# Reasoning about Photo Collections using Models of Outdoor Illumination

Daniel Hauagge, Scott Wehrwein, Paul Upchurch, Noah Snavely, and Kavita Bala

{hauagge,swehrwein,paulu,snavely,kb}@cs.cornell.edu

### Motivation

- Outdoor photo collections are rich in information, but difficult to reason about in part due to complexity of outdoor illumination. towards understanding the world through photo collections.
- Models of sun/sky illumination have been developed in the graphics community, but rarely leveraged in computer vision.

### Contributions

- An analysis of the interactions between outdoor illumination and scene properties.
- A method for estimating ambient occlusion, albedo, and lighting in outdoor photo collections.
- A method for estimating the capture time of outdoor photos

### Prior Work: Photometric Ambient Occlusion

[Hauagge et al. CVPR 2013]

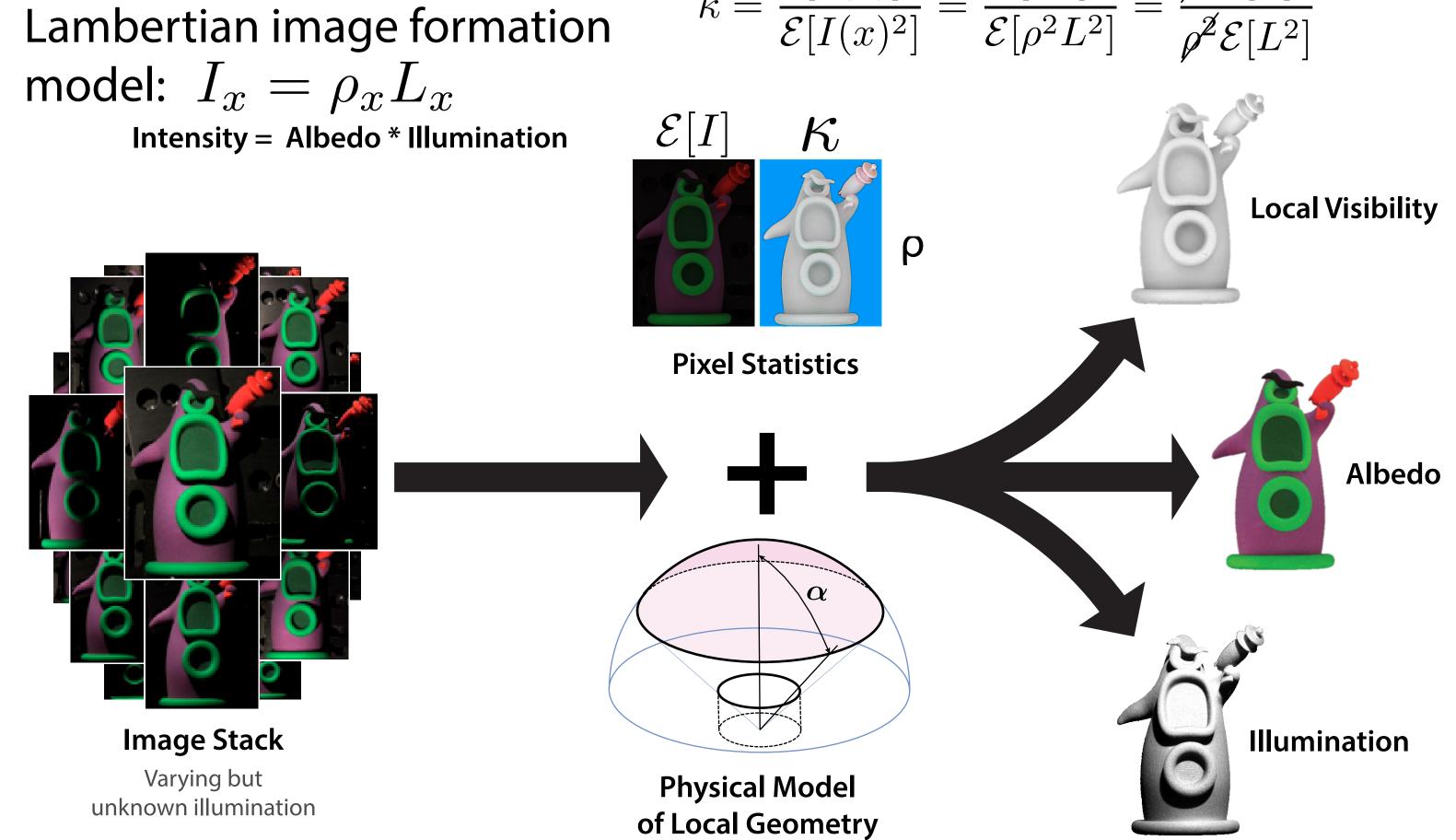
Albedo-invariant statistic:

