Yves Bouguet: http://www.vision.caltech.edu/bouguetj/calib_doc/index.html
 - Zhengyou Zhang's web site: <http://research.microsoft.com/~zhang/Calib/>

Some Related Techniques



Tour Into The Picture

Anjyo et al., SIGGRAPH 1997

http://koigakubo.hitachi.co.jp/little/DL_TipE.html

Some Related Techniques

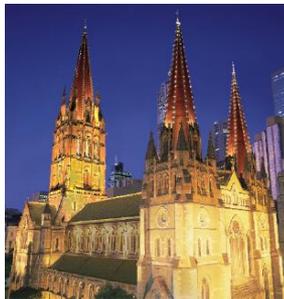


Image-Based Modeling and Photo Editing

- Mok et al., SIGGRAPH 2001
- <http://graphics.csail.mit.edu/ibedit/>



Single View Modeling of Free-Form Scenes

- Zhang et al., CVPR 2001
- <http://grail.cs.washington.edu/projects/svm/>

Automatic approaches

Make3D
Convert your image into 3d model

[Gallery](#)
[My 3D Images](#)
[Upload](#)

[Research/Code](#)
[Publications](#)
[Contact](#)
[FAQs](#)
[Guest Book](#)

Make3D --- convert your image into 3d model
☆☆☆☆☆

[Visit 3-D Gallery](#)
(Recent models.)

Our study has finished!
Thanks for uploading the pictures.
See the pictures uploaded by tens of thousands of users.
(Only public pictures are viewable).
Also, [download code](#).

139 diggs [digg it](#)

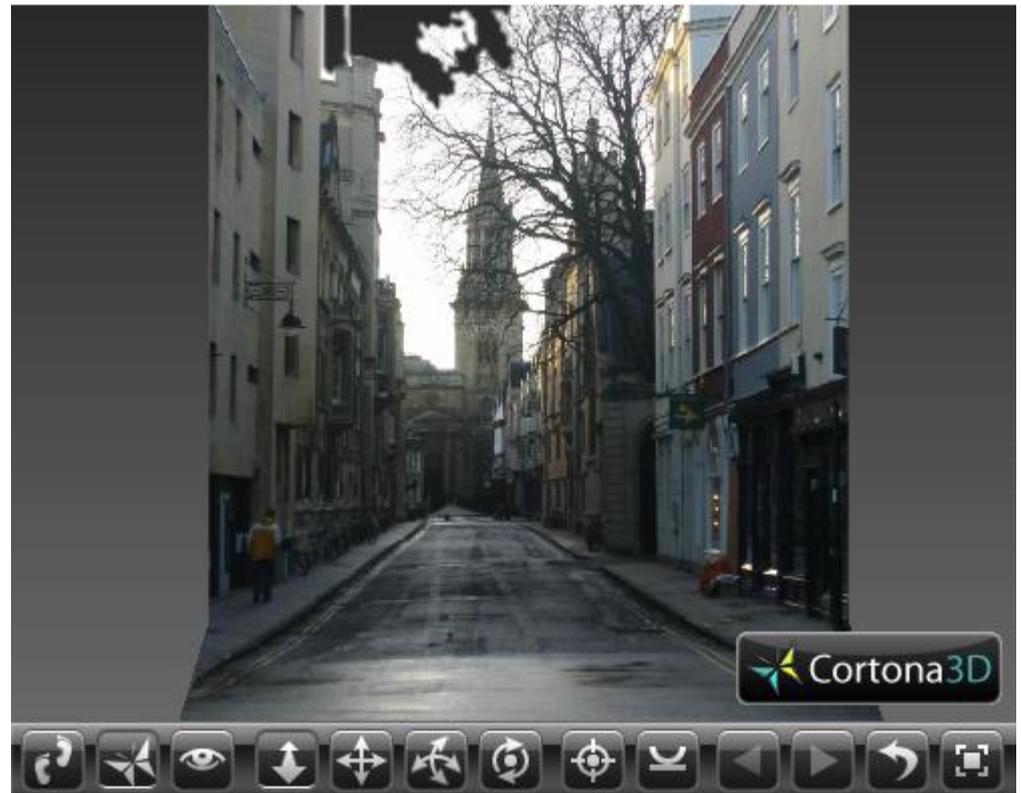
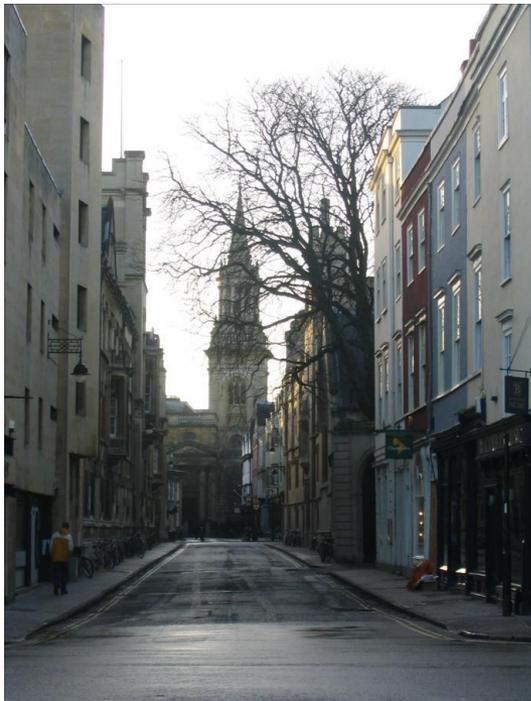
Make3D converts your single picture into a 3-D model, completely automatically.

