

# CS5670 : Computer Vision

Noah Snavely

## Photometric stereo

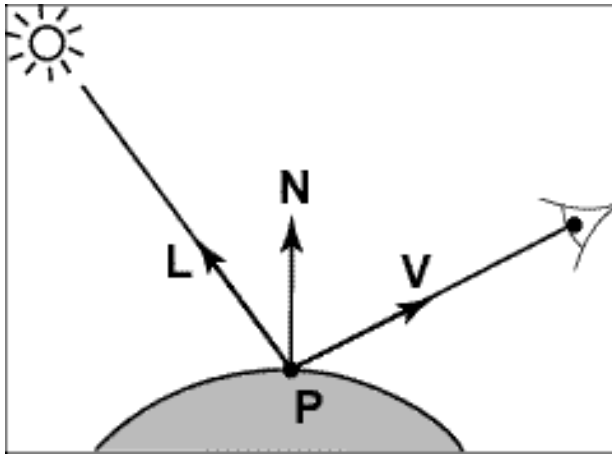


# Announcements

- Project 3 released
  - Due Monday, April 22, by 11:59pm
  - To be done in groups of 2
  - Demo by Kai

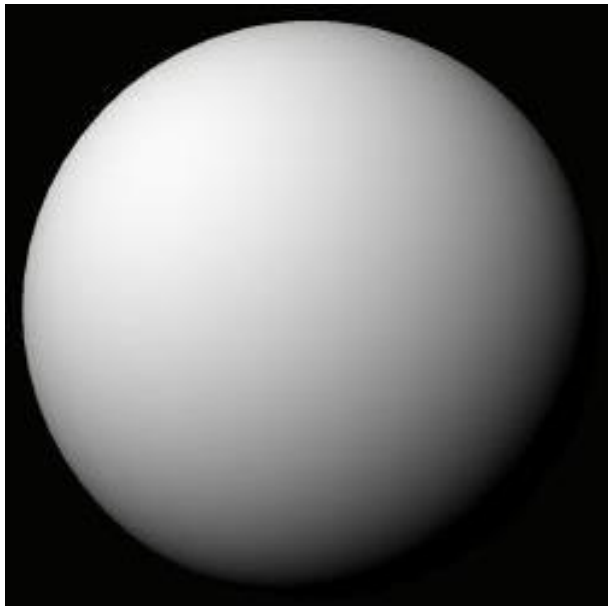
# Recap:

## Lambertian (Diffuse) Reflectance



$$I = k_d \mathbf{N} \cdot \mathbf{L}$$

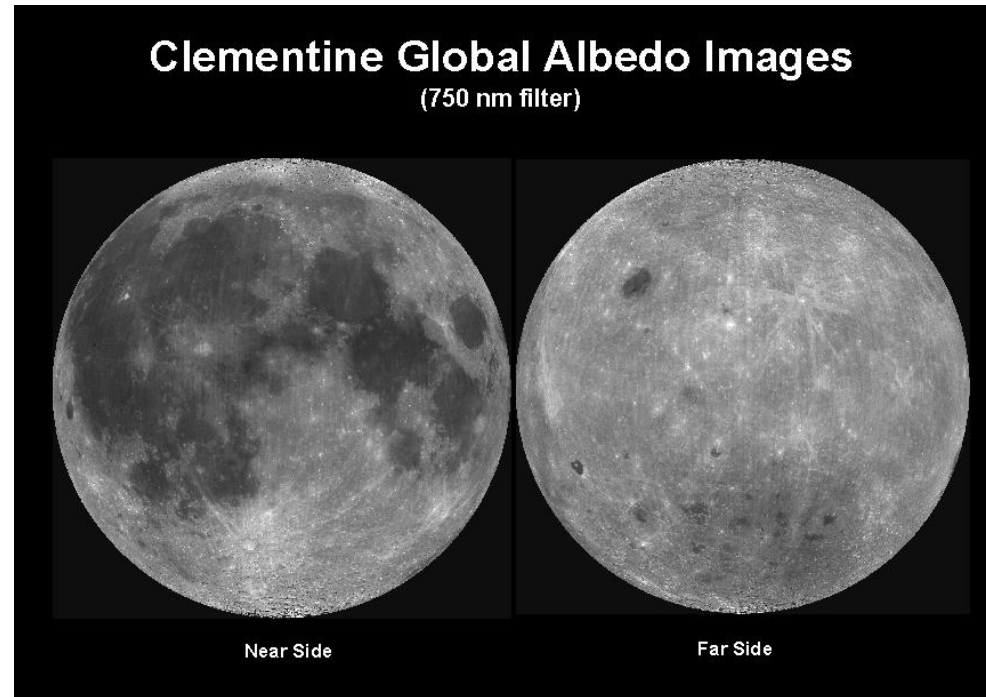
- $I$  : observed image intensity
- $k_d$  : object albedo
- $\mathbf{N}$  : surface normal
- $\mathbf{L}$  : light source direction



Lambertian sphere with constant albedo lit by a directional light source

## Sample albedos

Surface	Typical albedo
Fresh asphalt	0.04 <sup>[4]</sup>
Open ocean	0.06 <sup>[5]</sup>
Worn asphalt	0.12 <sup>[4]</sup>
Conifer forest (Summer)	0.08, <sup>[6]</sup> 0.09 to 0.15 <sup>[7]</sup>
Deciduous trees	0.15 to 0.18 <sup>[7]</sup>
Bare soil	0.17 <sup>[8]</sup>
Green grass	0.25 <sup>[8]</sup>
Desert sand	0.40 <sup>[9]</sup>
New concrete	0.55 <sup>[8]</sup>
Ocean ice	0.5–0.7 <sup>[8]</sup>
Fresh snow	0.80–0.90 <sup>[8]</sup>

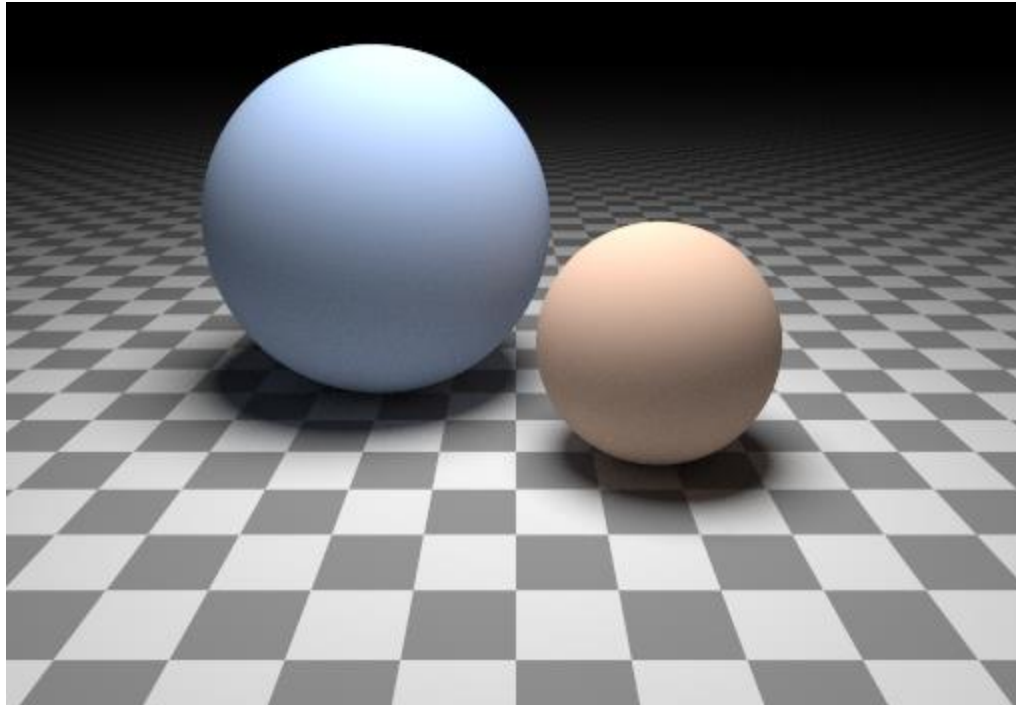


Objects can have varying albedo and albedo varies with wavelength

Source:

<https://en.wikipedia.org/wiki/Albedo>

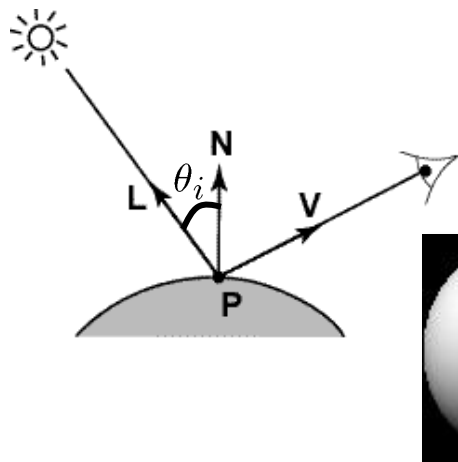
# Can we determine shape from lighting?



