

CS4670/5760: Computer Vision

Kavita Bala

Lecture 29: Photometric Stereo 2

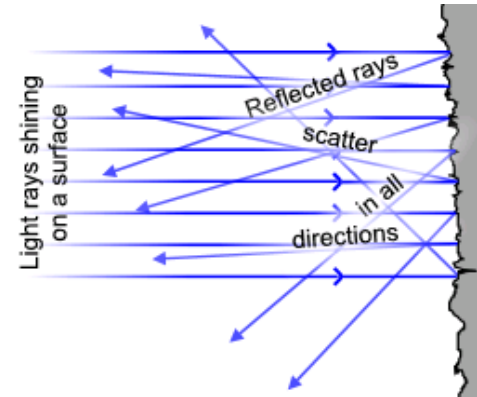
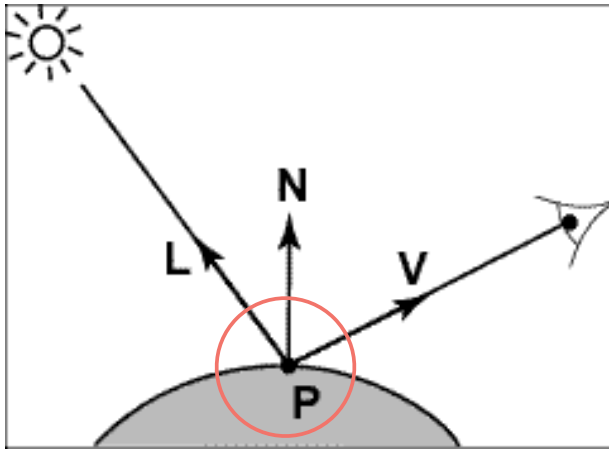


Thanks to Scott Wehrwein

Announcements

- PA 4 out tonight
- HW 2 out tonight
- Wed/Fri: MVS, sFM

Lambertian Reflectance



$$I = N \cdot L$$

Image intensity $=$ Surface normal \cdot Light direction

Image intensity \propto $\cos(\text{angle between } N \text{ and } L)$

Solving the Equations

$$\underbrace{\begin{bmatrix} I_1 & I_2 & I_3 \end{bmatrix}}_{\substack{\mathbf{I} \\ 1 \times 3}} = k_d \underbrace{\mathbf{N}^T}_{\substack{\mathbf{G} \\ 1 \times 3}} \underbrace{\begin{bmatrix} \mathbf{L}_1 & \mathbf{L}_2 & \mathbf{L}_3 \end{bmatrix}}_{\substack{\mathcal{L} \\ 3 \times 3}}$$

$$\mathbf{G} = \mathbf{I}\mathcal{L}^{-1}$$

- When is \mathcal{L} nonsingular (invertible)?
 - ≥ 3 light directions are linearly independent, or:
 - All light direction vectors cannot lie in a plane.
- What if we have more than one pixel?
 - Stack them all into one big system.

More than Three Lights

$$\begin{bmatrix} I_1 & \dots & I_n \end{bmatrix} = k_d \mathbf{N}^T \begin{bmatrix} \mathbf{L}_1 & \dots & \mathbf{L}_n \end{bmatrix}$$

- Solve using least squares (normal equations):

$$\mathbf{I} = \mathbf{G}\mathbf{L}$$

$$\mathbf{I}\mathbf{L}^T = \mathbf{G}\mathbf{L}\mathbf{L}^T$$

$$\mathbf{G} = (\mathbf{I}\mathbf{L}^T)(\mathbf{L}\mathbf{L}^T)^{-1}$$

- Equivalently use SVD
- Given \mathbf{G} , solve for \mathbf{N} and k_d as before.