It takes a two-dimensional image and creates a three-dimensional "fly around" model, giving the viewers access to the scene's depth and a range of points of view. After uploading your image, you can "fly" in the 3-D scene (requires VRML viewer or Adobe Shockwave), or watch a rendered 3-d movie (flash required). Visit [3-D Gallery](#), [Hall of fame](#), or [signup](#) to upload your photograph.

How does it work?
It uses powerful machine learning techniques to learn the 3-d structure of a scene as a function of the (single) image features. The team led by [Ashutosh Saxena](#) and [Andrew Y. Ng](#) has made the code available online; see more details [here](#).

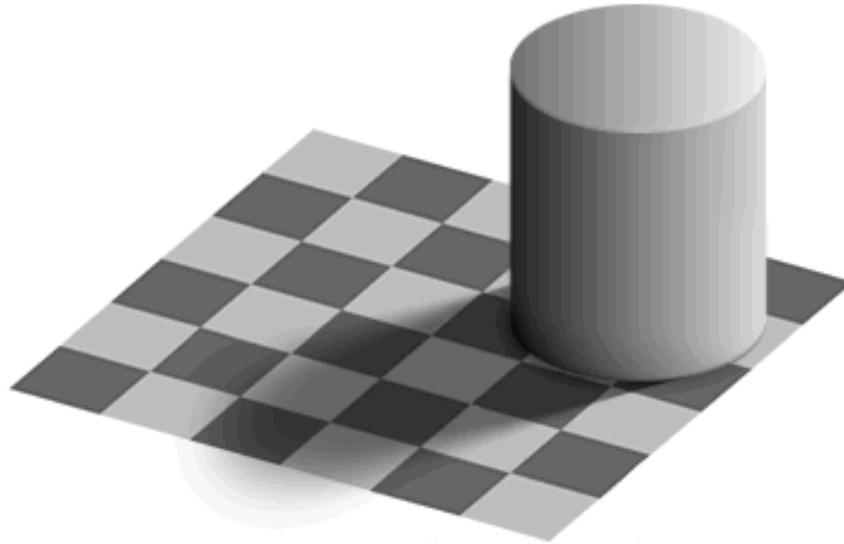
Make3D, Ashutosh Saxena

Automatic approaches



D. Hoiem, A.A. Efros, and M. Hebert, "Automatic Photo Pop-up", ACM SIGGRAPH 2005.

Light

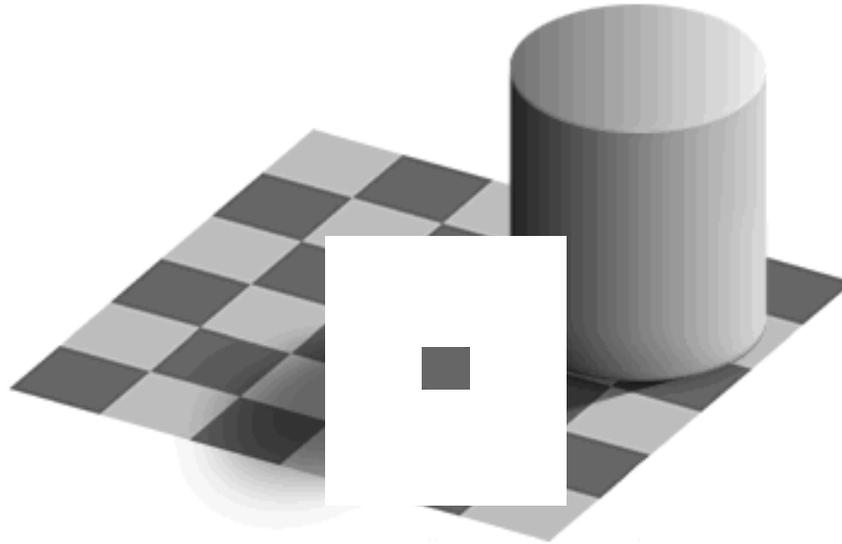


by Ted Adelson

Readings

- Szeliski, 2.2, 2.3.2

Light



by Ted Adelson

Readings

- Szeliski, 2.2, 2.3.2

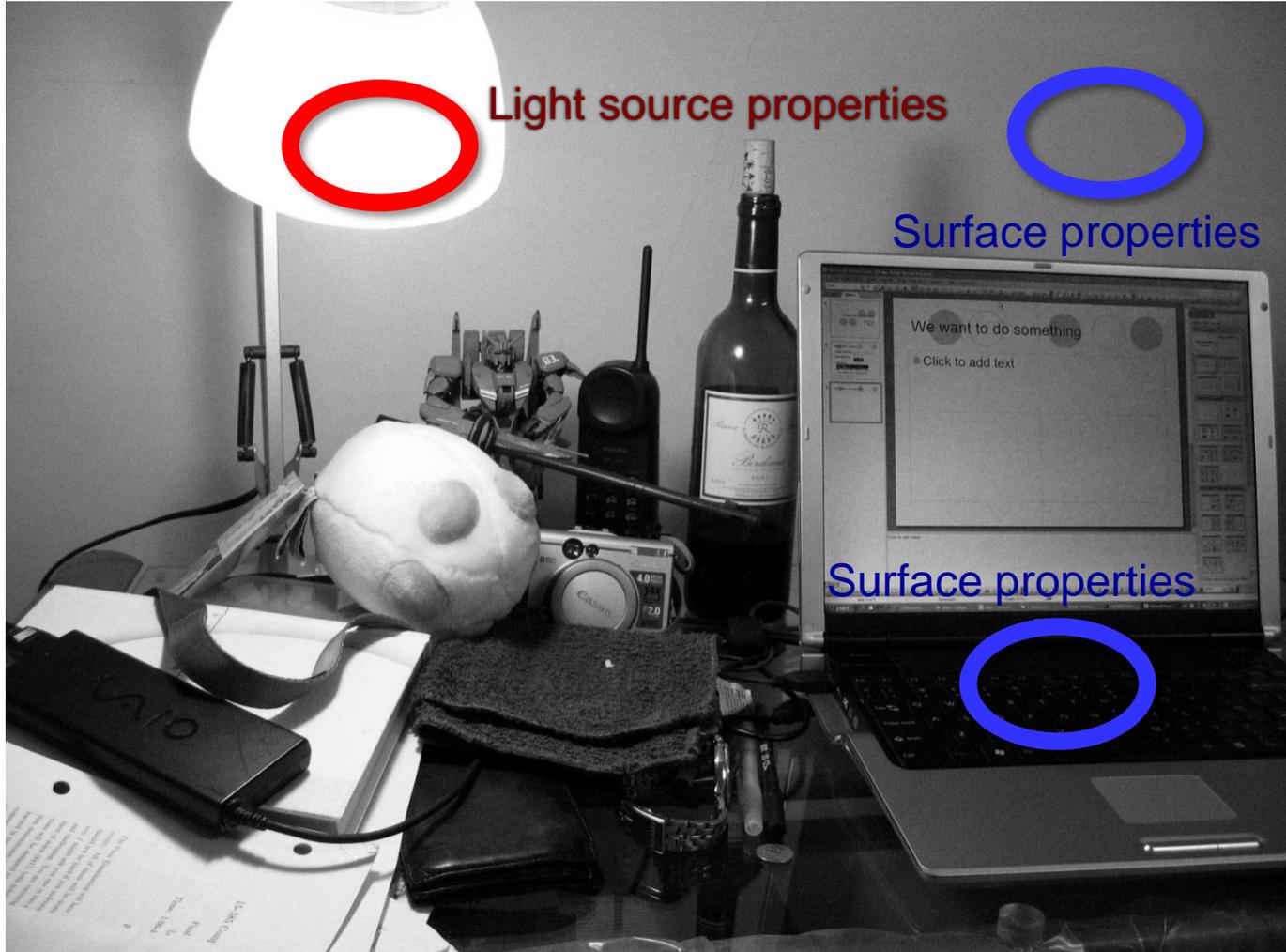
Properties of light

Today

- What is light?
- How do we measure it?
- How does light propagate?
- How does light interact with matter?