- Are these spheres?
  - Or just flat discs painted with varying albedo?

# Shape from shading

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Suppose  $k_d = 1$

$$\begin{aligned} I &= k_d \mathbf{N} \cdot \mathbf{L} \\ &= \mathbf{N} \cdot \mathbf{L} \\ &= \cos \theta_i \end{aligned}$$

You can directly measure angle between normal and light source

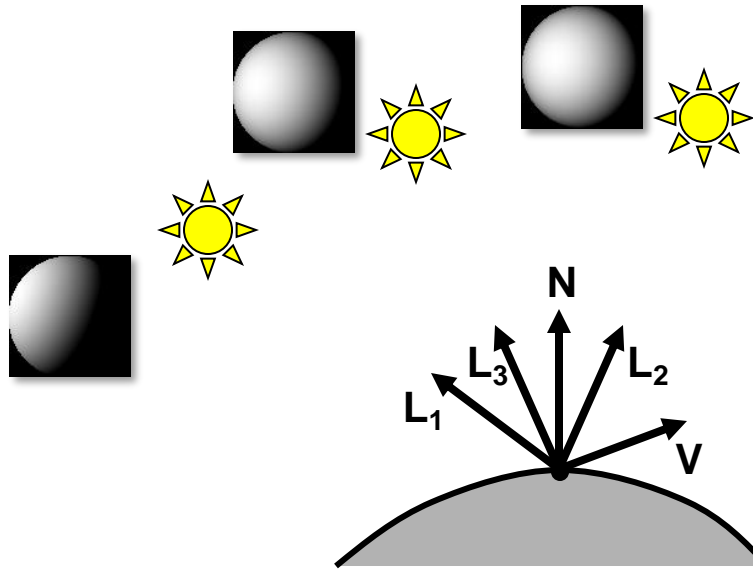
- Not quite enough information to compute surface shape
- But can be if you add some additional info, for example
  - assume a few of the normals are known (e.g., along silhouette)
  - constraints on neighboring normals—“integrability”
  - smoothness
- Hard to get it to work well in practice
  - plus, how many real objects have constant albedo?
  - But, deep learning can help

# Let's take more than one photo!

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# Photometric stereo

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$$I_1 = k_d \mathbf{N} \cdot \mathbf{L}_1$$

$$I_2 = k_d \mathbf{N} \cdot \mathbf{L}_2$$

$$I_3 = k_d \mathbf{N} \cdot \mathbf{L}_3$$

Can write this as a matrix equation:

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = k_d \begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix} \mathbf{N}$$



# Solving the equations

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$$\underbrace{\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix}}_{\substack{\mathbf{I} \\ 3 \times 1}} = \underbrace{\begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix}}_{\substack{\mathbf{L} \\ 3 \times 3}} \underbrace{k_d \mathbf{N}}_{\substack{\mathbf{G} \\ 3 \times 1}}$$

$$\mathbf{G} = \mathbf{L}^{-1} \mathbf{I}$$

$$k_d = \|\mathbf{G}\|$$

$$\mathbf{N} = \frac{1}{k_d} \mathbf{G}$$

Solve one such linear system **per pixel** to solve for that pixel's surface normal

# More than three lights

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Get better results by using more lights

$$\begin{bmatrix} I_1 \\ \vdots \\ I_n \end{bmatrix} = \begin{bmatrix} \mathbf{L}_1 \\ \vdots \\ \mathbf{L}_n \end{bmatrix} k_d \mathbf{N}$$

Least squares solution:

$$\begin{aligned} \mathbf{I} &= \mathbf{L}\mathbf{G} \\ \mathbf{L}^T \mathbf{I} &= \mathbf{L}^T \mathbf{L} \mathbf{G} \\ \mathbf{G} &= (\mathbf{L}^T \mathbf{L})^{-1} (\mathbf{L}^T \mathbf{I}) \end{aligned}$$

Solve for  $\mathbf{N}$ ,  $k_d$  as before

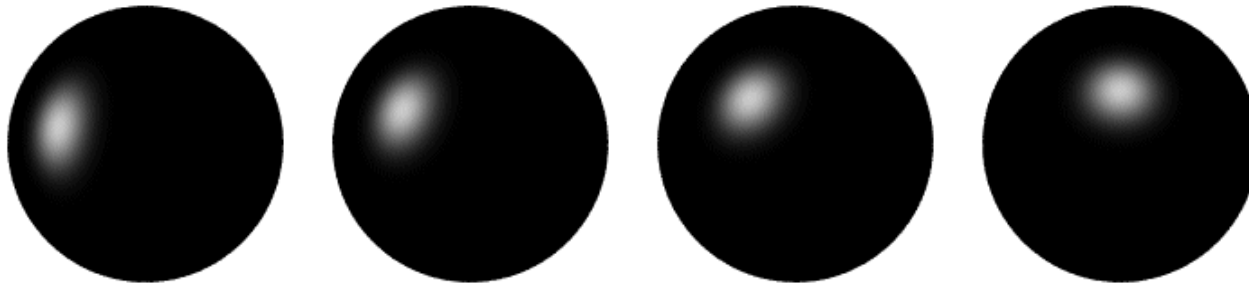
What's the size of  $\mathbf{L}^T \mathbf{L}$ ?



# Computing light source directions

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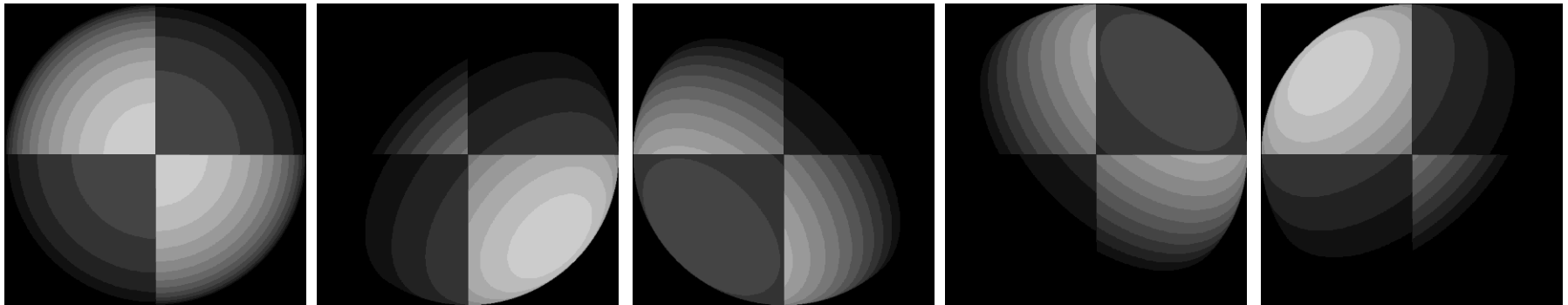
Trick: place a chrome sphere in the scene



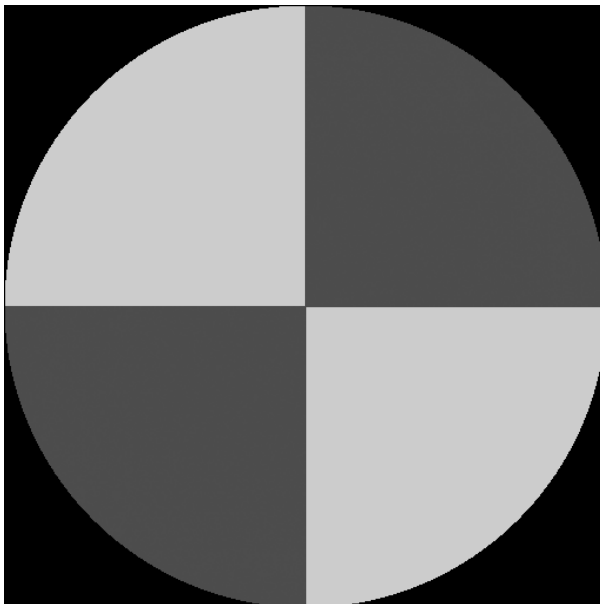
- the location of the highlight tells you where the light source is

