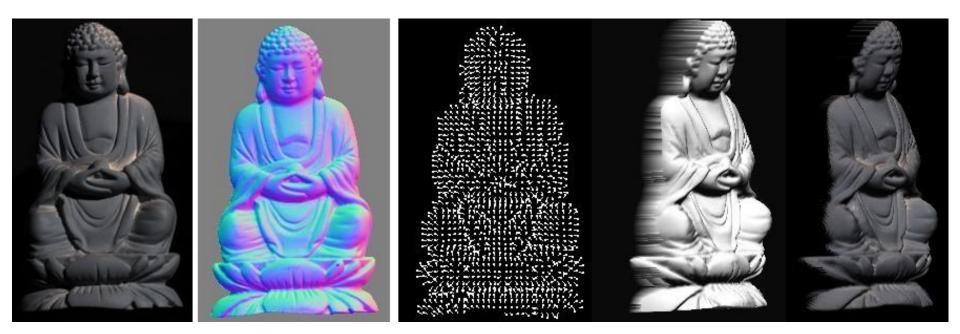
CS5670 : Computer Vision Noah Snavely

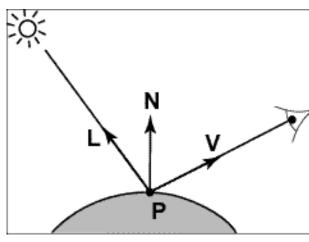
Photometric stereo



Announcements

- Project 3 code due tomorrow, 3/29, by 11:59pm to CMS
 - Artifact due Friday, 3/30, by 11:59pm to CMS

Recap: Lambertian (Diffuse) Reflectance





 $I = k_d \mathbf{N} \cdot \mathbf{L}$

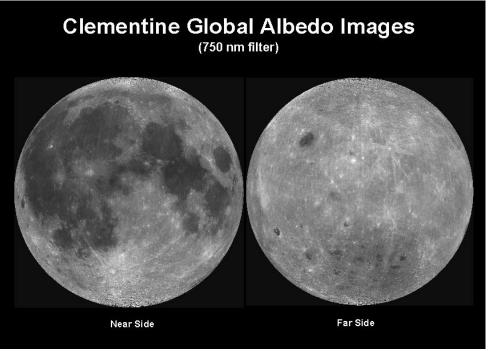
- *I* : observed image intensity
- k_d : object albedo
- N : surface normal
- L : light source direction



Lambertian sphere with constant albedo lit by a directional light source

Sample albedos

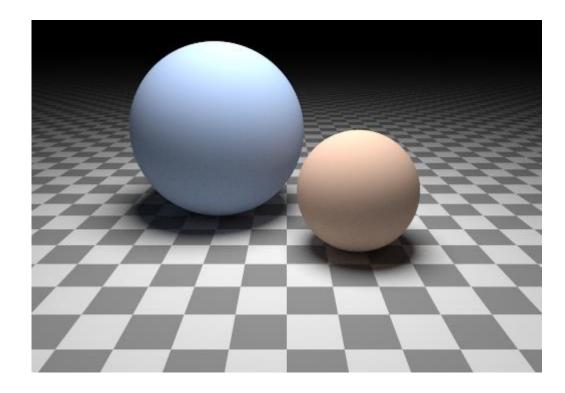
Surface	Typical albedo
Fresh asphalt	0.04 ^[4]
Open ocean	0.06 ^[5]
Worn asphalt	0.12 ^[4]
Conifer forest (Summer)	0.08, ^[6] 0.09 to 0.15 ^[7]
Deciduous trees	0.15 to 0.18 ^[7]
Bare soil	0.17 ^[8]
Green grass	0.25 ^[8]
Desert sand	0.40 ^[9]
New concrete	0.55 ^[8]
Ocean ice	0.5–0.7 ^[8]
Fresh snow	0.80–0.90 ^[8]



Objects can have varying albedo and albedo varies with wavelength

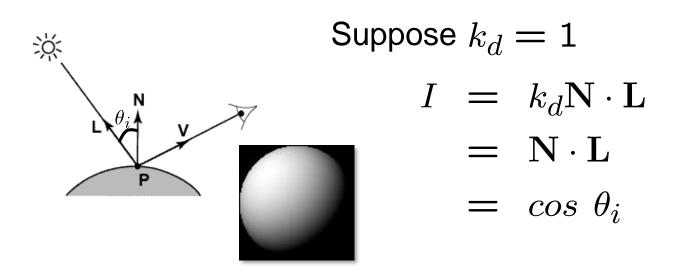
Source: <u>https://en.wikipedia.org/wiki/Albedo</u>

Can we determine shape from lighting?