Key limitation - lighting assumptions:

Uniformly distributed point source

Constant intensity and color

Constant white ambient term



Model local geometry as cylindrical crevice, parameterized by opening angle lpha

#### . Estimate $\alpha$ :

Analytically derive and invert  $\kappa(\alpha)$ , lookup observed  $\kappa$ 

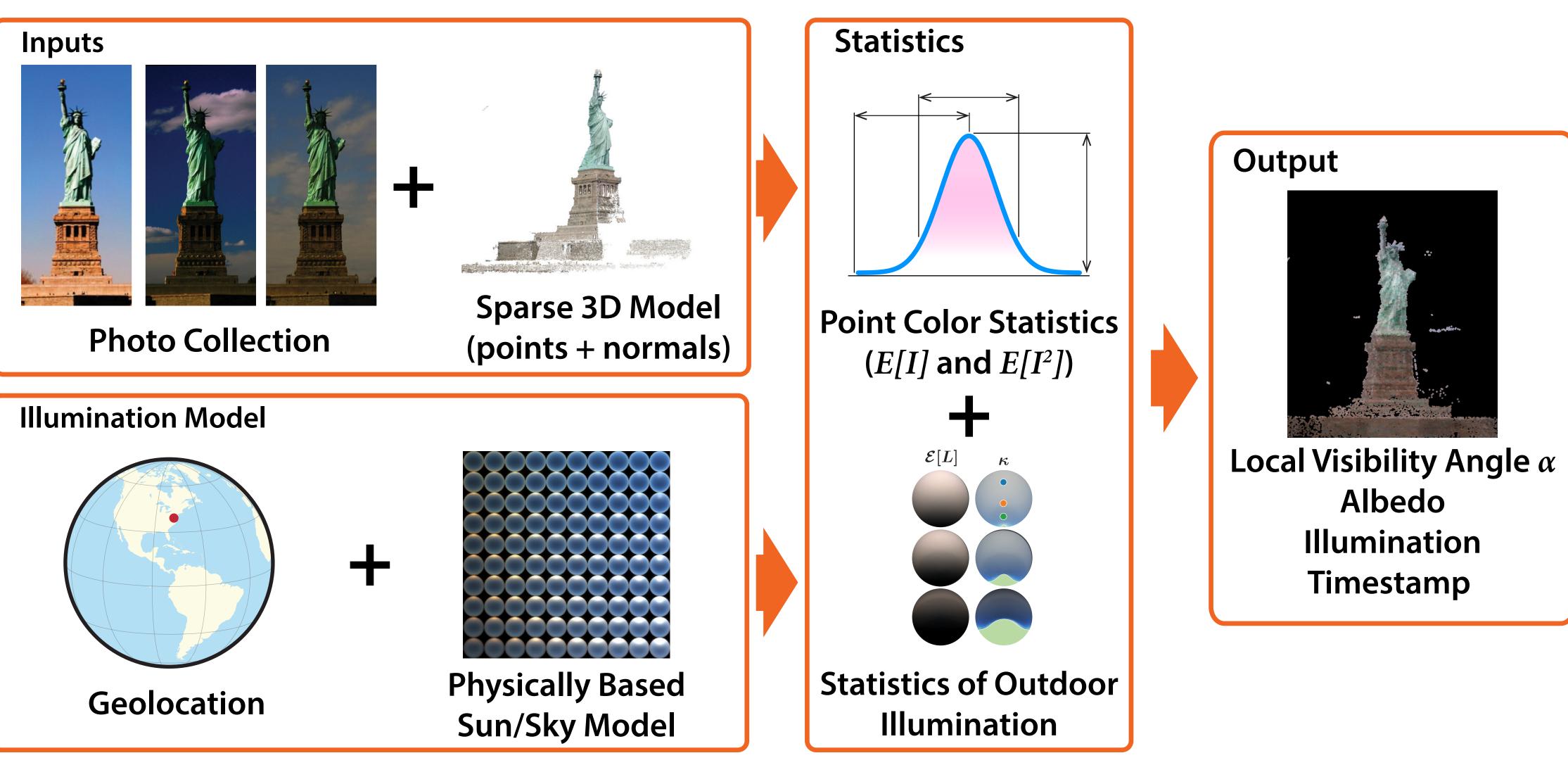
2. Estimate  $\rho$ :

