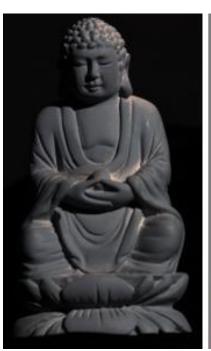
# **CS5670: Computer Vision**

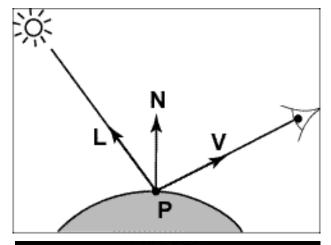
Photometric stereo







### Recap: Lambertian (Diffuse) Reflectance





$$I = k_d \mathbf{N} \cdot \mathbf{L}$$

• *I*: observed image intensity

•  $k_d$ : object albedo

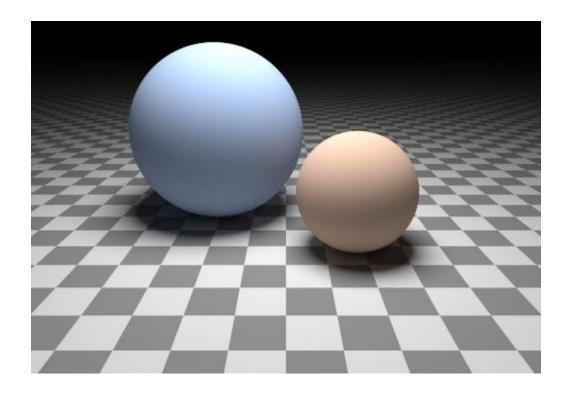
• N: surface normal

• L: light source direction



Lambertian sphere with constant albedo lit by a directional light source

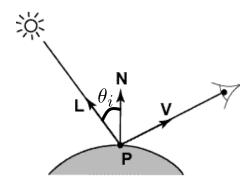
#### Can we determine shape from lighting?



- Are these spheres?
  - Or just flat discs painted with varying albedo?

# A Single Image: Shape from shading





Suppose (for now)  $k_d = 1$ 

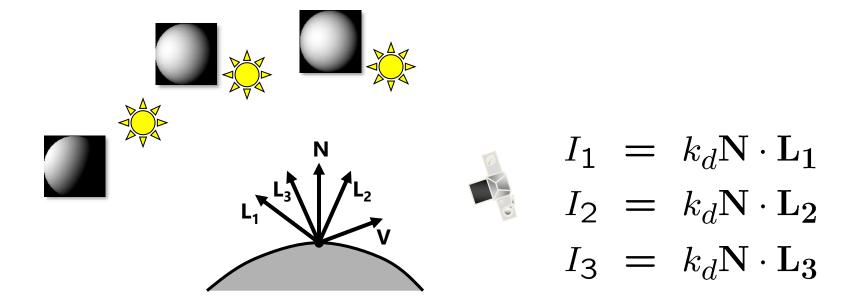
$$I = k_d \mathbf{N} \cdot \mathbf{L}$$
$$= \mathbf{N} \cdot \mathbf{L}$$
$$= \cos \theta_i$$

You can directly measure angle between normal and light source

- Not quite enough information to compute surface shape
- But can be if you add some additional info, for example
  - assume a few of the normals are known (e.g., along silhouette)
  - constraints on neighboring normals—"integrability"
  - smoothness
- Hard to get it to work well in practice
  - plus, how many real objects have constant albedo?
  - But, deep learning can help

## Let's take more than one photo!

#### Photometric stereo



Can write this as a matrix equation:

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = k_d \begin{vmatrix} \mathbf{L_1}^T \\ \mathbf{L_2}^T \\ \mathbf{L_3}^T \end{vmatrix} \mathbf{N}$$

# Solving the equations

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = \begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix} k_d \mathbf{N}$$

$$\mathbf{G} = \mathbf{L}^{-1} \mathbf{I}$$

$$k_d = \|\mathbf{G}\|$$

$$\mathbf{N} = \frac{1}{k_d} \mathbf{G}$$

$$\mathbf{N} = \frac{1}{k_d} \mathbf{G}$$

Solve one such linear system **per pixel** to solve for that pixel's surface normal