- Are these spheres?
 - Or just flat discs painted with varying albedo?

Shape from shading



You can directly measure angle between normal and light source

- Not quite enough information to compute surface shape
- But can be if you add some additional info, for example
 - assume a few of the normals are known (e.g., along silhouette)
 - constraints on neighboring normals—"integrability"
 - smoothness
- Hard to get it to work well in practice
 - plus, how many real objects have constant albedo?

Application: Detecting composite photos Real photo

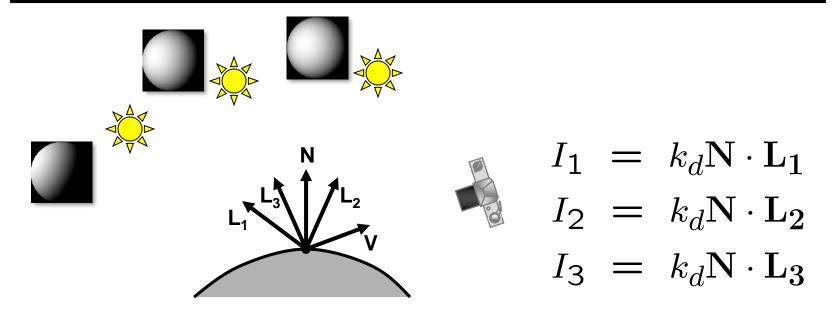
Fake photo





Let's take more than one photo!

Photometric stereo



Can write this as a matrix equation:

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = k_d \begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix} \mathbf{N}$$

Solving the equations

$$\begin{bmatrix} I_{1} \\ I_{2} \\ I_{3} \end{bmatrix} = \begin{bmatrix} \mathbf{L}_{1}^{T} \\ \mathbf{L}_{2}^{T} \\ \mathbf{L}_{3}^{T} \end{bmatrix} k_{d} \mathbf{N}$$

$$\mathbf{\mathbf{M}}$$

 $\mathbf{N} = \frac{1}{k_d} \mathbf{G}$

Solve one such linear system per pixel to solve for that pixel's surface normal

More than three lights

Get better results by using more lights

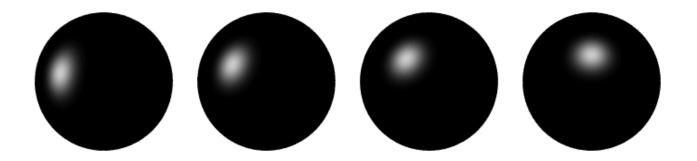
$$\begin{bmatrix} I_1 \\ \vdots \\ I_n \end{bmatrix} = \begin{bmatrix} \mathbf{L}_1 \\ \vdots \\ \mathbf{L}_n \end{bmatrix} k_d \mathbf{N}$$

Least squares solution:

$$\begin{split} \mathbf{I} &= \mathbf{L}\mathbf{G} \\ \mathbf{L}^T\mathbf{I} &= \mathbf{L}^T\mathbf{L}\mathbf{G} \\ \mathbf{G} &= (\mathbf{L}^T\mathbf{L})^{-1}(\mathbf{L}^T\mathbf{I}) \\ \end{split}$$
 Solve for N, k_d as before
$$\end{split}$$
 What's the size of L^TL?

