# CS4670: Computer Vision Noah Snavely

#### Lecture 31: Photometric stereo



#### What happens when a light ray hits an object?

Some of the light gets absorbed

• converted to other forms of energy (e.g., heat)

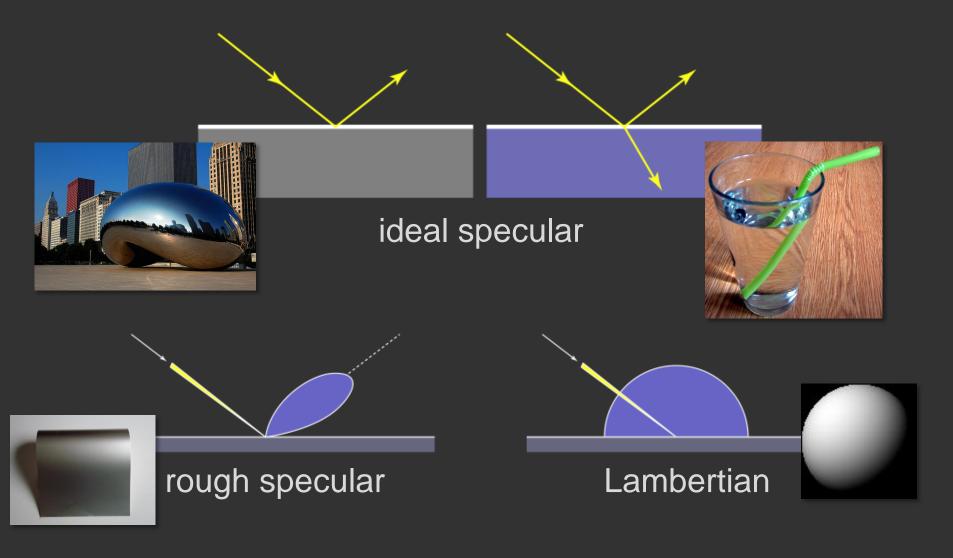
Some gets transmitted through the object

- possibly bent, through "refraction"
- a transmitted ray could possible bounce back

#### Some gets reflected

 as we saw before, it could be reflected in multiple directions (possibly all directions) at once

## **Classic reflection behavior**

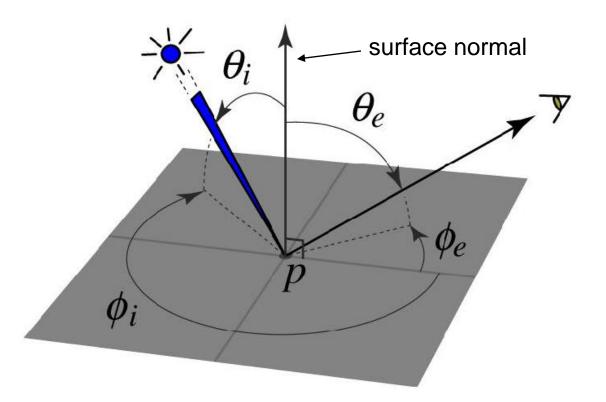


from Steve Marschner

# The BRDF

The Bidirectional Reflection Distribution Function

• Given an incoming ray  $(\theta_i, \phi_i)$  and outgoing ray  $(\theta_e, \phi_e)$  what proportion of the incoming light is reflected along outgoing ray?



Answer given by the BRDF:  $ho( heta_i,\phi_i, heta_e,\phi_e)$ 

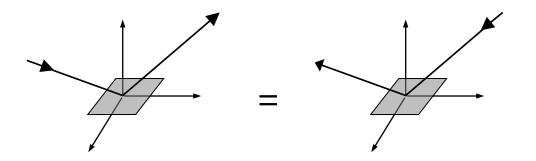
#### Constraints on the BRDF

Energy conservation

- Quantity of outgoing light ≤ quantity of incident light
  - − integral of BRDF  $\leq$  1