# Example

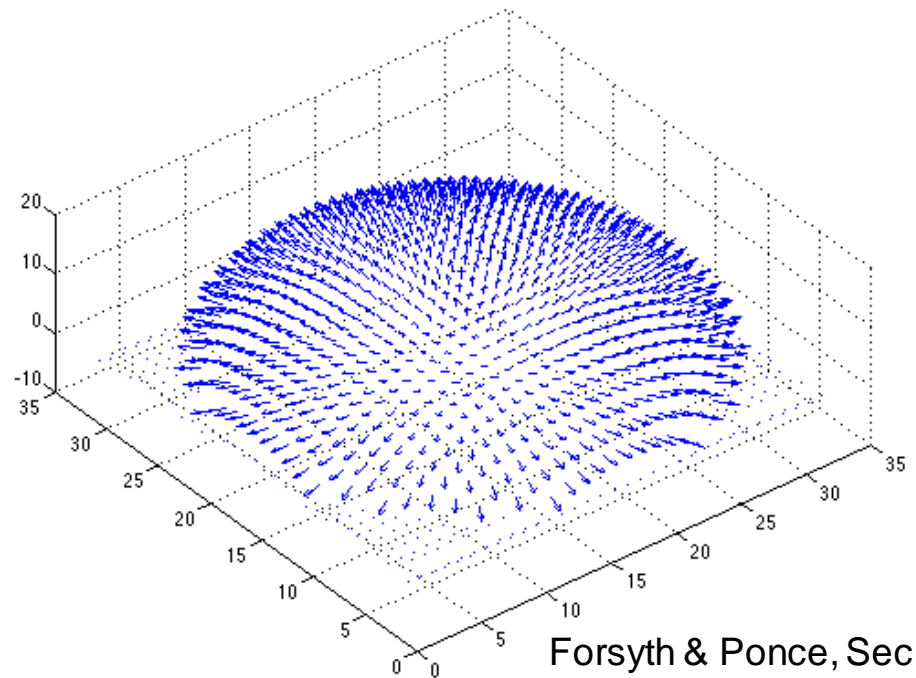
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Recovered albedo



Recovered normal field



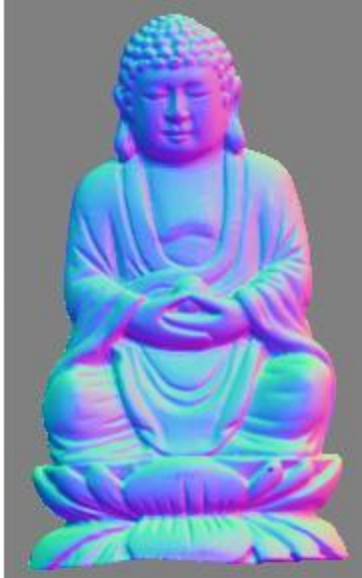
# Depth from normals

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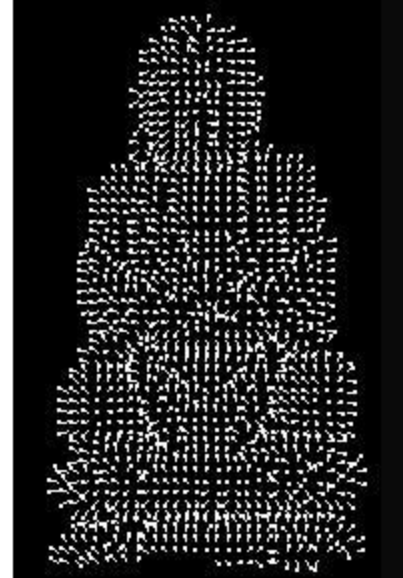
- Solving the linear system per-pixel gives us an estimated surface normal for each pixel



Input photo



Estimated normals



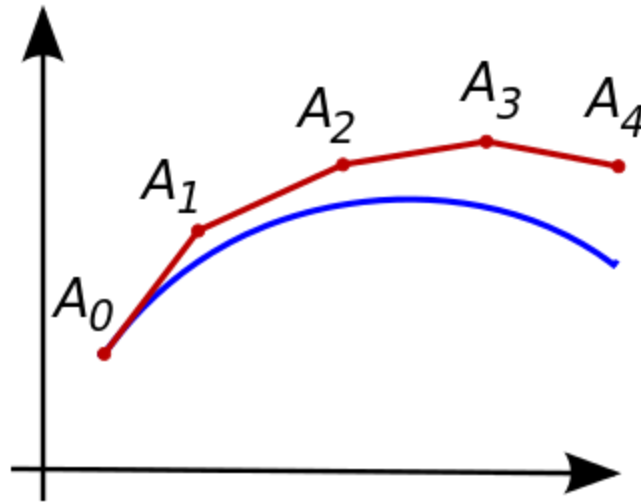
Estimated normals  
(needle diagram)

- How can we compute depth from normals?
  - Normals are like the “derivative” of the true depth

# Normal Integration

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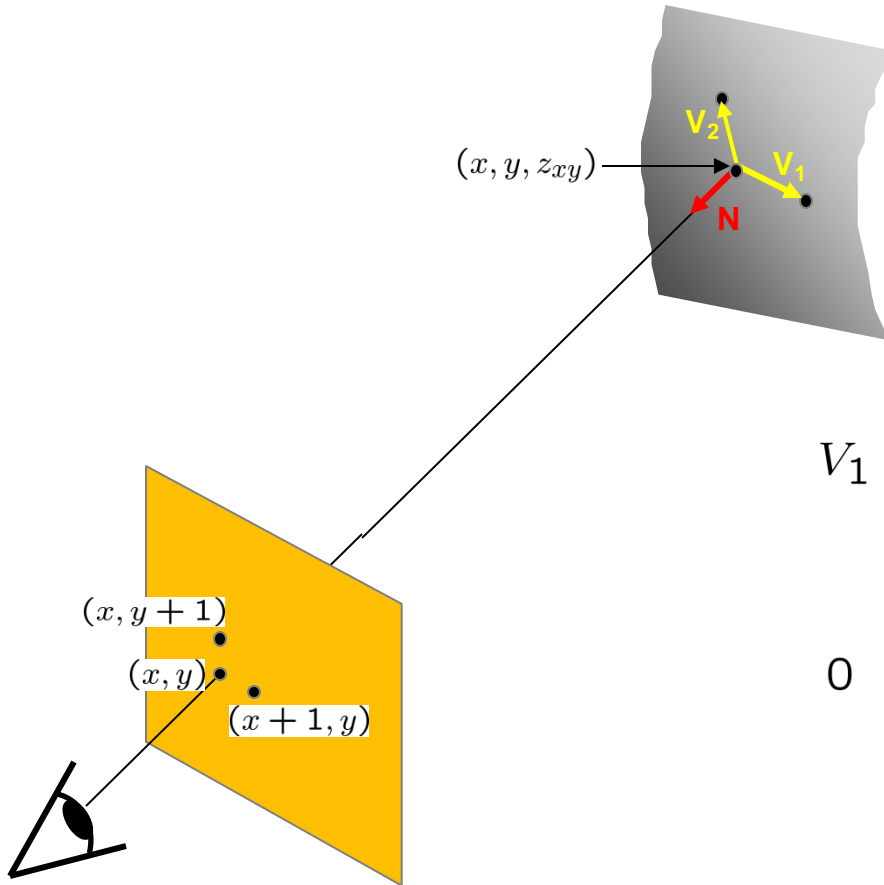
- Integrating a set of derivatives is easy in 1D
  - (similar to Euler's method from diff. eq. class)



- Could just integrate normals in each column / row separately
- Instead, we formulate as a linear system and solve for depths that *best agree with the surface normals*