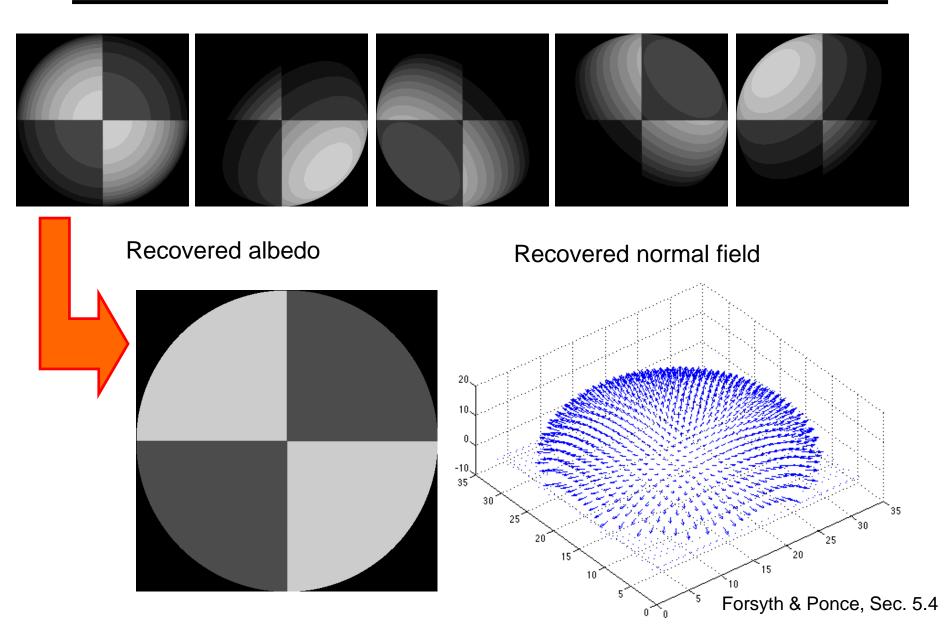
Computing light source directions

Trick: place a chrome sphere in the scene



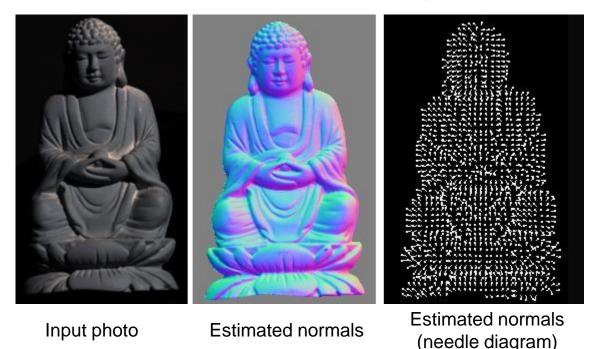
• the location of the highlight tells you where the light source is

Example



Depth from normals

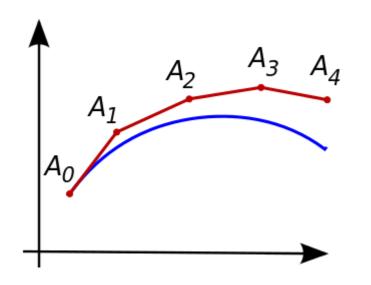
• Solving the linear system per-pixel gives us an estimated surface normal for each pixel



- How can we compute depth from normals?
 - Normals are like the "derivative" of the true depth

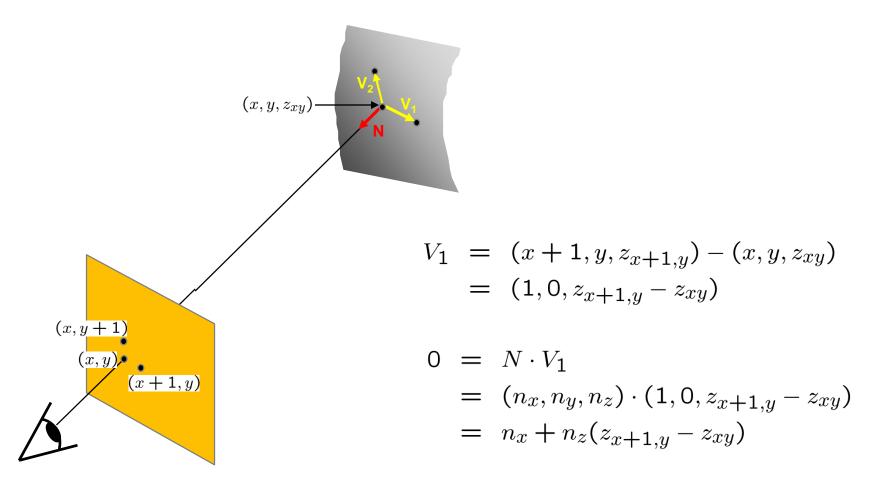
Normal Integration

- Integrating a set of derivatives is easy in 1D
 - (similar to Euler's method from diff. eq. class)