More than one pixel

Stack all pixels into one system:

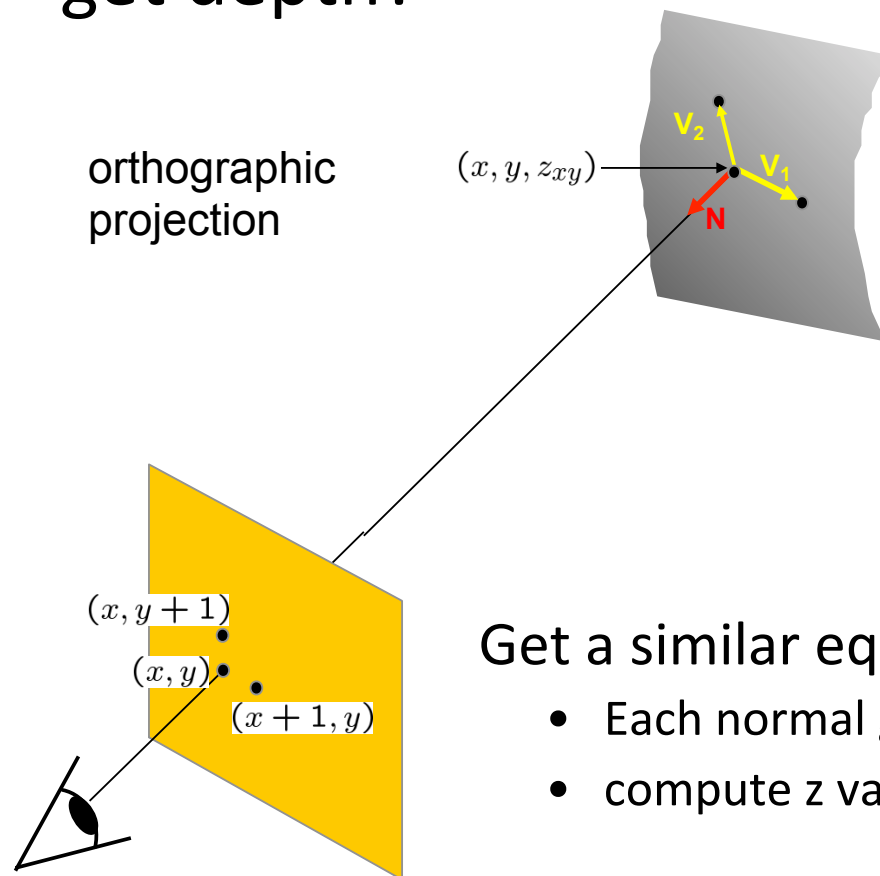
$$\begin{array}{c} p \times \# \text{ images} \\ \boxed{I} \end{array} = \begin{array}{c} p \times 3 \\ \boxed{N} \end{array} * \begin{array}{c} 3 \times \# \text{ images} \\ \boxed{L} \end{array}$$

Solve as before.

Depth Map from Normal Map

- We now have a surface normal, but how do we get depth?

orthographic projection



Assume a smooth surface

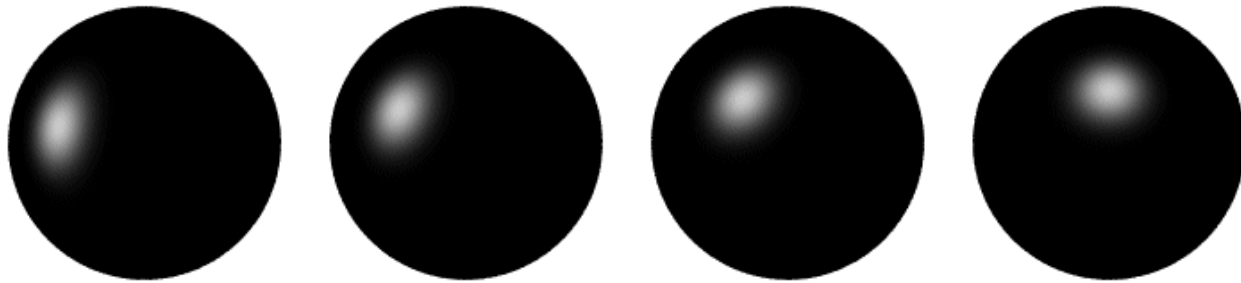
$$\begin{aligned}V_1 &= (x + 1, y, z_{x+1,y}) - (x, y, z_{xy}) \\ &= (1, 0, z_{x+1,y} - z_{xy})\end{aligned}$$
$$\begin{aligned}0 &= N \cdot V_1 \\ &= (n_x, n_y, n_z) \cdot (1, 0, z_{x+1,y} - z_{xy}) \\ &= n_x + n_z(z_{x+1,y} - z_{xy})\end{aligned}$$

