

Lecture 3: Radiometry

Fall 2004
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Information

- Office Hours
 - Wed: 2-3 Upson 5142
- Web-page
- www.cs.cornell.edu/courses/cs665/2004fa/
 - Tentative schedule
 - Homeworks, lecture notes, will be on-line
 - Check for updates and announcements

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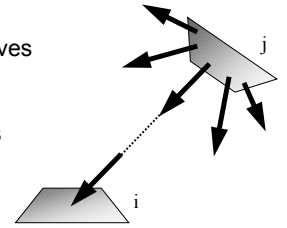
Radiosity Algorithm

- Subdivide scene in polygons
 - mesh that determines final solution
- Compute Form Factors
 - transfer of energy between polygons
- Solve linear system
 - results in power (color) per polygon
- Pick a viewpoint and display
 - loop

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Form Factor

- $F_{j \rightarrow i}$ = the fraction of power emitted by j , which is received by i
- Area
 - if i is smaller, it receives less power
- Orientation
 - if i faces j , it receives more power
- Distance
 - if i is further away, it receives less power



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Form Factors - how to compute?

- Closed Form
 - Analytic
- Hemicube
- Monte-Carlo

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How To Solve Linear System

- Matrix Inversion
- Gathering methods
 - Jacobi iteration
 - Gauss-Seidel
- Shooting
 - Southwell iteration
 - Improved Southwell iteration

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Matrix Inversion

$$\begin{bmatrix} B_1 \\ B_2 \\ \dots \\ B_n \end{bmatrix} = \begin{bmatrix} 1 - \rho_1 F_{1 \rightarrow 1} & -\rho_1 F_{1 \rightarrow 2} & \dots & -\rho_1 F_{1 \rightarrow n} \\ -\rho_2 F_{2 \rightarrow 1} & 1 - \rho_2 F_{2 \rightarrow 2} & \dots & -\rho_2 F_{2 \rightarrow n} \\ \dots & \dots & \dots & \dots \\ -\rho_n F_{n \rightarrow 1} & -\rho_n F_{n \rightarrow 2} & \dots & 1 - \rho_n F_{n \rightarrow n} \end{bmatrix}^{-1} \begin{bmatrix} B_{e,1} \\ B_{e,2} \\ \dots \\ B_{e,n} \end{bmatrix}$$

$O(n^3)$

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Jacobi

- For all patches i ($i=1 \dots N$) : $B_i^{(0)} = B_{e,i}$

- while not converged:
 - for all patches i ($i=1 \dots N$)

$$B_i^{(g)} = B_{e,i} + \rho_i \sum_{j=1}^N B_j^{(g-1)} F(i \rightarrow j)$$

update of 1 patch requires evaluation of N FFs

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Improved Gathering

- Jacobi iteration only uses values of previous iterations to compute new values
- Gauss-Seidel iteration
 - New values used immediately
 - Slightly better convergence

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Gauss-Seidel

- For all patches i ($i=1 \dots N$) : $B_i^{(0)} = B_{e,i}$

- while not converged:
 - for all patches i ($i=1 \dots N$)

$$B_i^{(g)} = B_{e,i} + \rho_i \sum_{j=1}^{i-1} B_j^{(g)} F(i \rightarrow j) + \rho_i \sum_{j=i}^N B_j^{(g-1)} F(i \rightarrow j)$$

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Southwell Iteration

- “Shooting” method
- Start with initial guess for light distribution (light sources)
- Select patch and distribute its energy over all polygons

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Progressive Refinement

- Southwell selects shooting patches in no particular order
- Progressive refinement radiosity selects patch with largest unshot energy
- First image is generated fairly quickly!

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PR + Ambient term

- PR gives an estimate for each radiosity value that is smaller than the real value
- Estimate can be improved by using ambient term
 - Add all unshot energy
 - Distribute total unshot energy equally over all patches
- Solution has improved energy distribution

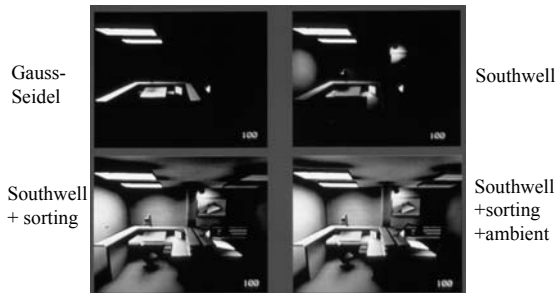
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Gathering vs. Shooting