# Depth from normals

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$$\begin{aligned} V_1 &= (x+1, y, z_{x+1,y}) - (x, y, z_{xy}) \\ &= (1, 0, z_{x+1,y} - z_{xy}) \end{aligned}$$

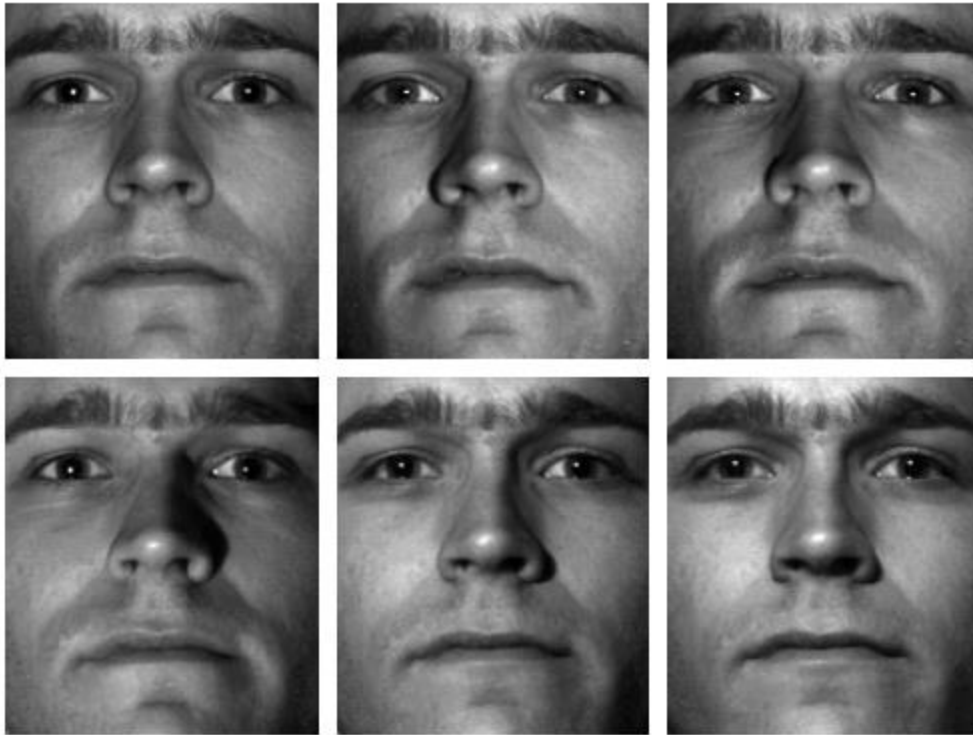
$$\begin{aligned} 0 &= N \cdot V_1 \\ &= (n_x, n_y, n_z) \cdot (1, 0, z_{x+1,y} - z_{xy}) \\ &= n_x + n_z(z_{x+1,y} - z_{xy}) \end{aligned}$$

Get a similar equation for  $V_2$

- Each normal gives us two linear constraints on  $z$
- compute  $z$  values by solving a matrix equation

# Results

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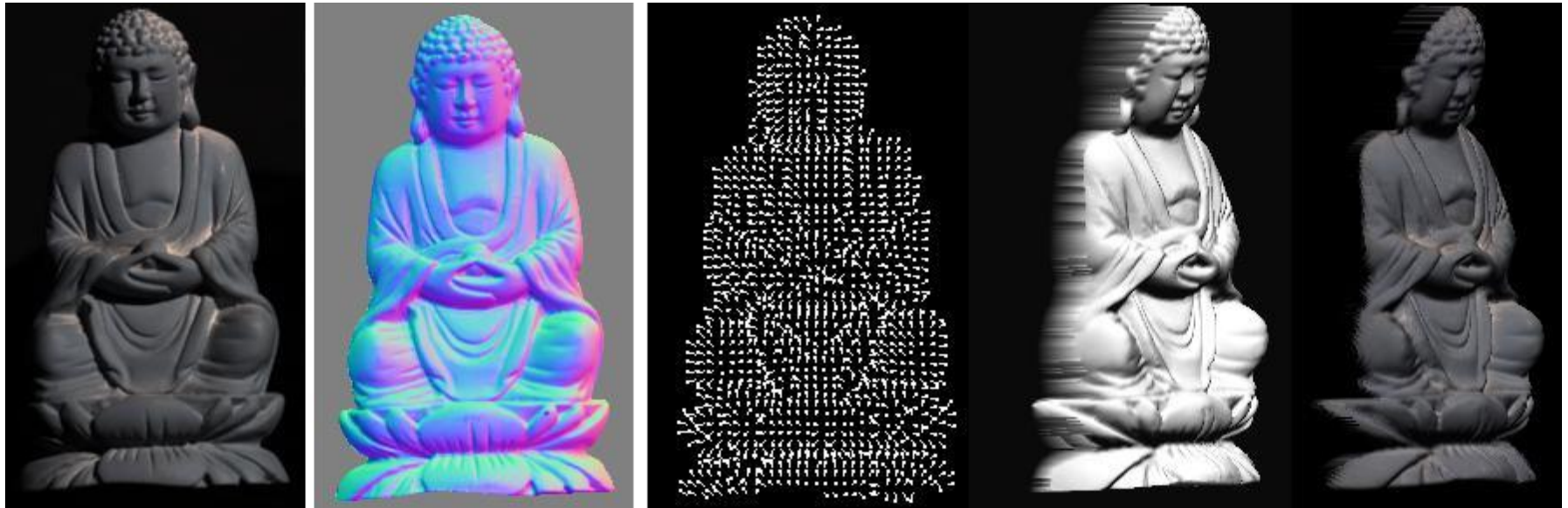


from Athos Georghiades



# Results

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# Extension

## Photometric Stereo from Colored Lighting

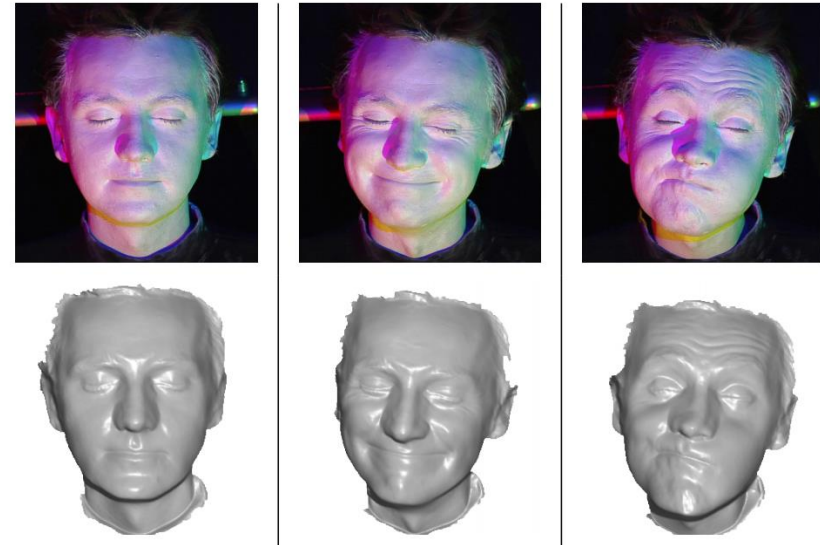
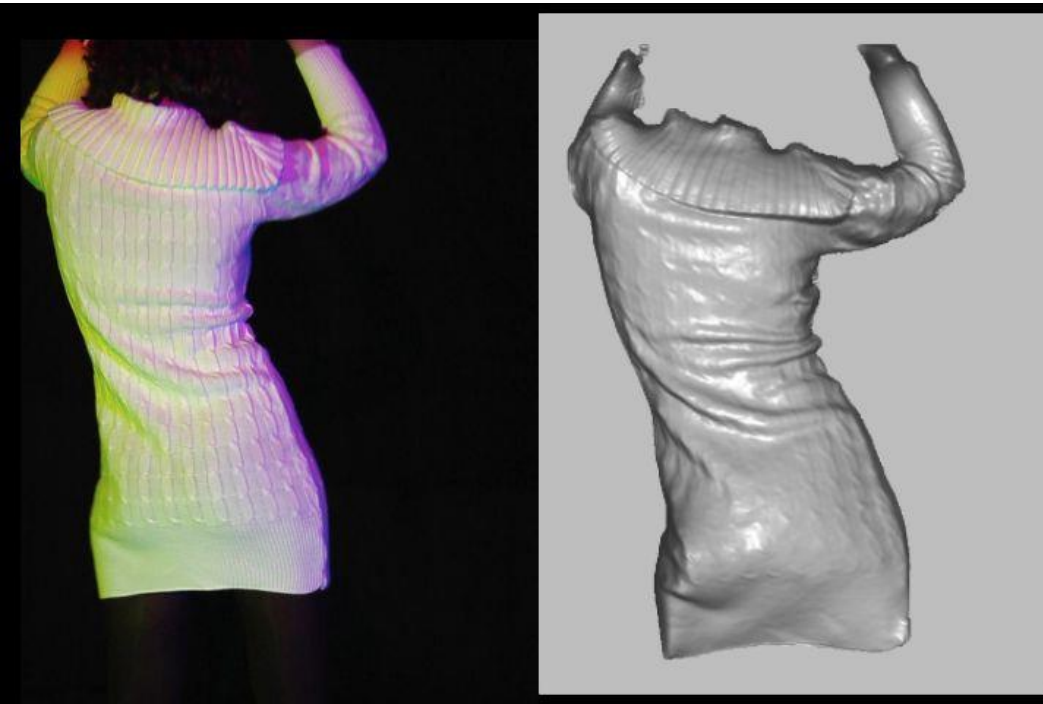