Radiometry

What determines the brightness of an image pixel?



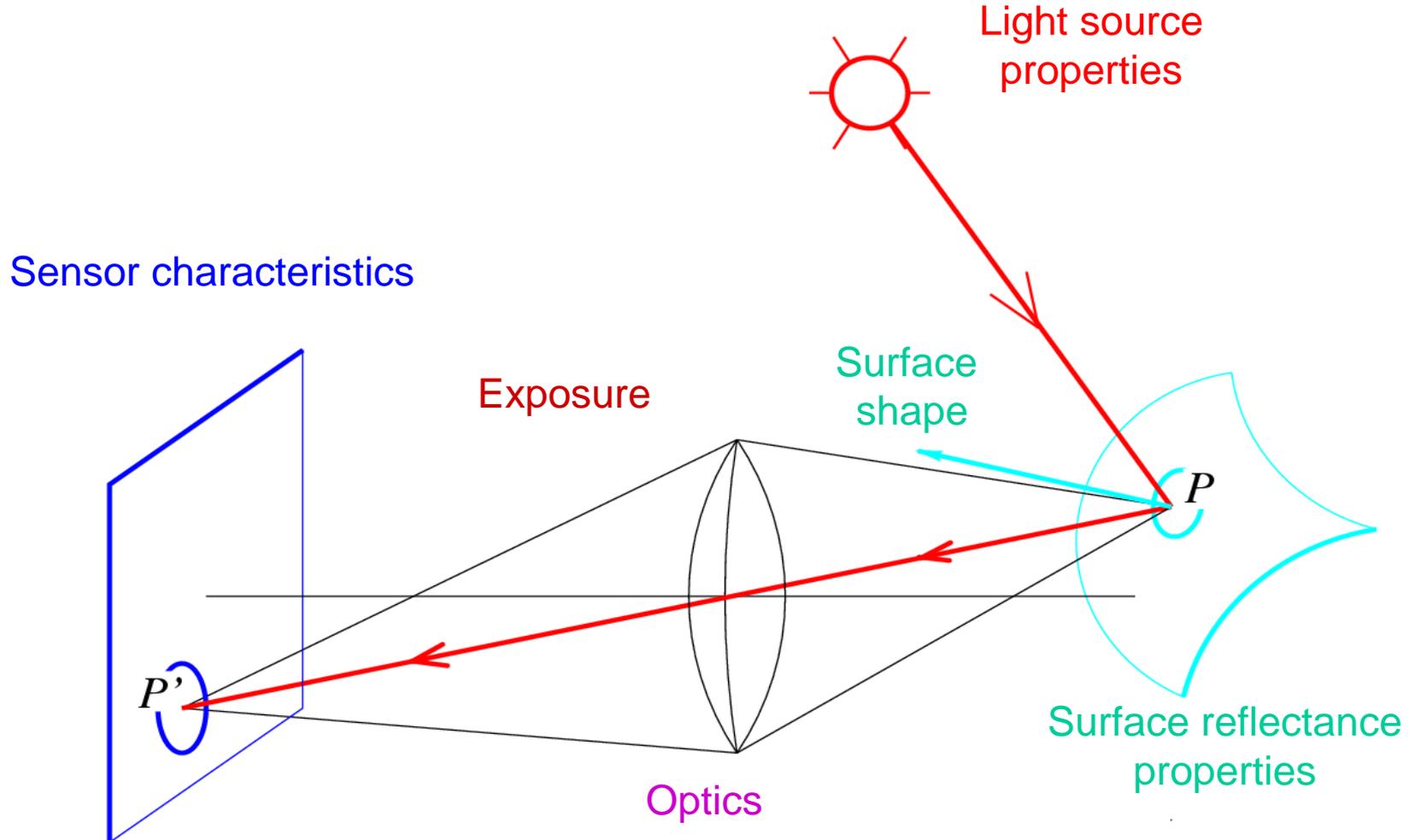
Radiometry

What determines the brightness of an image pixel?