(Observed)

3. Estimate L:

L = -

# **Algorithm Overview**



### **Outdoor Illumination Model**

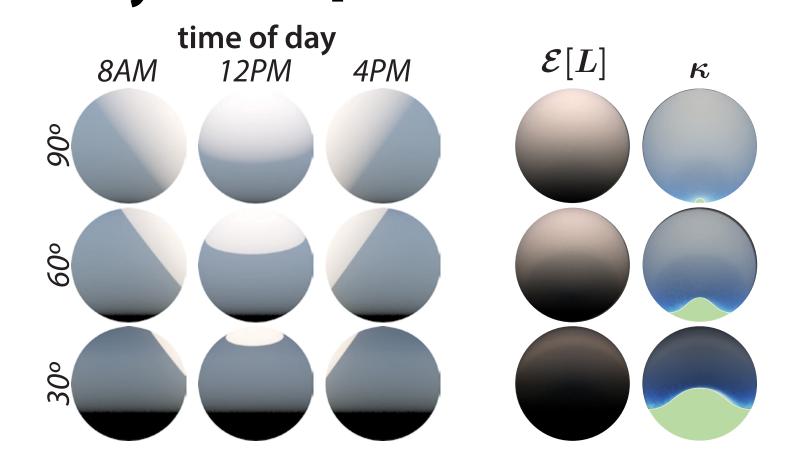
Sun/sky illumination at a point is determined by:

- Geographic Location  $(\phi, \lambda)$
- Time and Date (t)
- Ambient Occlusion  $(\alpha)$
- Surface Normal  $(\vec{n})$