Fig. 2. **Applying the original algorithm to a face with white makeup.**  
Top: example input frames from video of an actor smiling and grimacing.  
Bottom: the resulting integrated surfaces.

### Video Normals from Colored Lights

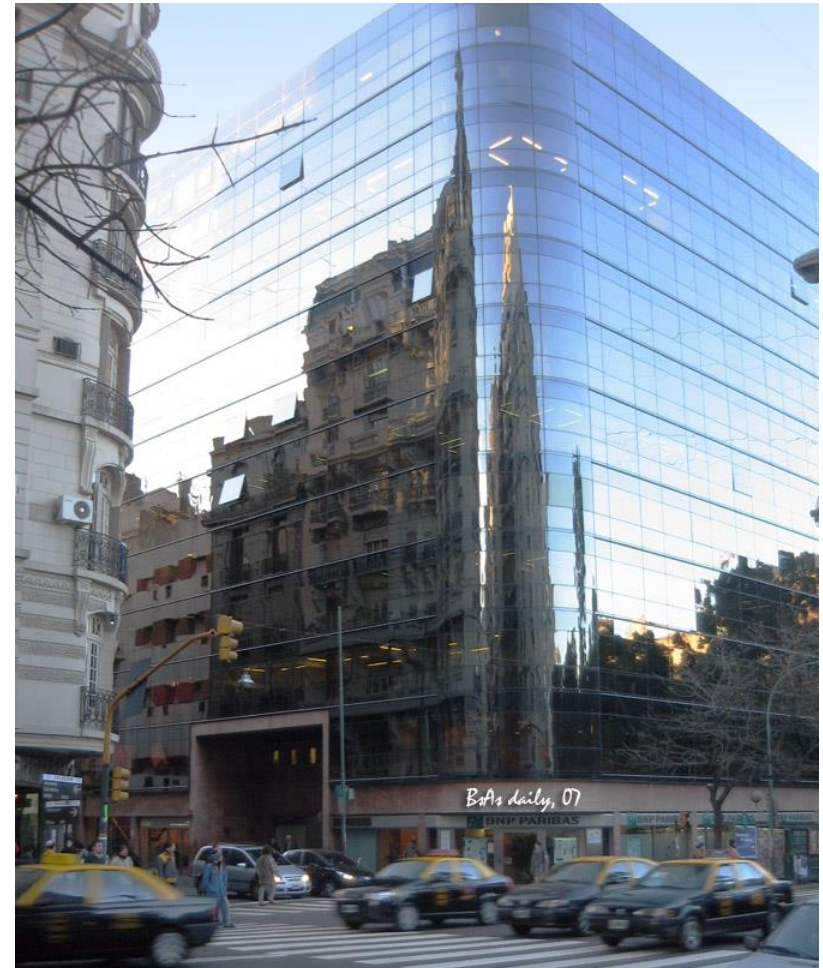
Gabriel J. Brostow, Carlos Hernández, George Vogiatzis, Björn Stenger, Roberto Cipolla  
[IEEE TPAMI](#), Vol. 33, No. 10, pages 2104-2114, October 2011.

# Questions?

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# For now, ignore specular reflection

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# And Refraction...

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# And Interreflections...

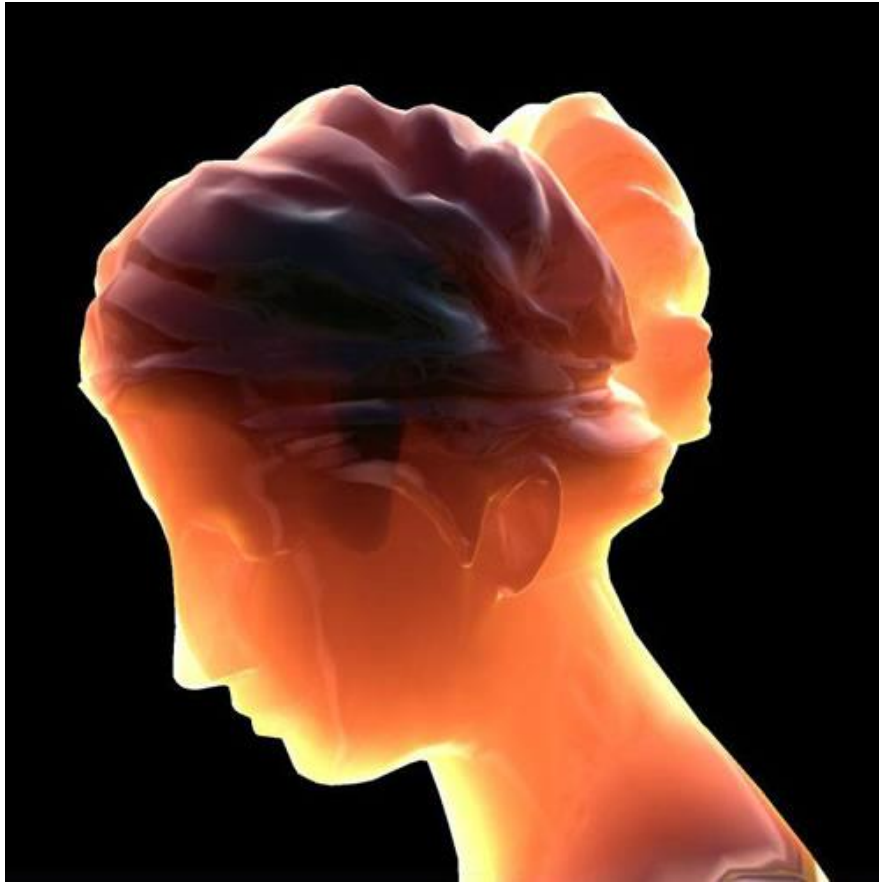
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Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

# And Subsurface Scattering...

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# Limitations

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## Bigger problems

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections

## Smaller problems

- camera and lights have to be distant
- calibration requirements
  - measure light source directions, intensities
  - camera response function

Newer work addresses some of these issues

## Some pointers for further reading:

- Zickler, Belhumeur, and Kriegman, "[Helmholtz Stereopsis: Exploiting Reciprocity for Surface Reconstruction](#)." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann & Seitz, "[Example-Based Photometric Stereo: Shape Reconstruction with General, Varying BRDFs](#)." IEEE Trans. PAMI 2005