## More than three lights

Can get better results by using more than 3 lights

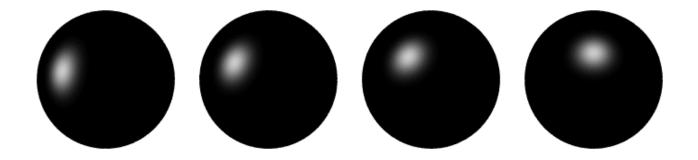
$$\begin{bmatrix} I_1 \\ \vdots \\ I_n \end{bmatrix} = \begin{bmatrix} \mathbf{L_1} \\ \vdots \\ \mathbf{L_n} \end{bmatrix} k_d \mathbf{N}$$

Least squares solution:

$$egin{array}{lll} \mathbf{I} &=& \mathbf{L}\mathbf{G} \\ \mathbf{L}^T\mathbf{I} &=& \mathbf{L}^T\mathbf{L}\mathbf{G} \\ \mathbf{G} &=& (\mathbf{L}^T\mathbf{L})^{-1}(\mathbf{L}^T\mathbf{I}) \end{array}$$
 Solve for N,  $\mathsf{k}_\mathsf{d}$  as before

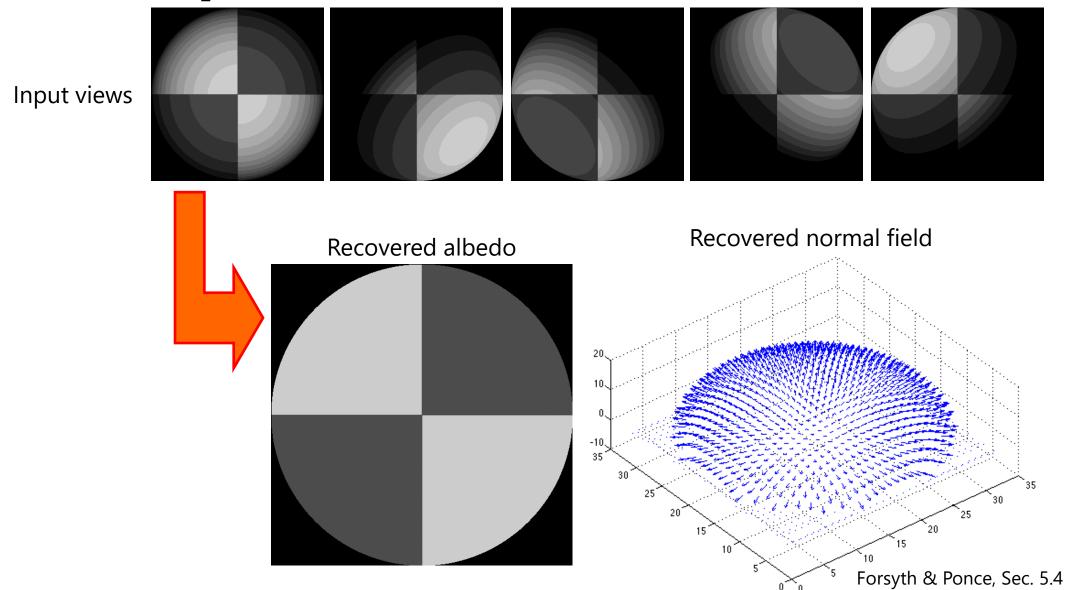
## Computing light source directions

Trick: place a chrome sphere in the scene



- the location of the highlight tells you where the light source is

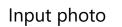
# Example

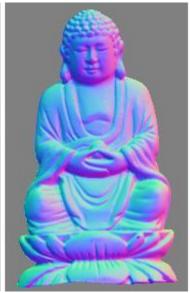


## **Depth from normals**

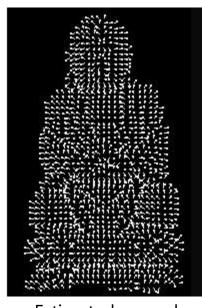
- Solving the linear system perpixel gives us an estimated surface normal for each pixel
- How can we compute depth from normals?
  - Normals are like the "derivative" of the true depth