- Gathering
 - Jacobi Gauss-Seidel
- Shooting
 - Sowell Progressive Radiosity

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Comparison



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Radiosity Algorithms

- Object (scene) based
- Assumptions
 - Polygons
 - Diffuse BRDFs
 - Diffuse light sources
 - Static scenes
 - “Constant polygon” assumption does not capture high frequency illumination (e.g. shadow cast by a fence)

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Radiosity

- Does not handle non-diffuse surfaces
- “Constant polygon” assumption does not capture high frequency illumination (e.g. shadow cast by a fence)
- Non-polygonal objects are a problem

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Improvements

- Hierarchical Radiosity
- Discontinuity-Driven Radiosity

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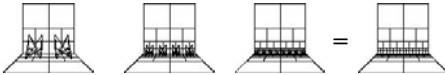
Hierarchical Radiosity

- Iterative techniques are $O(n^2)/\text{iteration}$
- Insight:
 - Radiosity is similar to N-body problem
 - Form Factor F proportional to $1/r^2$
 - Use N-body algorithms to get $O(n)$

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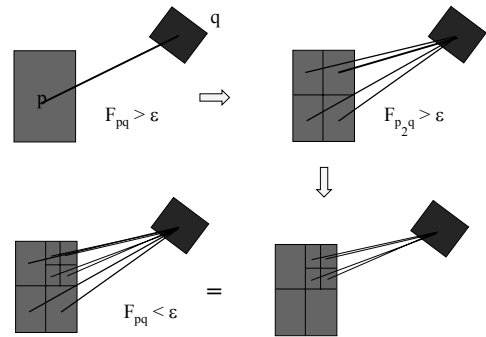
HR: Main idea

- Hierarchical refinement of patches
- Interactions at different levels of hierarchy
- Higher level interactions replace whole blocks of element-element interactions



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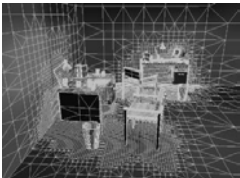
Refining a Link



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Algorithm

- Top-down
- Each node connected to constant # of nodes
- Number of interactions $O(k^2 + n)$
 - Note that k^2 interactions required



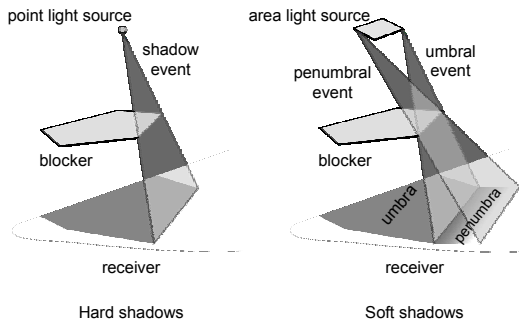
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Discontinuity Meshing



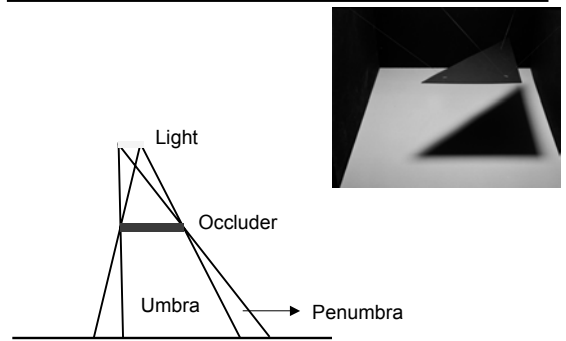
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Shadows: Hard and Soft



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Discontinuity Meshing



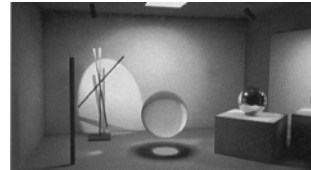
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Hierarchical Radiosity with Discontinuity Meshing



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Complex BRDFs



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Summary

- Discrete form of rendering equation
- Form Factor computations
- Different ways of solving the linear system (Jacobi, Gauss Seidel, Southwell)
- HR, Importance, Discontinuities

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CS 665
Radiometry and
the Rendering Equation
Chapter 2 in Advanced GI

What is this course about

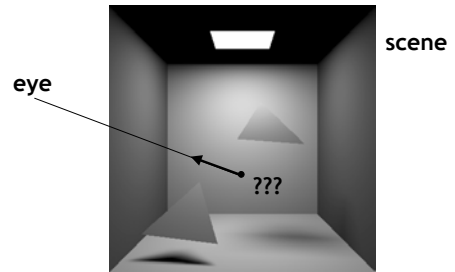
- What does image generation mean?
 - Physics of light
- How to generate images?
 - Global illumination algorithms
- How do we do this efficiently?
 - Alternative representations, data structures, image-based rendering etc.