# GELSIGHT

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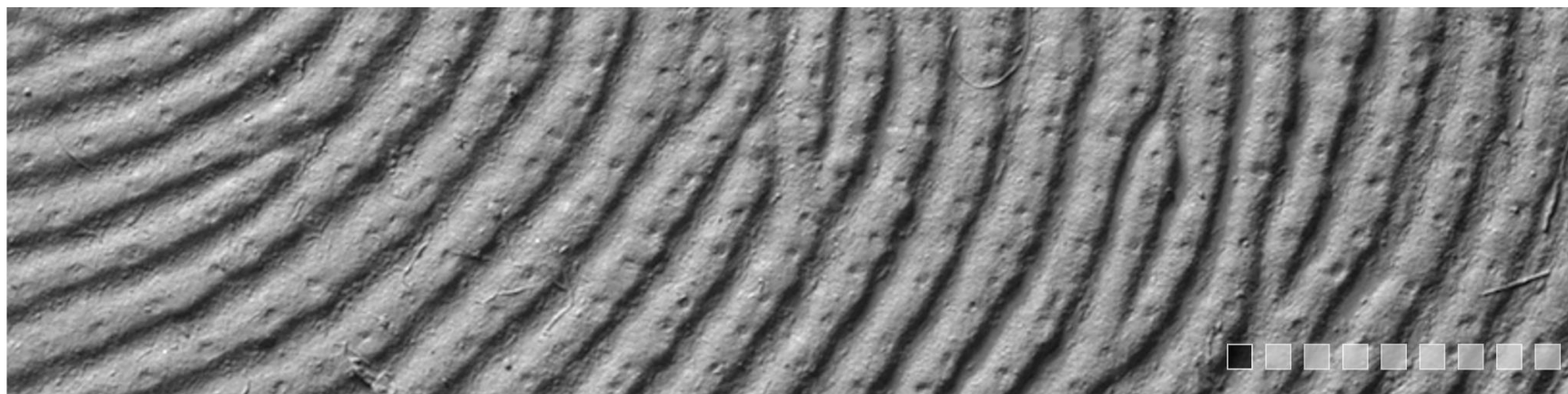
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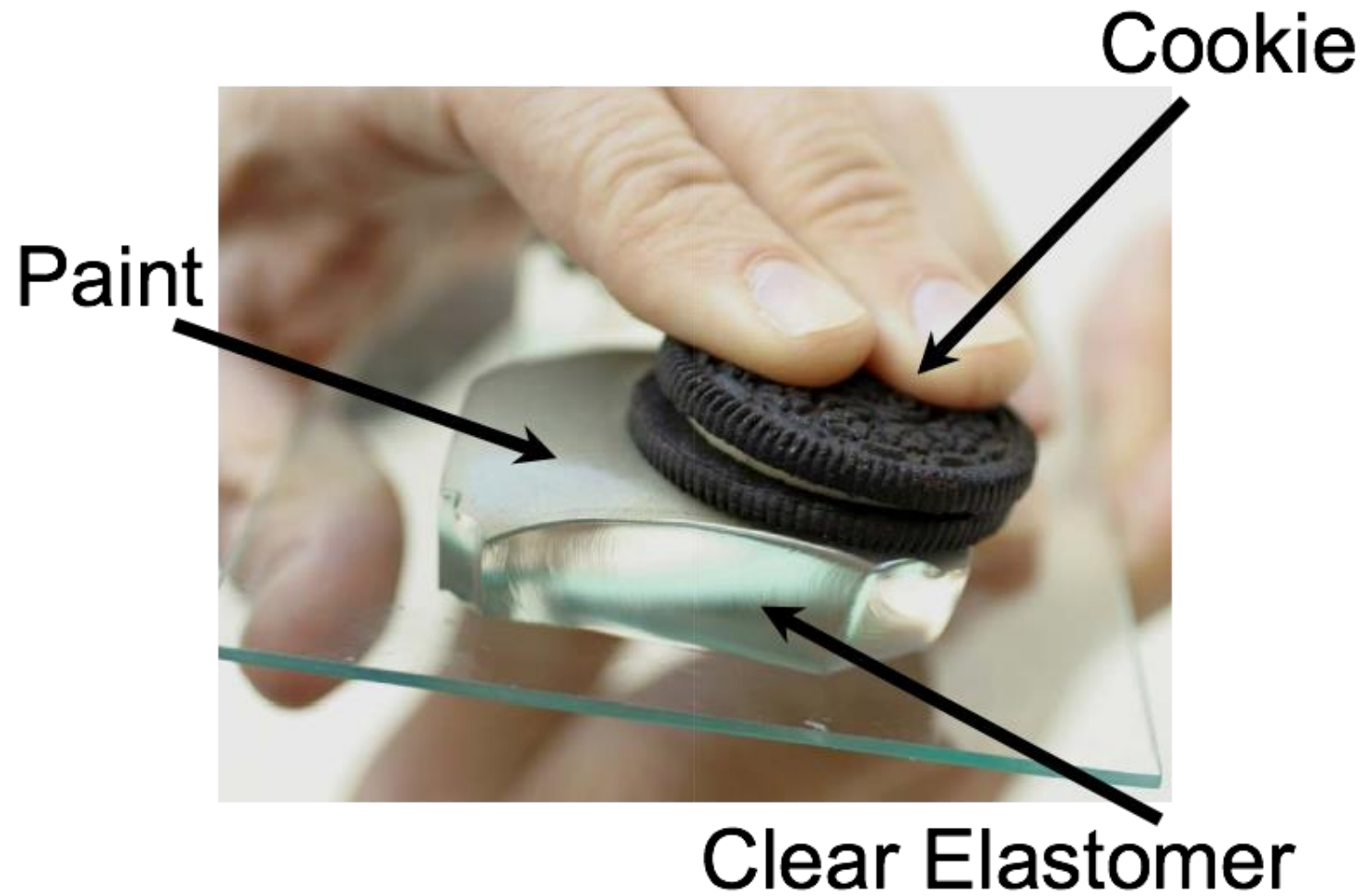
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Johnson and Adelson, 2009