Get a similar equation for V_2

- Each normal gives us two linear constraints on z
- compute z values by solving a matrix equation

Determining Light Directions

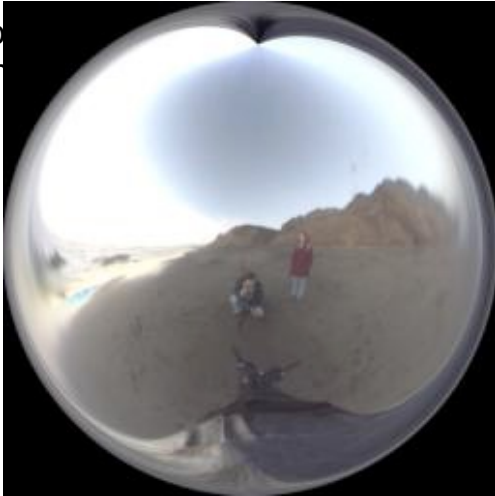
- Trick: Place a mirror ball in the scene.



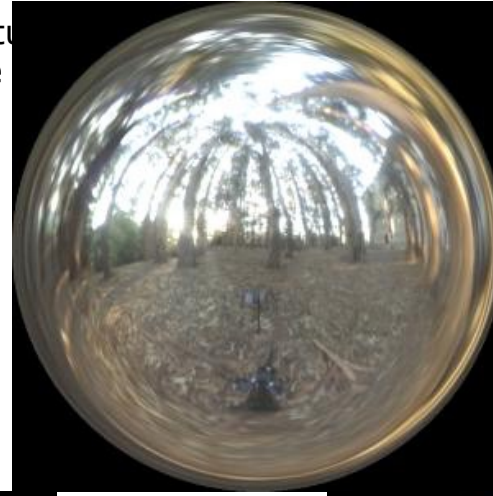
- The location of the highlight is determined by the light source direction.