RTGI, Program of Computer Graphics

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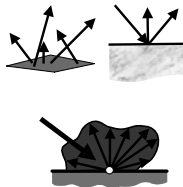
Motivation



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What is the behavior of light?

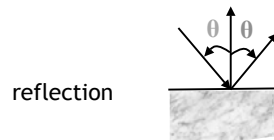
- Physics of light
- Radiometry
- Material properties
- **Rendering Equation**



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Brief history of Optics: 350 B.C.

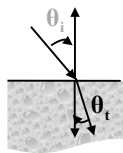
- Greek philosophers (350 B.C.)
 - Pythagoras, Plato, Aristotle
 - Light emanated from the eye
- By 300 B.C.,
 - Understood rectilinear propagation of light
 - Euclid described law of reflection



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Brief history of Optics: Dark Ages

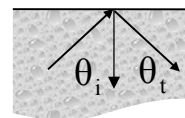
- Ptolemy (130 A.D.)
 - Refraction
- Not much in Dark Ages except for Alhazen (elaborated law of reflection)
 - Angles of incidence and reflection in the same plane as normal N



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Brief history of Optics: 17th century

- 17th century: telescopes, microscopes
- Kepler (1611)
 - Total internal reflection



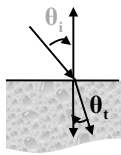
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Brief history of Optics: 17th century

- Snell (1621): Law of refraction

$$\frac{\sin \theta_i}{\sin \theta_t} = \frac{\eta_t}{\eta_i}$$

$$\eta_i \sin \theta_i = \eta_t \sin \theta_t$$

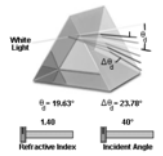


- Descartes published the sine form of law of refraction
- Fermat (1657)
 - Law of refraction from *principle of least time*

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Brief history of Optics: 17th century

- Newton (1642-1727)
 - Dispersion
 - Light splits into component colors
 - Corpuscular/emission theory

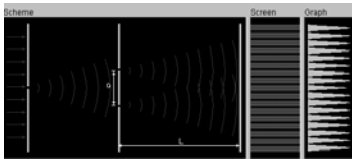


- Huygens (1629-1695)
 - Developed wave theory of light
 - Discovered polarization
- Two conflicting theories: wave vs. emission

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Brief history of Optics (contd.)

- Young (1801)
 - Principle of interference
 - Double-slit experiment



- Fresnel
 - 1816: Diffraction and interference
 - 1821: Fresnel equations for reflection and refraction

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Brief history of Optics (contd.)

- Maxwell (1831-1879)
 - Electricity and magnetism
 - Maxwell's equations
 - $\nabla \cdot \mathbf{E} = 0$
 - $\nabla \cdot \mathbf{B} = 0$
 - $\nabla \times \mathbf{E} = -\partial \mathbf{B} / \partial t$
 - $\nabla \times \mathbf{B} = \mu_0 \epsilon_0 \partial \mathbf{E} / \partial t$
 - Theoretical validation of measured speed of light
- Light is electromagnetic radiation

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Brief history of Optics (contd.)

- Hertz (1887)
 - Discovered the photoelectric effect
 - The process whereby electrons are liberated from materials under the action of radiant energy*
- Einstein (1905)
 - Explained photoelectric effect
 - Light is a stream of quantized energy packets called quanta (photons)
 - $E = h \nu$

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Brief history of Optics (contd.)