- Could just integrate normals in each column / row separately
- Instead, we formulate as a linear system and solve for depths that *best agree with the surface normals*

Depth from normals



Get a similar equation for V_2

- Each normal gives us two linear constraints on z
- compute z values by solving a matrix equation

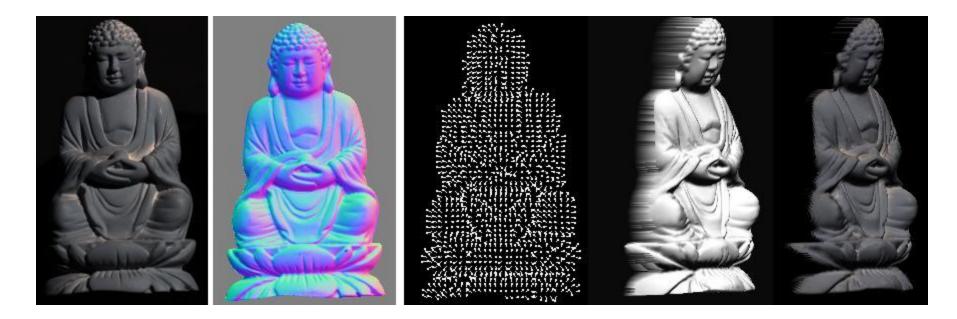
Results





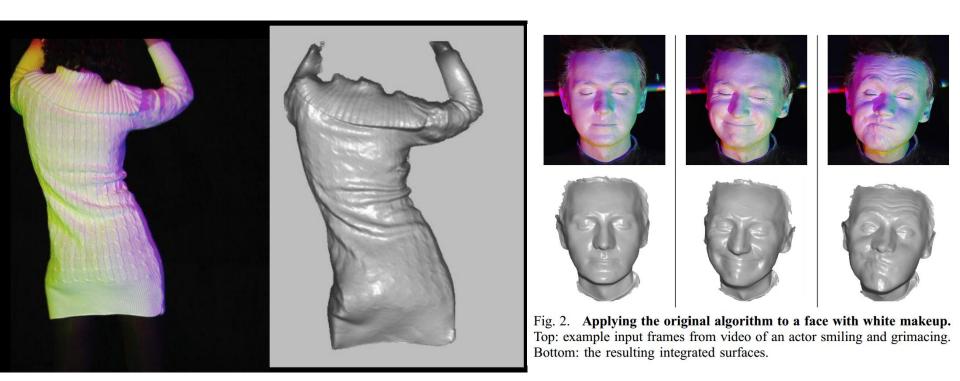
from Athos Georghiades

Results



Extension

Photometric Stereo from Colored Lighting



Video Normals from Colored Lights

Gabriel J. Brostow, Carlos Hernández, George Vogiatzis, Björn Stenger, Roberto Cipolla <u>IEEE TPAMI</u>, Vol. 33, No. 10, pages 2104-2114, October 2011.

Questions?

For now, ignore specular reflection



Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

And Refraction...



Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

And Interreflections...



Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

And Subsurface Scattering...



Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

Limitations

Bigger problems

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections

Smaller problems

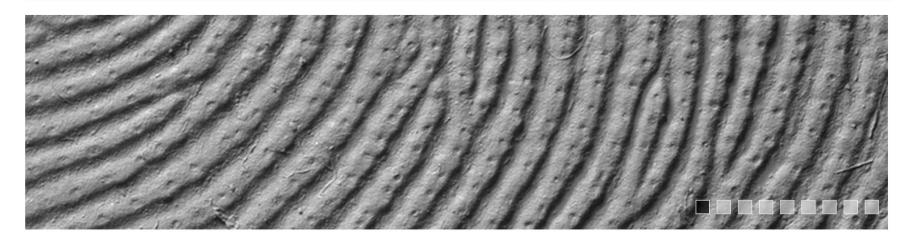
- camera and lights have to be distant
- calibration requirements
 - measure light source directions, intensities
 - camera response function