Real-World HDR Lighting Environments

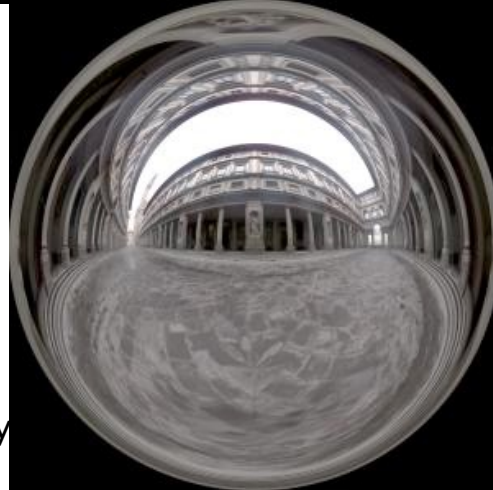
Funston
Beach



Eucalyptus
Grove



Uffizi
Gallery



Grace
Cathedral



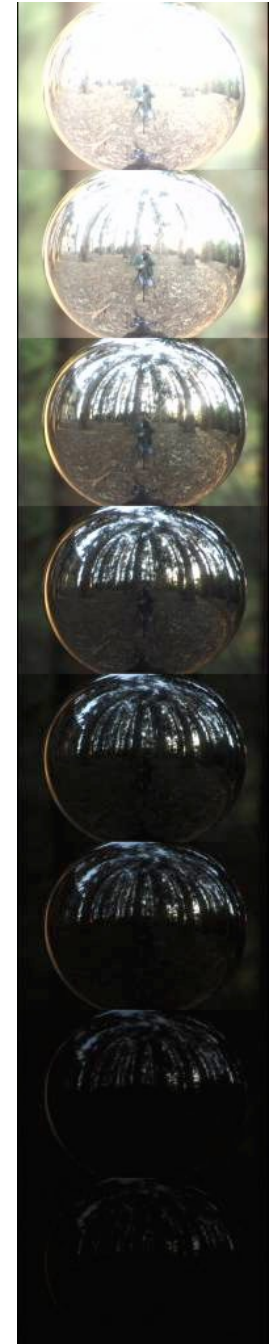
Lighting Environments from the Light Probe Image Gallery:
<http://www.debevec.org/Probes/>