Helmholtz reciprocity

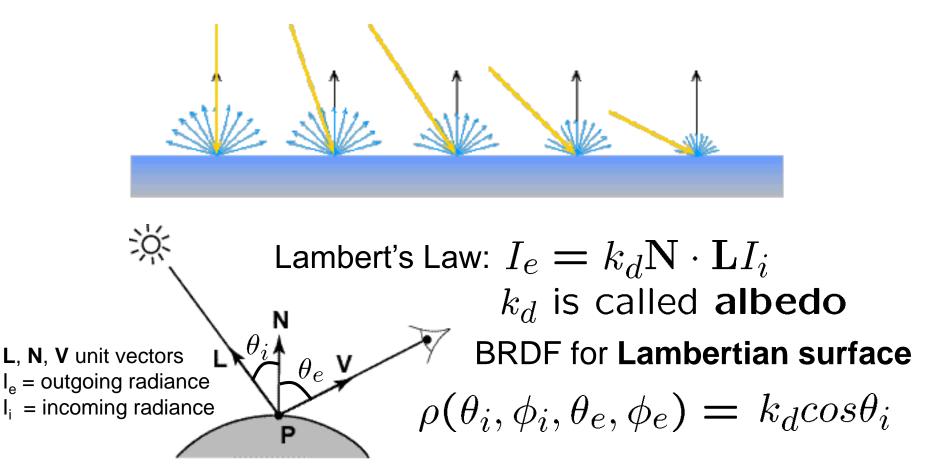
• reversing the path of light produces the same reflectance



## Diffuse reflection

Diffuse reflection governed by Lambert's law

- Viewed brightness does not depend on viewing direction
- Brightness does depend on direction of illumination
- This is the model most often used in computer vision

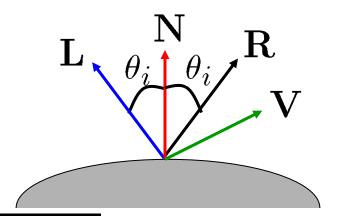




http://www.math.montana.edu/frankw/ccp/multiworld/twothree/lighting/applet1.htm http://www.math.montana.edu/frankw/ccp/multiworld/twothree/lighting/learn2.htm

## Specular reflection

For a perfect mirror, light is reflected about N

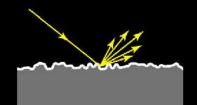


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$$I_e = \begin{cases} I_i & \text{if } \mathbf{V} = \mathbf{R} \\ 0 & \text{otherwise} \end{cases}$$



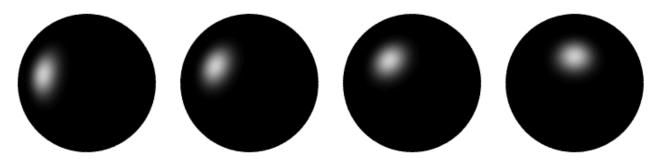
conductor plus microgeometry



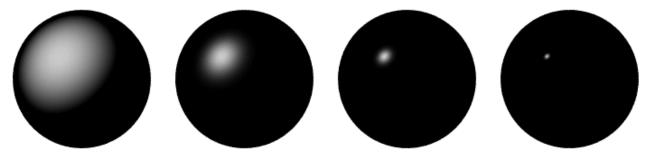
Near-perfect mirrors have a highlight around R

common model:  $I_e = k_s (\mathbf{V} \cdot \mathbf{R})^{n_s} I_i$  $\cos^{n_{s}}\phi$ 0.9 0.8 *n*<sub>s</sub> = 1 0.7 0.6 0.5 0.4 0.3 0.2 0.1  $n_{\rm S} = 128$ 0 100 80 60 40 20 0 20 40 60 80 100

#### Specular reflection

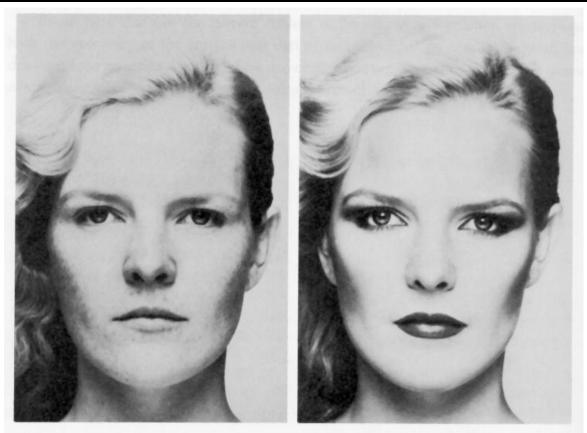


#### Moving the light source



Changing n<sub>s</sub>

#### **Photometric Stereo**

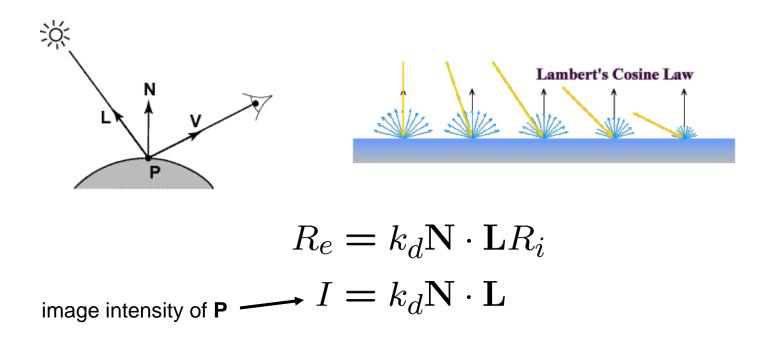


Merle Norman Cosmetics, Los Angeles

#### Readings

R. Woodham, *Photometric Method for Determining Surface Orientation from Multiple Images*. Optical Engineering 19(1)139-144 (1980). (<u>PDF</u>)