Newer work addresses some of these issues

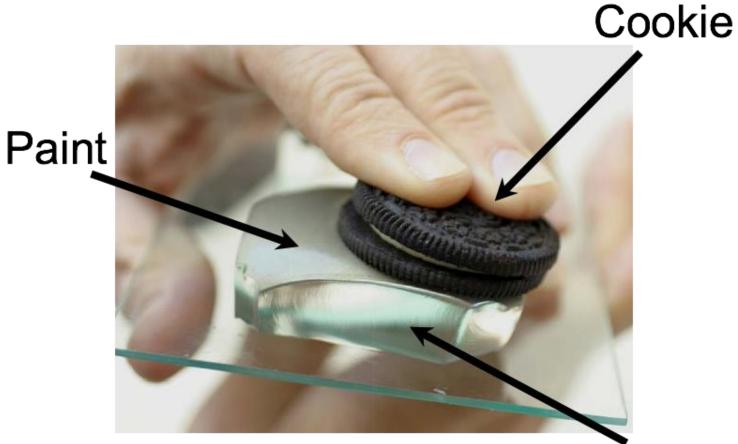
Some pointers for further reading:

- Zickler, Belhumeur, and Kriegman, "<u>Helmholtz Stereopsis: Exploiting</u> <u>Reciprocity for Surface Reconstruction</u>." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann & Seitz, "<u>Example-Based Photometric Stereo: Shape</u> <u>Reconstruction with General, Varying BRDFs</u>." IEEE Trans. PAMI 2005