Radiometry

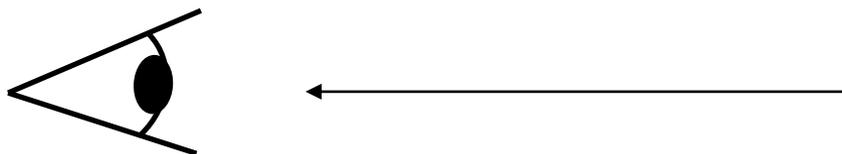
What determines the brightness of an image pixel?



What is light?

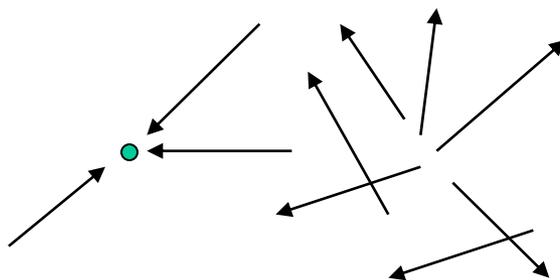
Electromagnetic radiation (EMR) moving along rays in space

- $R(\lambda)$ is EMR, measured in units of power (watts)
 - λ is wavelength



Light field

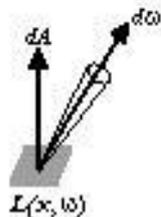
- We can describe all of the light in the scene by specifying the radiation (or “**radiance**” along all light rays) arriving at every point in space and from every direction



$$R(X, Y, Z, \theta, \phi, \lambda, t)$$

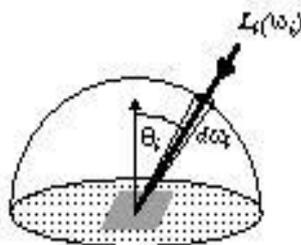
Radiometry

Radiometry is the science of light energy measurement



Radiance

**The energy carried by a ray
energy/(area solidangle)**



Irradiance

**The energy per unit area
falling on a surface**

Radiosity

**The energy per unit area
leaving a surface**

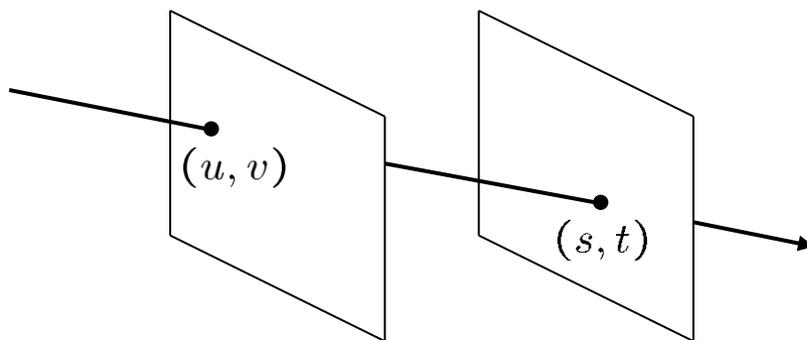
The light field

$$R(X, Y, Z, \theta, \phi, \lambda, t)$$

- Known as the **plenoptic function**
- If you know R , you can predict how the scene would appear from any viewpoint.