#### **Diffuse reflection**

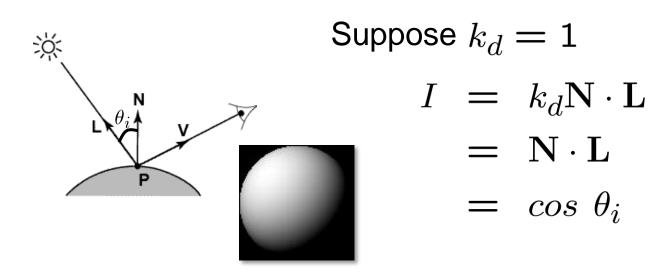


Simplifying assumptions

- $I = R_e$ : camera response function is the identity function:
- $R_i = 1$ : light source intensity is 1

- can achieve this by dividing each pixel in the image by R<sub>i</sub>

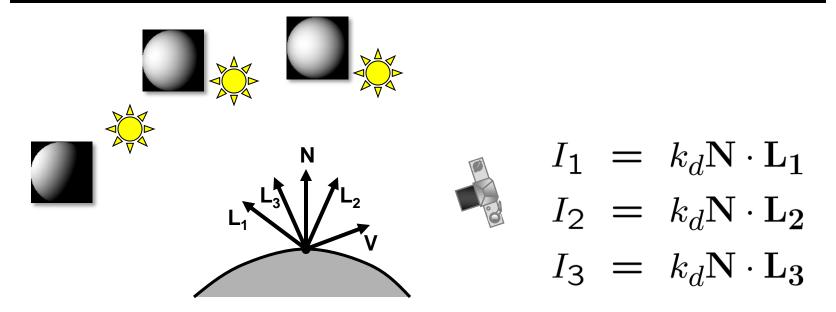
# Shape from shading



You can directly measure angle between normal and light source

- Not quite enough information to compute surface shape
- But can be if you add some additional info, for example
  - assume a few of the normals are known (e.g., along silhouette)
  - constraints on neighboring normals—"integrability"
  - smoothness
- Hard to get it to work well in practice
  - plus, how many real objects have constant albedo?

#### Photometric stereo



Can write this as a matrix equation:

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = k_d \begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix} \mathbf{N}$$

#### Solving the equations

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = \begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix} k_d \mathbf{N}$$
$$\underbrace{\mathbf{L}_3^T}_{\mathbf{L}_3^T} = \underbrace{\mathbf{L}_3^T}_{\mathbf{L}_3^T} \underbrace{\mathbf{L}_3^T} \underbrace{\mathbf{L}_3^T}_{\mathbf{L}_3^T} \underbrace{\mathbf{L}_3^T} \underbrace{\mathbf{L}_3^T}_{\mathbf{L}_3^T} \underbrace{\mathbf{L}_3^T} \underbrace{\mathbf{L}_3$$

- $\mathbf{G} = \mathbf{L}^{-1}\mathbf{I}$
- $k_d = \|\mathbf{G}\|$
- $\mathbf{N} = \frac{1}{k_d} \mathbf{G}$

## More than three lights

Get better results by using more lights

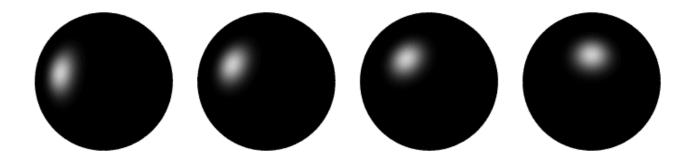
$$\begin{bmatrix} I_1 \\ \vdots \\ I_n \end{bmatrix} = \begin{bmatrix} \mathbf{L}_1 \\ \vdots \\ \mathbf{L}_n \end{bmatrix} k_d \mathbf{N}$$

Least squares solution:

$$\begin{split} \mathbf{I} &= \mathbf{L}\mathbf{G} \\ \mathbf{L}^T\mathbf{I} &= \mathbf{L}^T\mathbf{L}\mathbf{G} \\ \mathbf{G} &= (\mathbf{L}^T\mathbf{L})^{-1}(\mathbf{L}^T\mathbf{I}) \\ \end{split}$$
 Solve for N, k<sub>d</sub> as before What's the size of L<sup>T</sup>L?