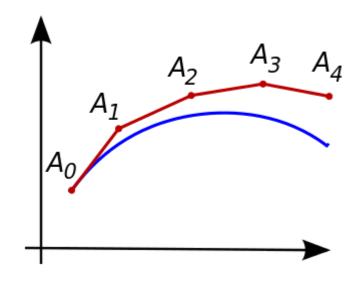
Estimated normals



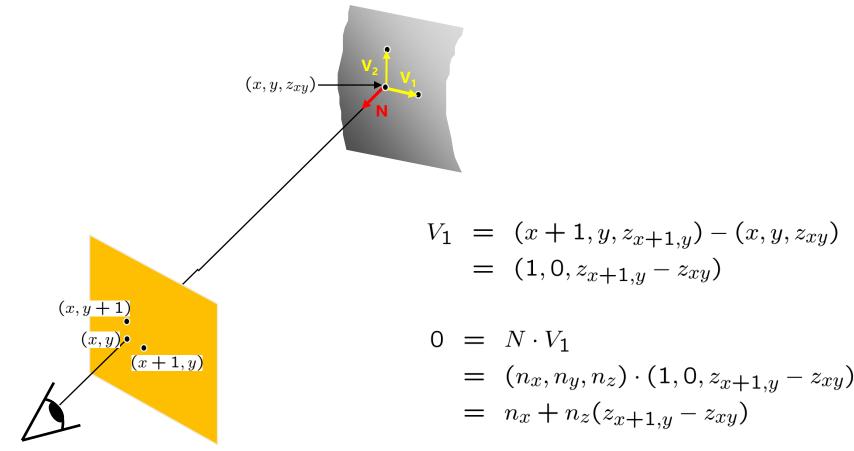
Estimated normals (needle diagram)

## **Normal Integration**

- Integrating a set of derivatives is easy in 1D
  - (similar to Euler's method from diff. eq. class)
- Could integrate normals in each column / row separately
  - Wouldn't give a good surface
- Instead, we formulate as a linear system and solve for depths that best agree with the surface normals



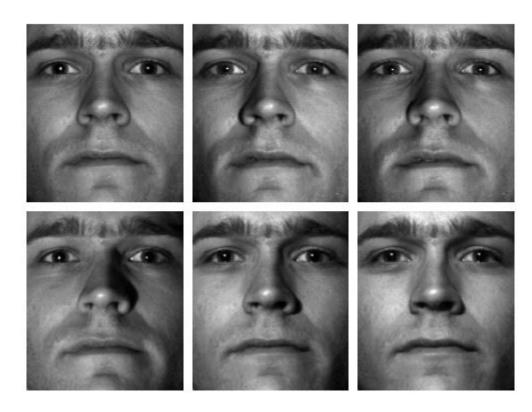
## **Depth from normals**



#### Get a similar equation for V<sub>2</sub>

- Each normal gives us two linear constraints on z
- compute z values by solving a matrix equation

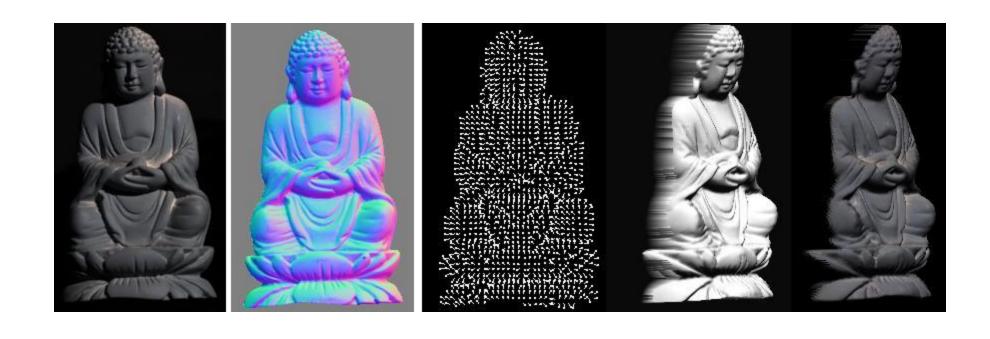
### Results





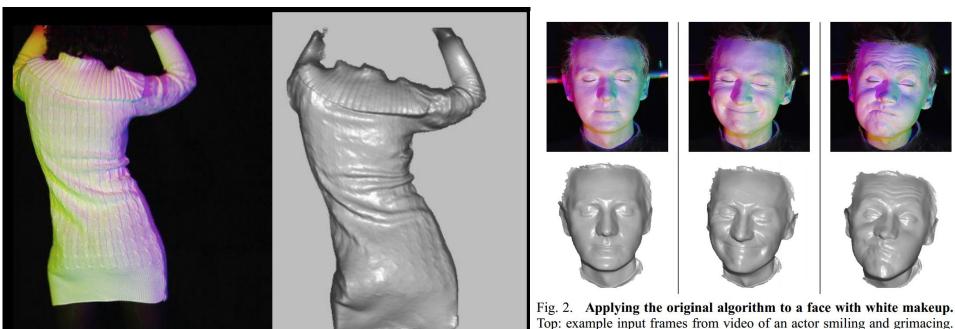


### Results



#### **Extension**

Photometric Stereo from Colored Lighting



Top: example input frames from video of an actor smiling and grimacing. Bottom: the resulting integrated surfaces.