 $L(\phi, \lambda, t, \alpha, \vec{n})$ 

# **Estimating Albedo**

**Key Idea: Predict statistics of natural** illumination using a physically-based sun/sky model [Hosek and Wilkie 2012]

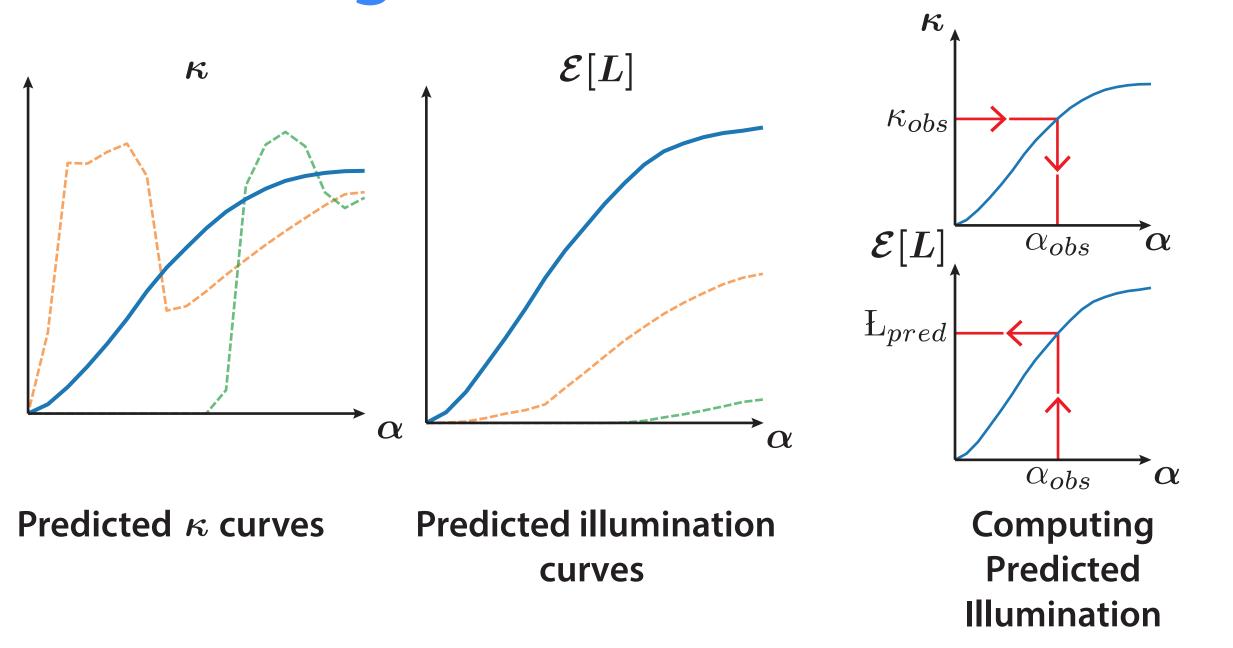


$$ho = rac{\mathcal{E}[I]}{\mathcal{E}[L(\phi,\lambda,t,lpha,ec{n})]}$$
 (Observed) (Predicted)

# 45° North **Equator**

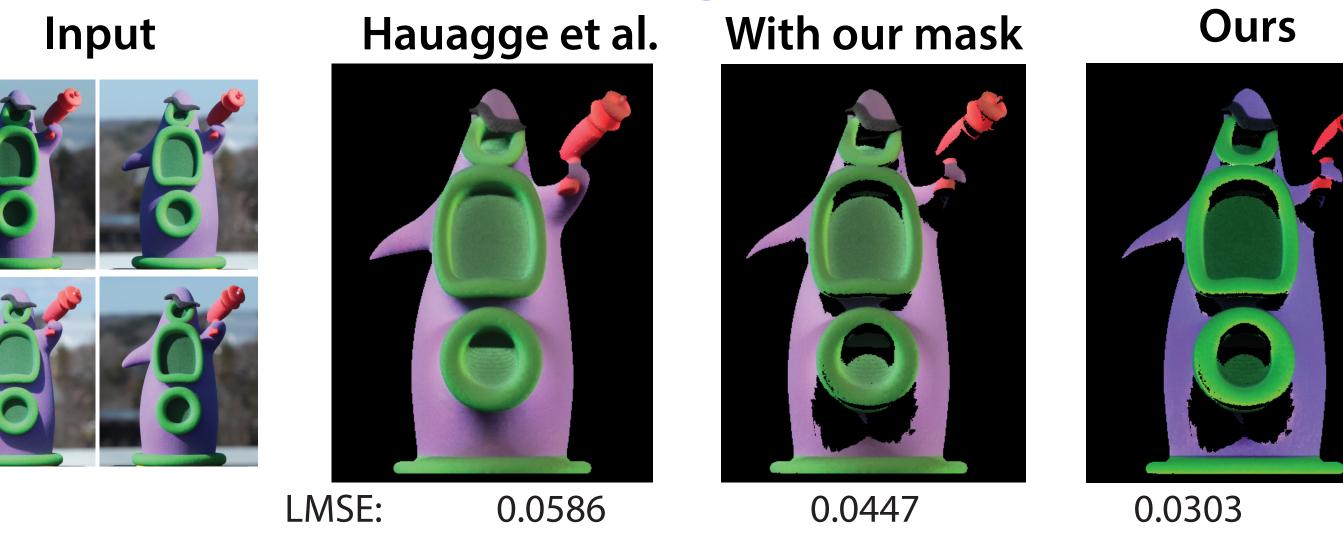
Sun visibility is determined by geographic location, time of day and year, crevice depth, and surface normal.