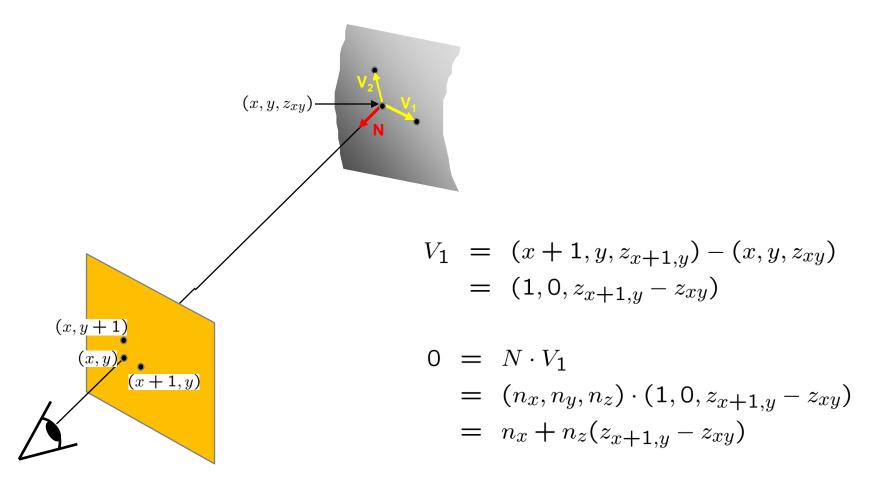
## Computing light source directions

Trick: place a chrome sphere in the scene



· the location of the highlight tells you where the light source is

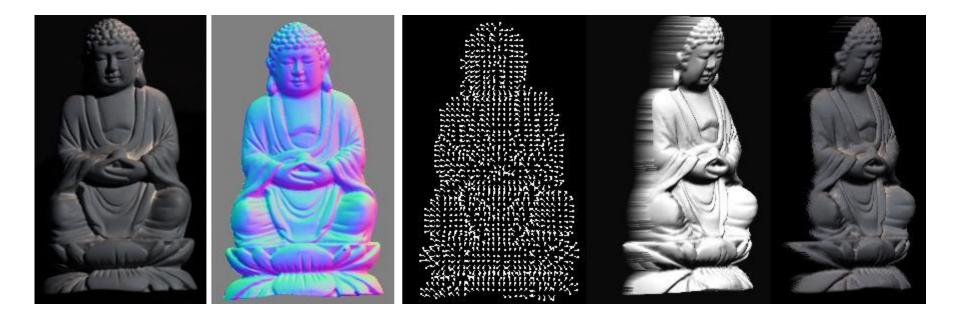
## Depth from normals



Get a similar equation for  $V_2$ 

- Each normal gives us two linear constraints on z
- compute z values by solving a matrix equation

# Example



## What if we don't have mirror ball?

Hayakawa, Journal of the Optical Society of America, 1994, <u>Photometric stereo under a light source with</u> <u>arbitrary motion</u>.

## Limitations

#### **Big problems**

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections

Smaller problems

- camera and lights have to be distant
- calibration requirements
  - measure light source directions, intensities
  - camera response function

Newer work addresses some of these issues

Some pointers for further reading:

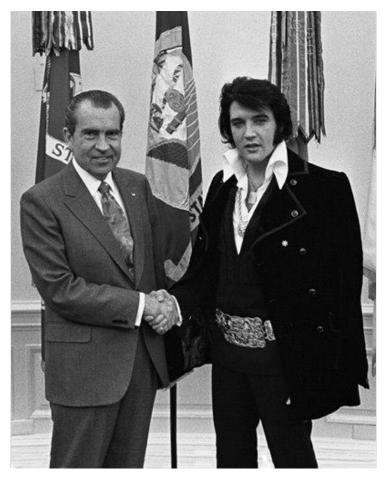
- Zickler, Belhumeur, and Kriegman, "<u>Helmholtz Stereopsis: Exploiting</u> <u>Reciprocity for Surface Reconstruction</u>." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann & Seitz, "<u>Example-Based Photometric Stereo: Shape</u> <u>Reconstruction with General, Varying BRDFs</u>." IEEE Trans. PAMI 2005

# Application: Detecting composite photos

#### Which is the real photo?



Fake photo



Real photo