Johnson and Adelson, 2009



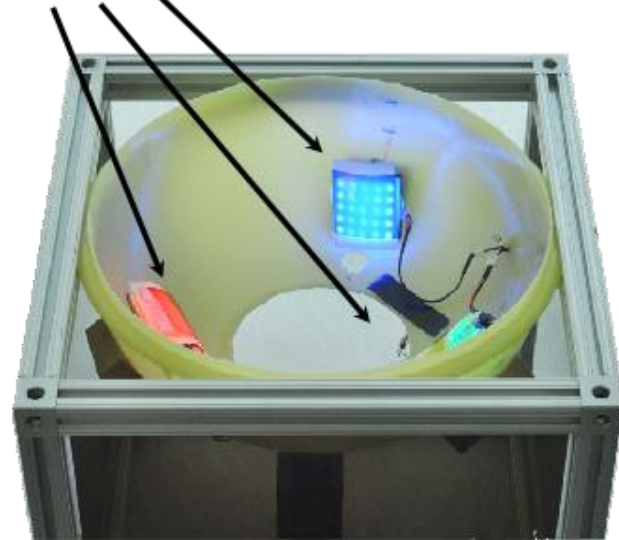


# Lights, camera, action

Sensor



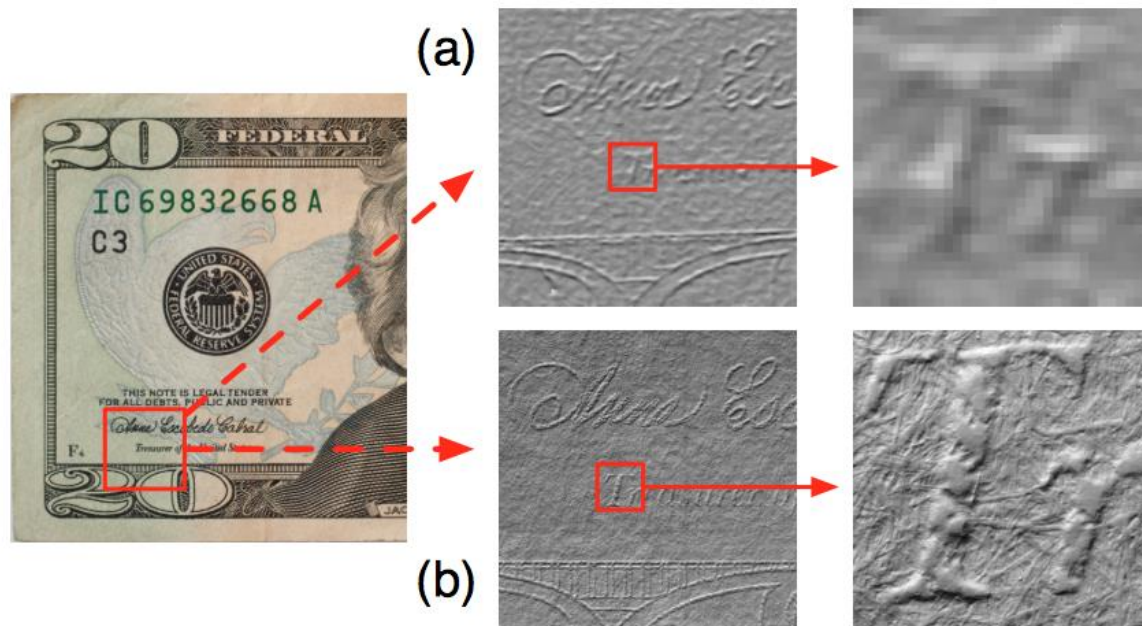
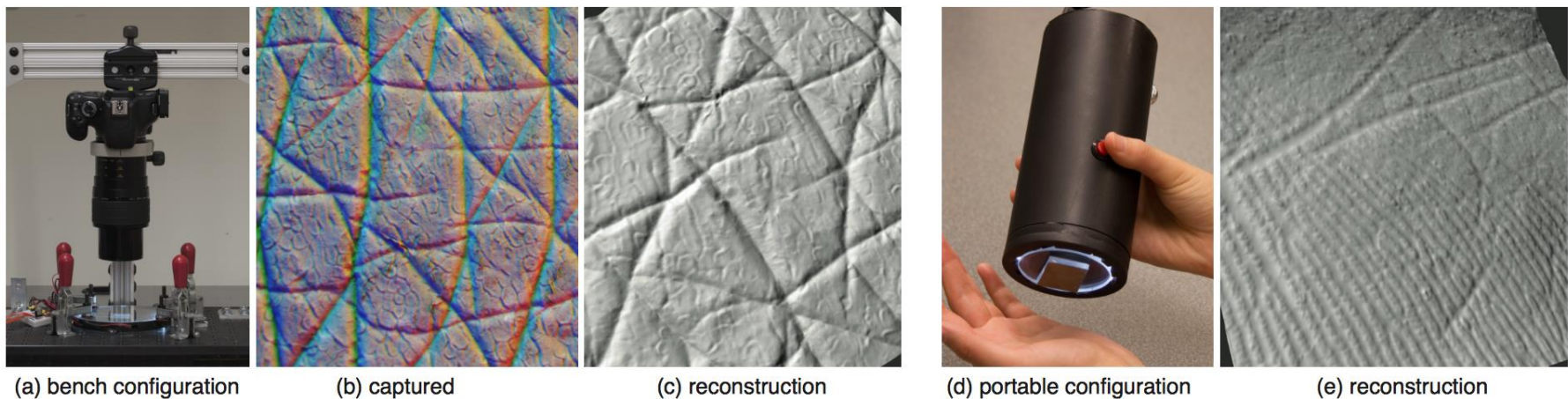
Lights



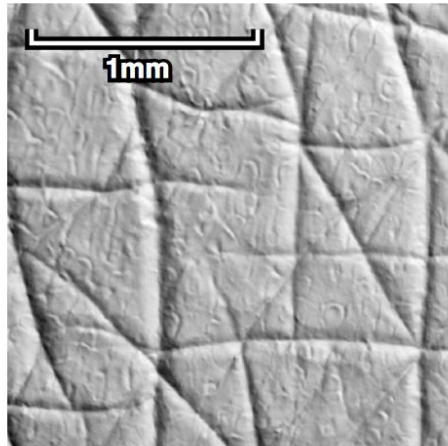
Camera



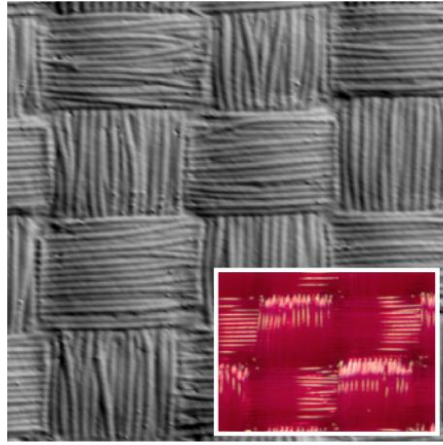




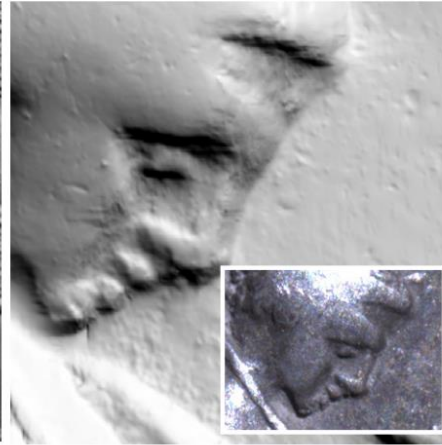
**Figure 7:** Comparison with the high-resolution result from the original retrographic sensor. (a) Rendering of the high-resolution \$20 bill example from the original retrographic sensor with a close-up view. (b) Rendering of the captured geometry using our method.