- Particle and wave theory seemingly mutually exclusive
- Quantum Mechanics
 - Bohr, Born, Heisenberg, Schrodinger, Pauli
 - Sub-microscopic phenomena
 - Unite particle and wave behavior of light
 - QED by Feynman

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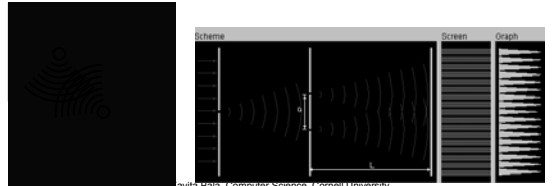
Models of Light

- Geometric Optics
 - Emission
 - Reflection / Refraction
 - Absorption
- Simplest model
- Size of objects > wavelength of light

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Models of light

- Wave Model
 - Maxwell's Equations
 - Object size comparable to wavelength
 - Diffraction, Interference, Polarization



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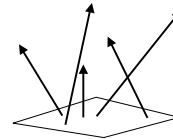
Models of light

- Quantum Model
 - Fluorescence
 - Phosphorescence
 - Relativistic effects
- Most complete model

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Geometric Optics: Properties

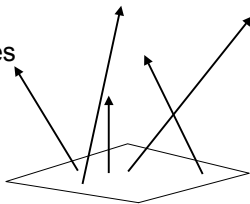
- Light travels in straight lines
- Rays do not interact with each other
- Rays have color(wavelength), intensity



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Emission

- New rays due to:
 - chemical
 - electrical
 - nuclear processes

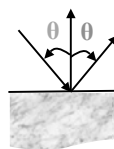


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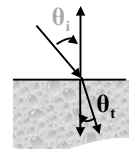
Reflections/Refractions

- Interface between 2 materials

reflection



refraction

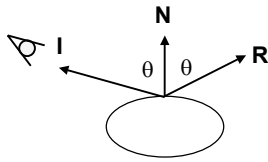


- Specular reflections and refractions
 - One direction

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Specular Reflections

Reflected ray R

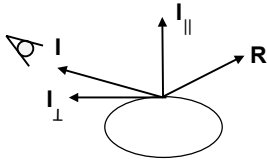


$$I_{\parallel} = (I \cdot N) N$$

$$I_{\perp} = I - (I \cdot N) N$$

$$R = I_{\parallel} + (-I_{\perp})$$

$$= 2 (I \cdot N) N - I$$



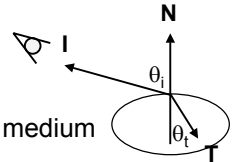
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Specular Refractions

- Compute refracted ray T
- Index of refraction: $n = \frac{c}{v} = \frac{c}{\frac{c}{\eta}} = \eta$
- Snell's law:

$$\frac{\sin \theta_i}{\sin \theta_t} = \frac{n_2}{n_1}$$

$$n_1 \sin \theta_i = n_2 \sin \theta_t$$

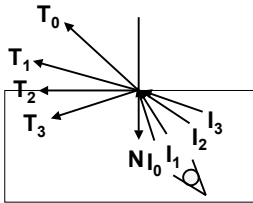


- Ray from rare to dense medium

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Total Internal Reflection

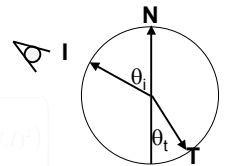
- Consider ray from dense to rare



$$\frac{\sin \theta_c}{1} = \frac{n_2}{n_1}$$

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Transmitted Ray



$$T = \frac{2(n_1 n_2)}{n_1 + n_2} \left[\frac{2(n_1 - n_2)}{n_1 + n_2} \cos \theta_i - \sqrt{1 - \left(\frac{n_2}{n_1} \right)^2 \sin^2 \theta_i} \right] N$$

$$T = \frac{2(n_1 n_2)}{n_1 + n_2} \left[\cos \theta_i - \sqrt{1 - \left(\frac{n_2}{n_1} \right)^2 \sin^2 \theta_i} \right] N$$

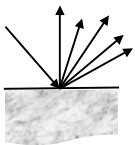
$$\cos \theta_t = \left(\frac{n_1}{n_2} \right) \cos \theta_i$$

$$n = \frac{c}{v}$$

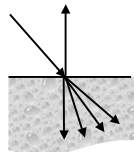
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Realistic Reflections/Refractions

reflection

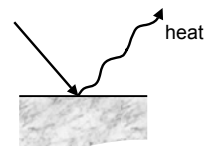


refraction



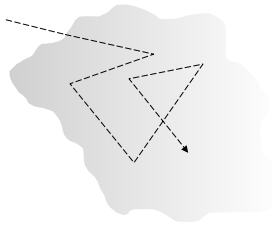
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Absorption



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Interaction with matter



participating medium

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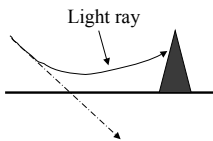
Interaction with matter



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Interaction with matter

- Continuously varying refractive index



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Interaction with matter



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