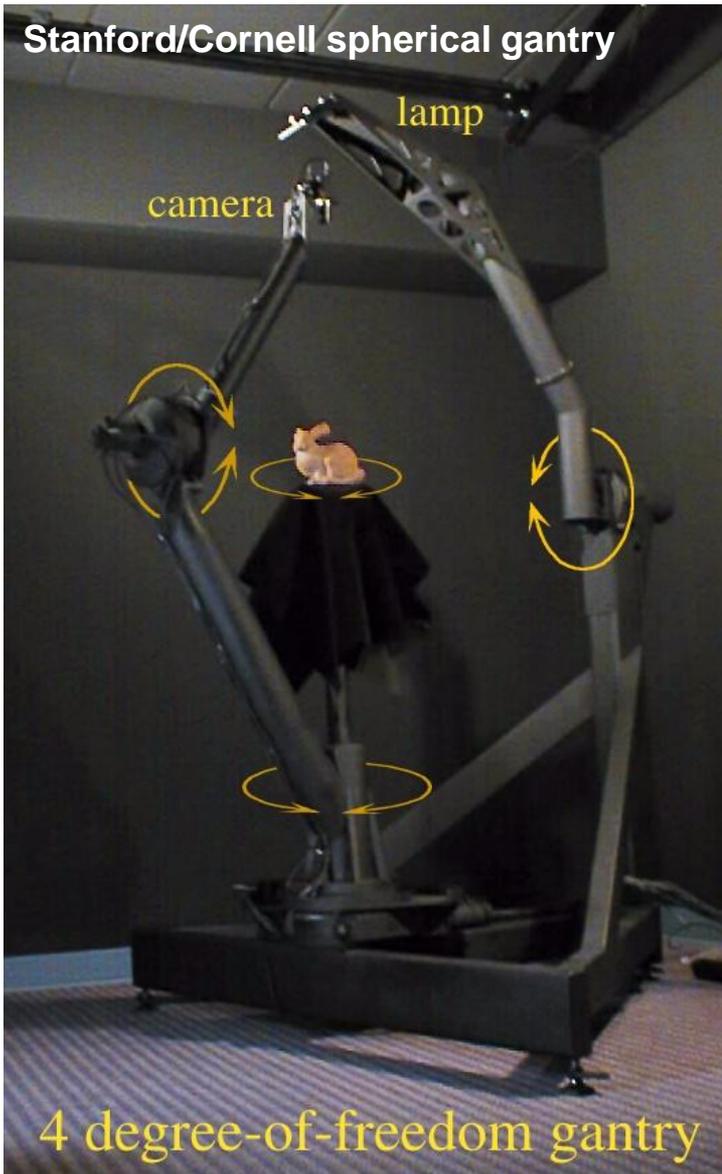
The **light field** $R(u, v, s, t)$ — t is *not* time (different from above t !)

- Assume radiance does not change along a ray
 - what does this assume about the world?
- Parameterize rays by intersection with two planes:

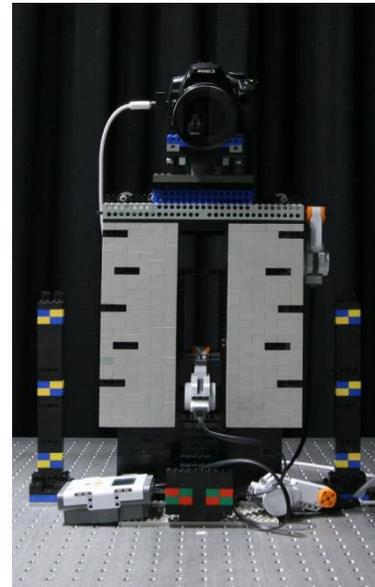


- Usually drop λ and time parameters
- How could you capture a light field?

Capturing light fields



Stanford Multi-Camera Array



Lego Mindstorms Gantry



Handheld light field camera

Light field example



More info on light fields

If you're interested to read more:

The plenoptic function

- **Original reference:** E. Adelson and J. Bergen, "[The Plenoptic Function and the Elements of Early Vision](#)," in M. Landy and J. A. Movshon, (eds) Computational Models of Visual Processing, MIT Press 1991.
- L. McMillan and G. Bishop, "[Plenoptic Modeling: An Image-Based Rendering System](#)", Proc. SIGGRAPH, 1995, pp. 39-46.

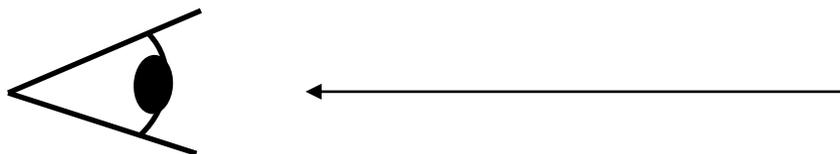
The light field

- M. Levoy and P. Hanrahan, "[Light Field Rendering](#)", Proc SIGGRAPH 96, pp. 31-42.
- S. J. Gortler, R. Grzeszczuk, R. Szeliski, and M. F. Cohen, "[The lumigraph](#)," in Proc. SIGGRAPH, 1996, pp. 43-54.

What is light?

Electromagnetic radiation (EMR) moving along rays in space

- $R(\lambda)$ is EMR, measured in units of power (watts)
 - λ is wavelength



Perceiving light

- How do we convert radiation into “color”?
- What part of the spectrum do we see?

Visible light

We “see” electromagnetic radiation in a range of wavelengths