#### **Video Normals from Colored Lights**

Gabriel J. Brostow, Carlos Hernández, George Vogiatzis, Björn Stenger, Roberto Cipolla <u>IEEE TPAMI</u>, Vol. 33, No. 10, pages 2104-2114, October 2011.

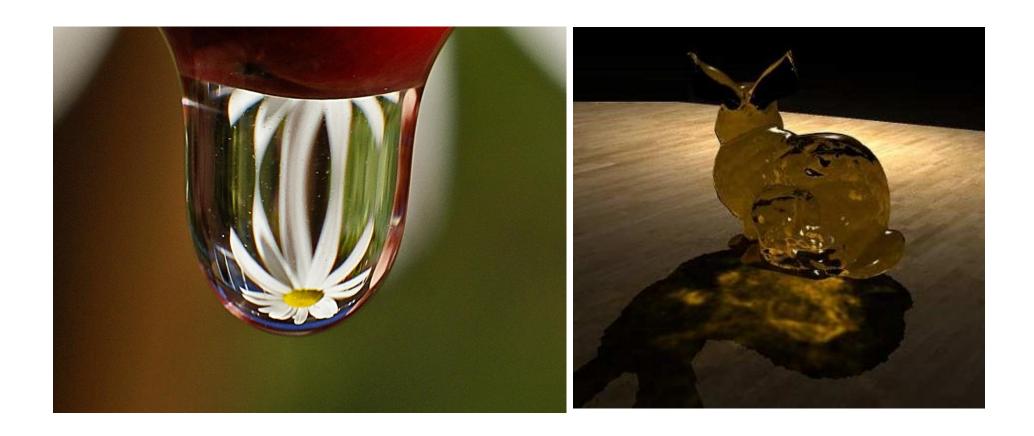
## **Questions?**

# For now, ignore specular reflection





#### **And Refraction...**



#### And Interreflections...





Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

# And Subsurface Scattering...





#### Limitations

#### Bigger problems

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections

#### Smaller problems

- camera and lights have to be distant
- calibration requirements
  - measure light source directions, intensities
  - camera response function

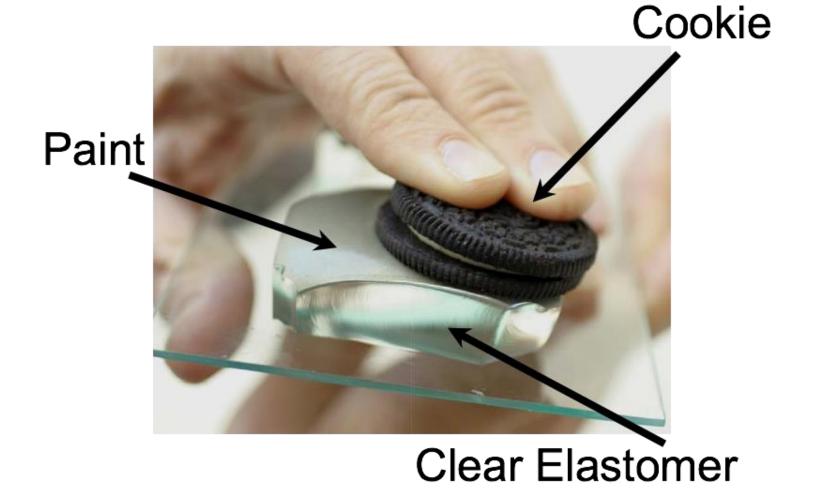
#### Newer work addresses some of these issues

#### Some pointers for further reading:

- Zickler, Belhumeur, and Kriegman, "<u>Helmholtz Stereopsis: Exploiting Reciprocity for Surface Reconstruction</u>." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann & Seitz, "<u>Example-Based Photometric Stereo: Shape Reconstruction with General, Varying BRDFs.</u>" IEEE Trans. PAMI 2005