# Estimating Ambient Occlusion (a)



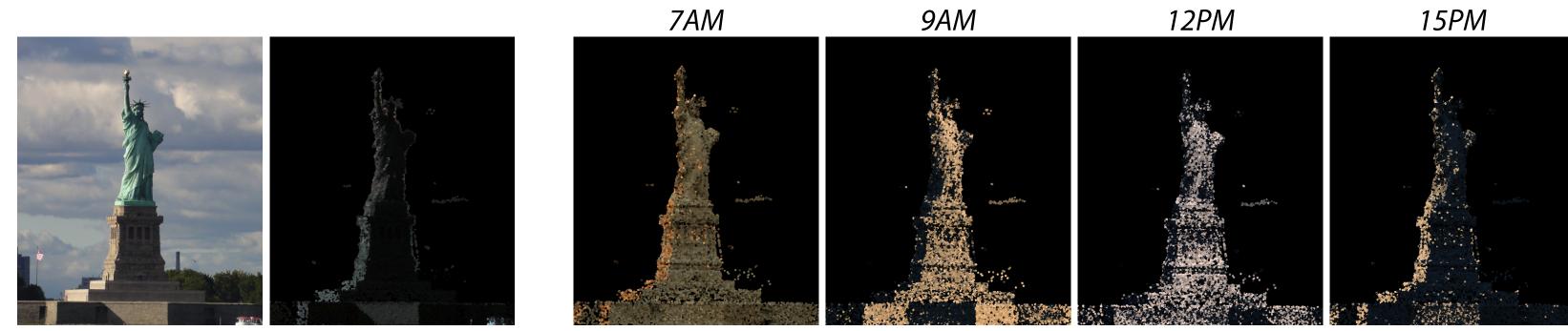
For an observed value of  $\kappa$ , we look up  $\alpha$  and the predicted illumination E[L], which allows us to estimate the albedo.

## Application: Intrinsic Image Decomposition

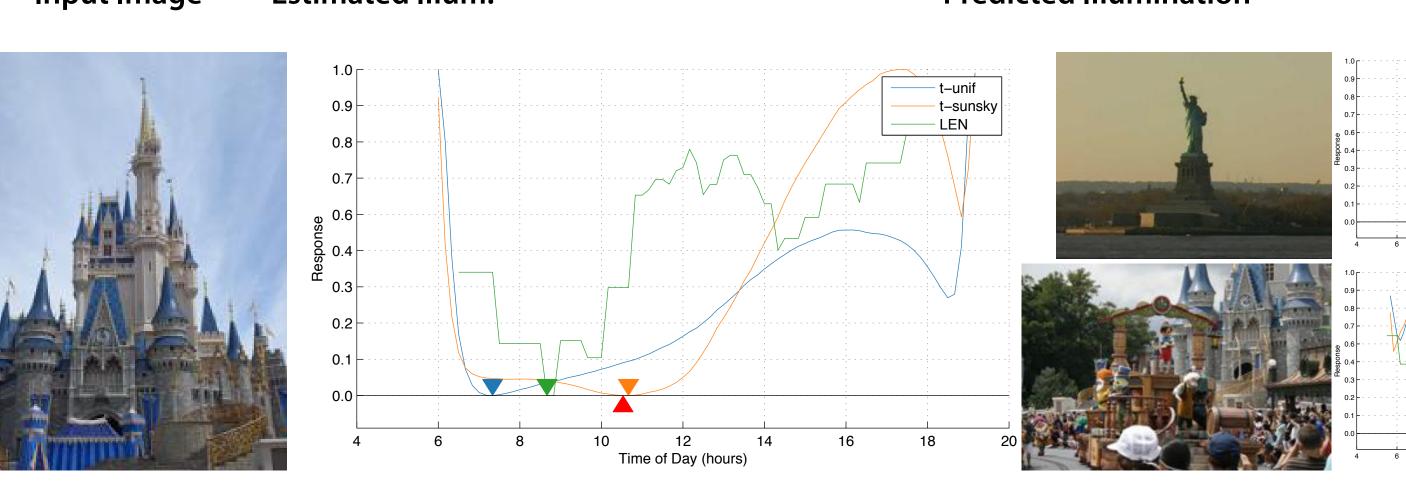


## **Application: Timestamp Estimation**

### Key Idea: Estimate illumination and find closest predicted illumination



**Predicted Illumination** 



t-sunsky: our method t-unif: albedo from [Hauagge et al., 2013] Rand: the average expected error of guessing a random time during the day **Exif:** timestamp stored in image LEN: method proposed in [Lalonde et al., Timestamp errors (in minutes) IJCV2011]

## **Visualizing Timestamp Results**

Exif (baseline)





#### t-sunsky (our method) Reference Lighting