Johnson and Adelson, 2009



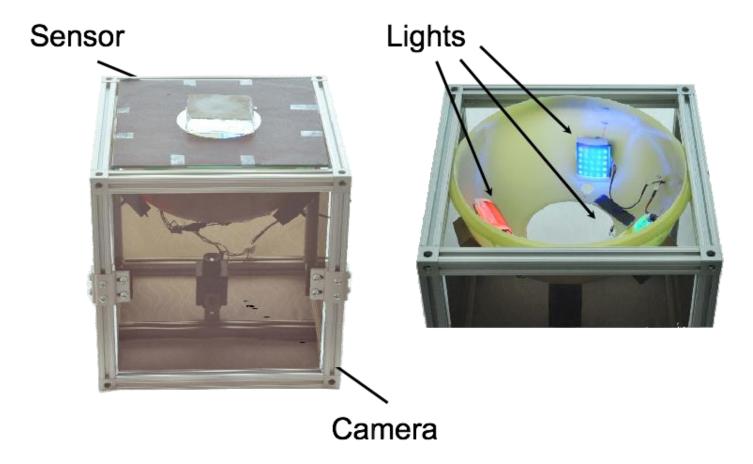
Clear Elastomer

Johnson and Adelson, 2009





Lights, camera, action





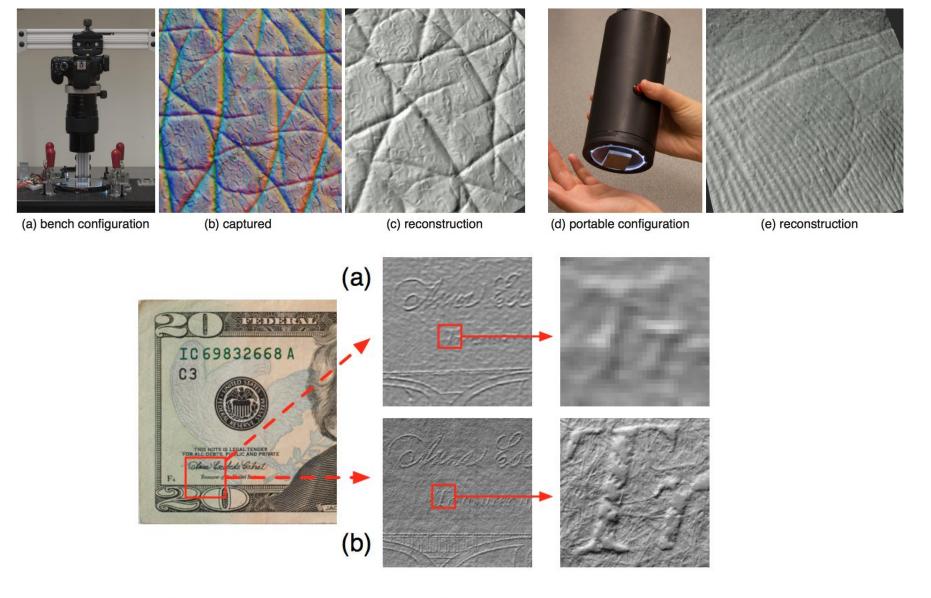


Figure 7: Comparison with the high-resolution result from the original retrographic sensor. (a) Rendering of the high-resolution \$20 bill example from the original retrographic sensor with a close-up view. (b) Rendering of the captured geometry using our method.

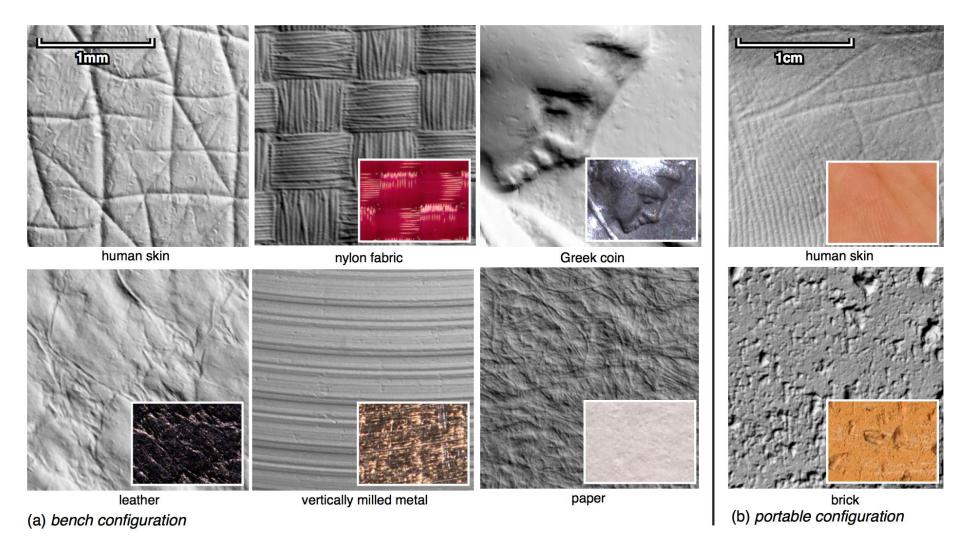
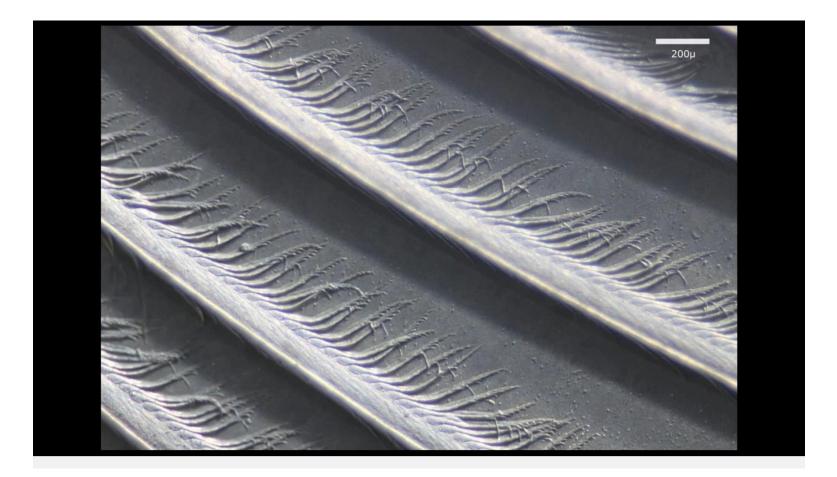


Figure 9: Example geometry measured with the bench and portable configurations. Outer image: rendering under direct lighting. Inset: macro photograph of original sample. Scale shown in upper left. Color images are shown for context and are to similar, but not exact scale.



Sensing Surfaces with GelSight

kimoatmit Subscribe 64 138,850 views https://www.youtube.com/watch?v=S7gXih4XS7A

Questions?