HOME PRODUCTS VIDEOS IMAGES PAPERS NEWS ABOUT US CONTACT

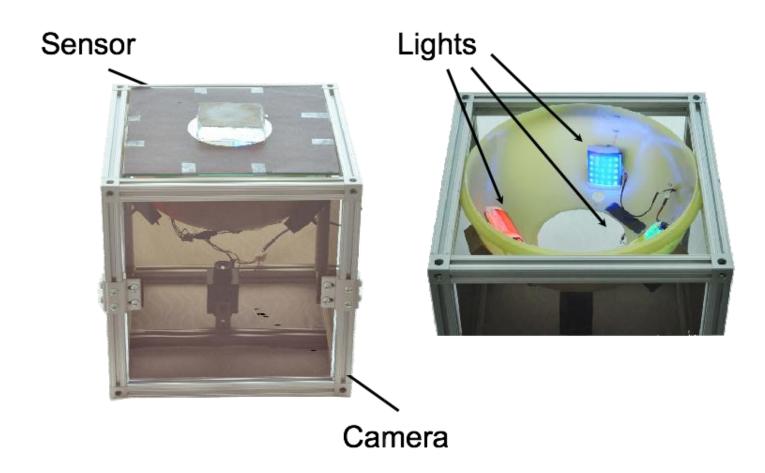


Johnson and Adelson, 2009

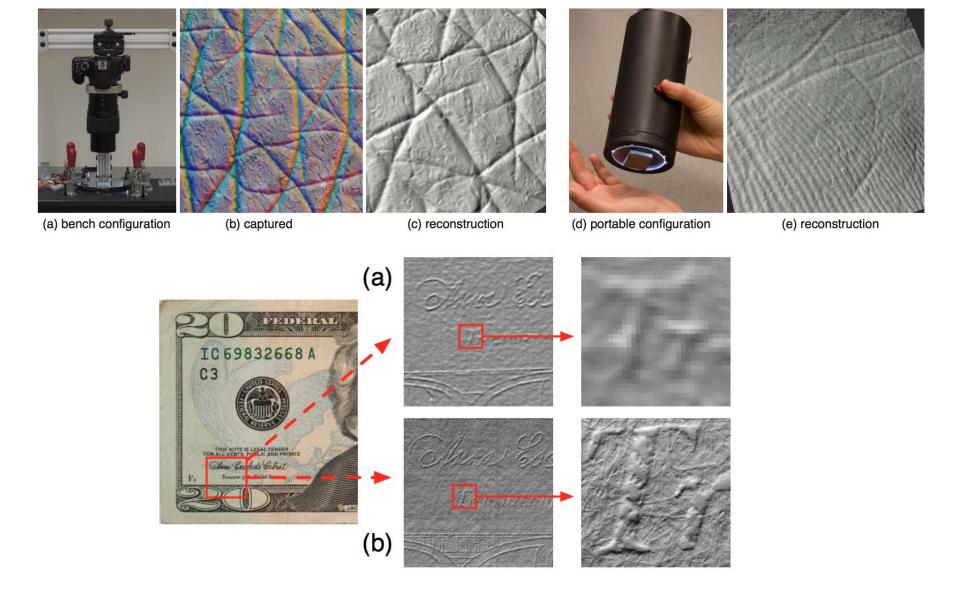




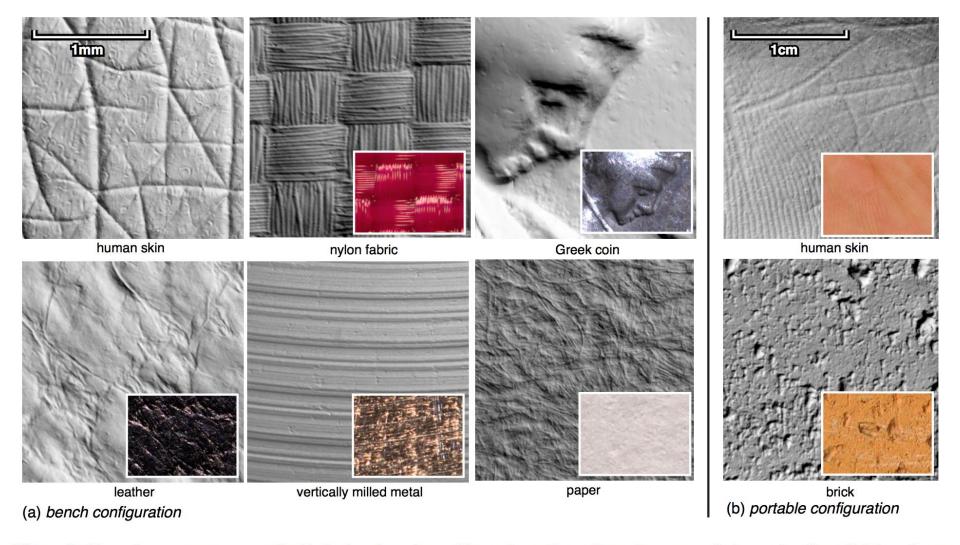
### Lights, camera, action



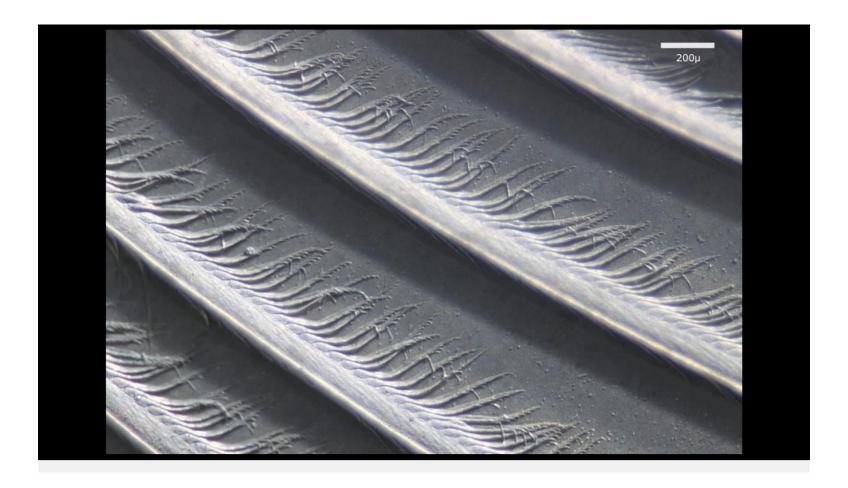




**Figure 7:** Comparison with the high-resolution result from the original retrographic sensor. (a) Rendering of the high-resolution \$20 bill example from the original retrographic sensor with a close-up view. (b) Rendering of the captured geometry using our method.



**Figure 9:** Example geometry measured with the bench and portable configurations. Outer image: rendering under direct lighting. Inset: macro photograph of original sample. Scale shown in upper left. Color images are shown for context and are to similar, but not exact scale.



#### Sensing Surfaces with GelSight



138,850 views

#### **InverseRenderNet:** Learning single image inverse rendering

#### Ye Yu and William A. P. Smith Department of Computer Science, University of York, UK

{yy1571, william.smith}@york.ac.uk

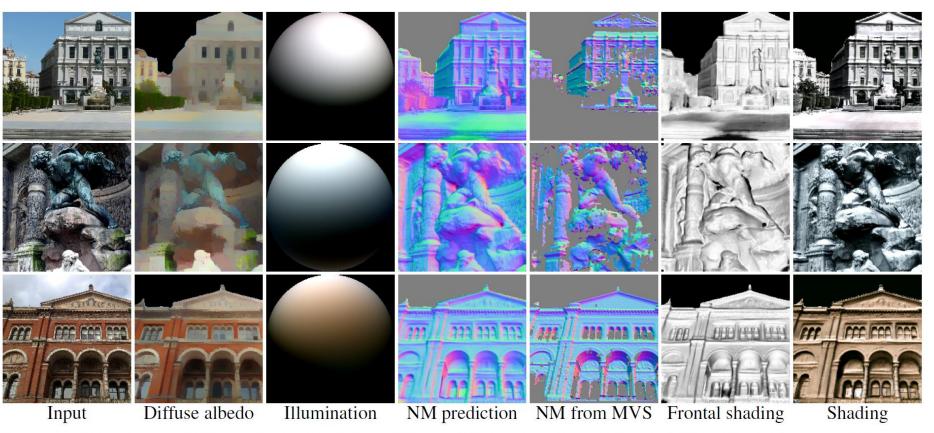


Figure 1: From a single image (col. 1), we estimate albedo and normal maps and illumination (col. 2-4); comparison multiview stereo result from several hundred images (col. 5); re-rendering of our shape with frontal/estimated lighting (col. 6-7).

## **Questions?**