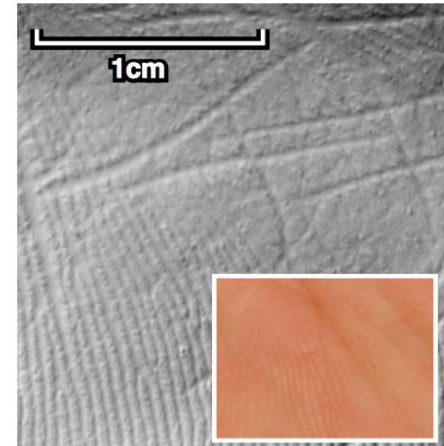
human skin



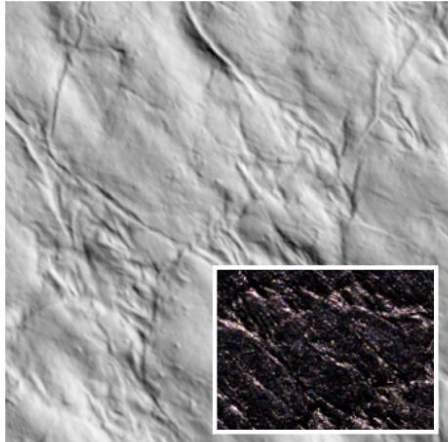
nylon fabric



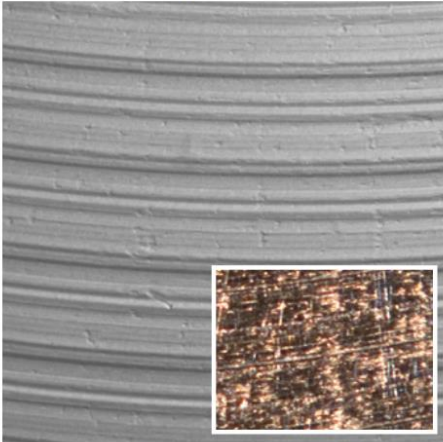
Greek coin



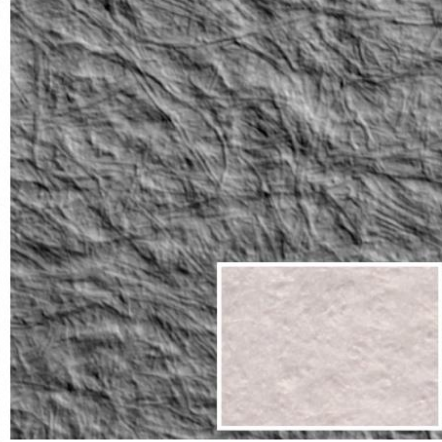
human skin



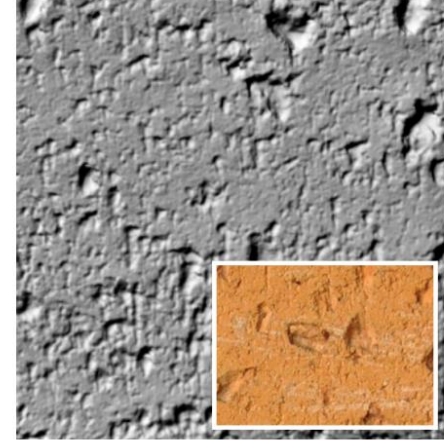
leather



vertically milled metal



paper



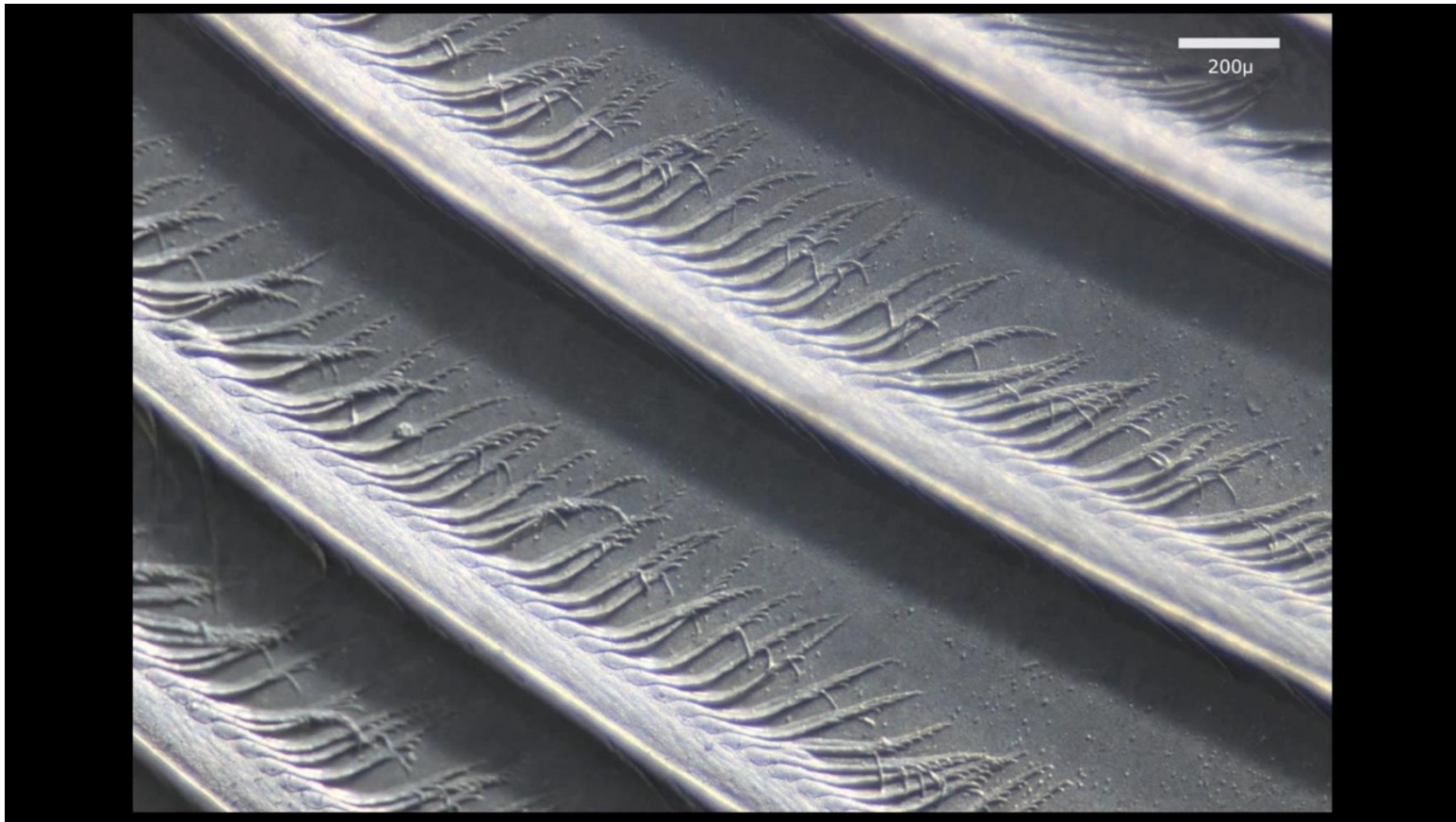
brick

(a) bench configuration

(b) portable configuration

**Figure 9:** Example geometry measured with the bench and portable configurations. Outer image: rendering under direct lighting. Inset: macro photograph of original sample. Scale shown in upper left. Color images are shown for context and are to similar, but not exact scale.





## Sensing Surfaces with GelSight



kimoatmit

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138,850 views

<https://www.youtube.com/watch?v=S7gXih4XS7A>

# InverseRenderNet: Learning single image inverse rendering

Ye Yu and William A. P. Smith

Department of Computer Science, University of York, UK

{yy1571, william.smith}@york.ac.uk

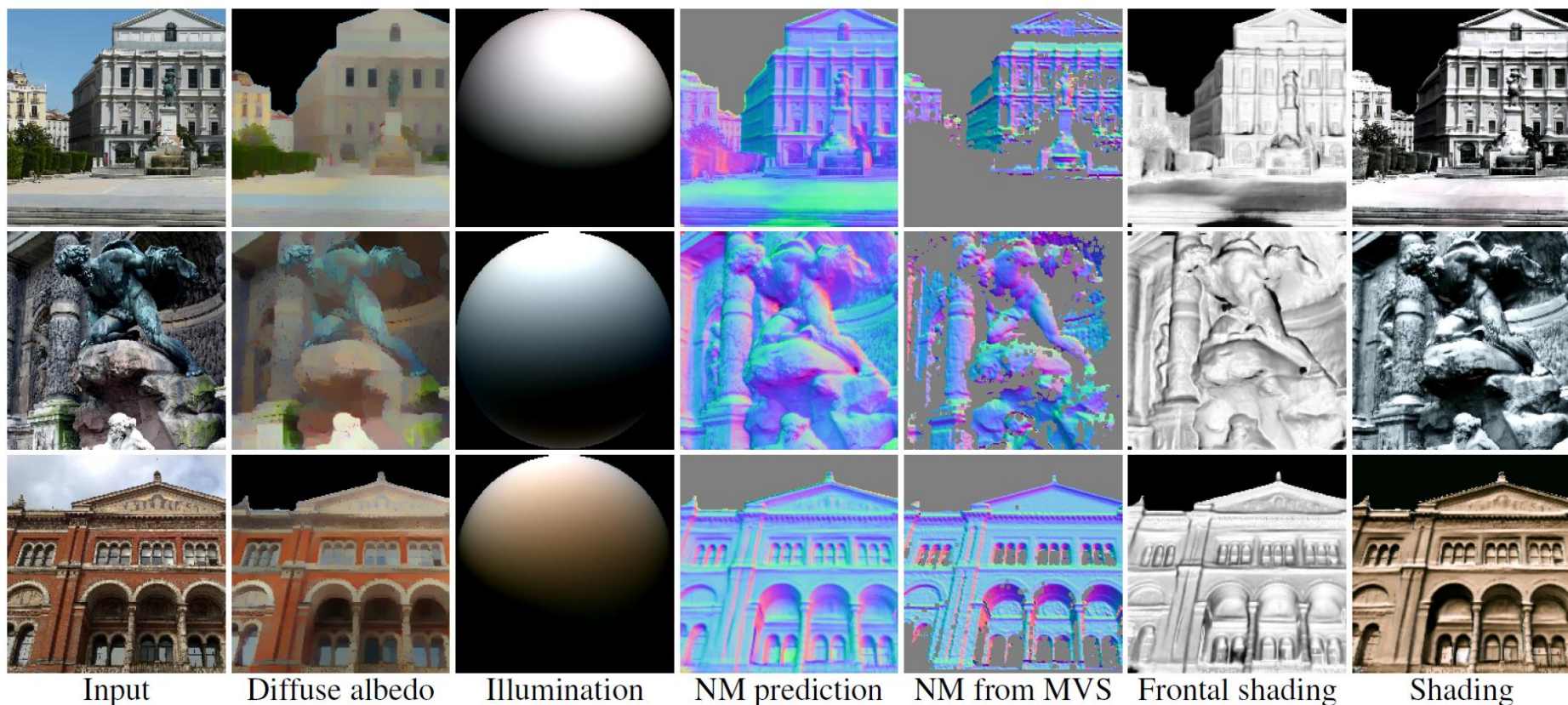


Figure 1: From a single image (col. 1), we estimate albedo and normal maps and illumination (col. 2-4); comparison multi-view stereo result from several hundred images (col. 5); re-rendering of our shape with frontal/estimated lighting (col. 6-7).

# Questions?