Mirrored Sphere

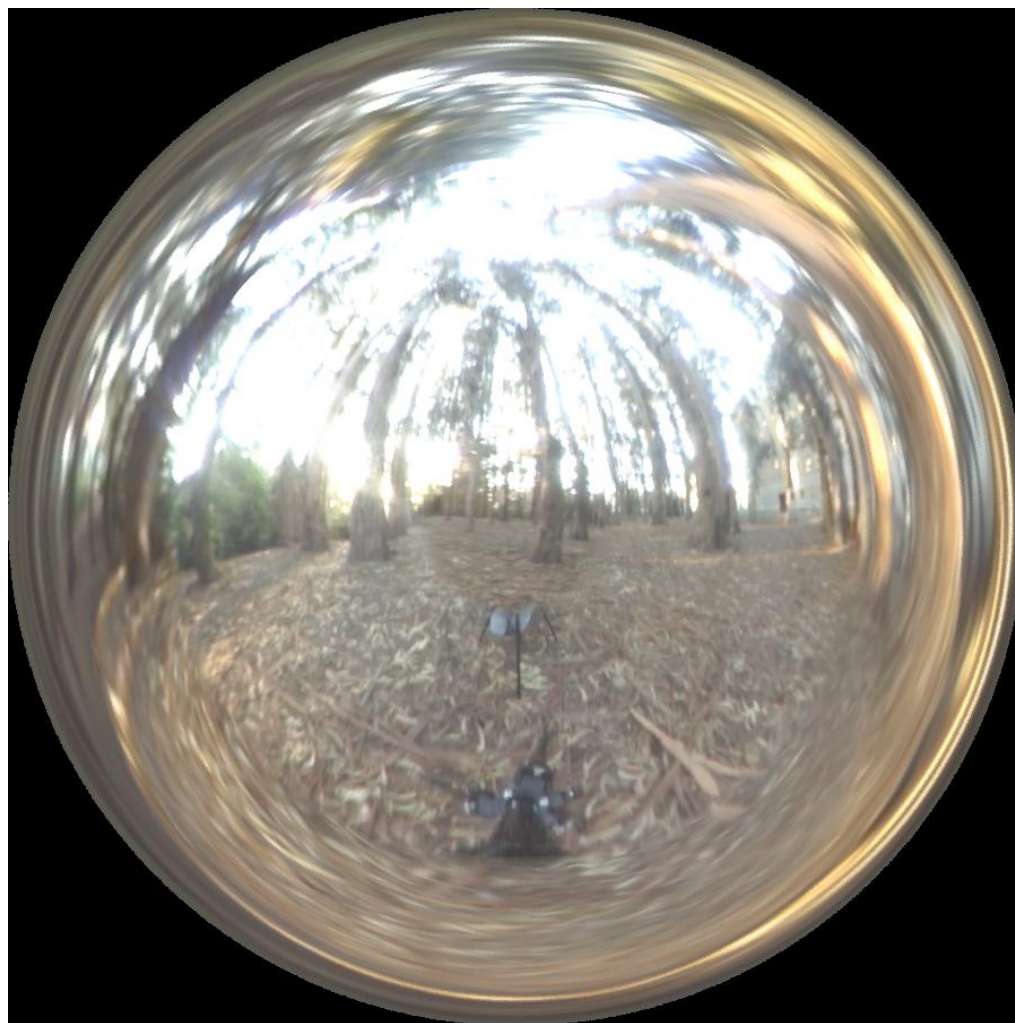




Acquiring the Light Probe



Assembling the Light Probe



Extreme HDR Image Series



1 sec
f/4



1/4 sec
f/4



1/30 sec
f/4



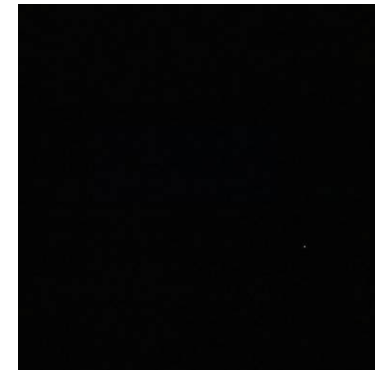
1/30 sec
f/16



1/250 sec
f/16



1/1000 sec
f/16



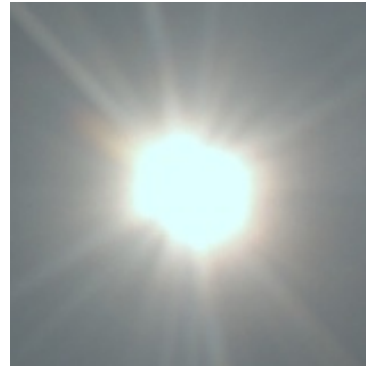
1/8000 sec f/16

Extreme HDR Image Series

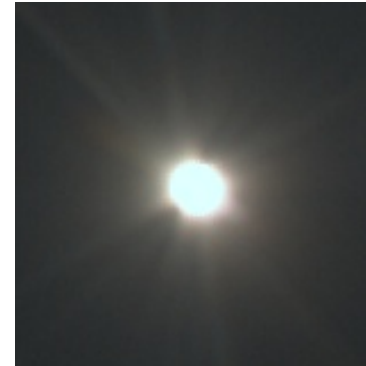
sun closeup



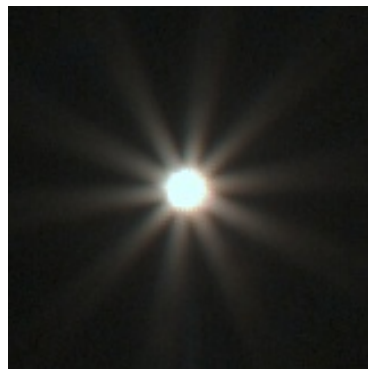
1 sec
f/4



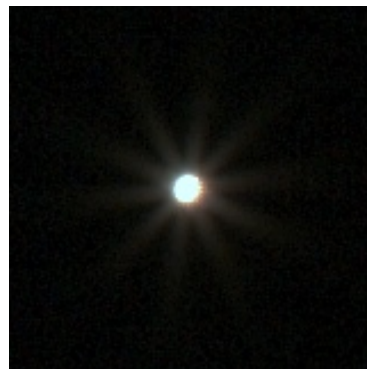
1/4 sec
f/4



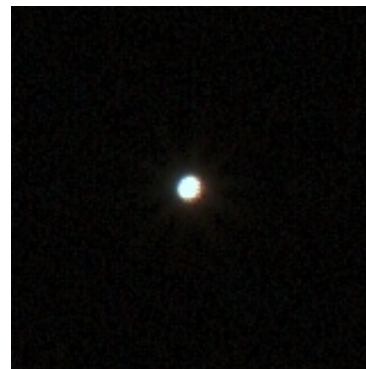
1/30 sec
f/4



1/30 sec
f/16



1/250 sec
f/16

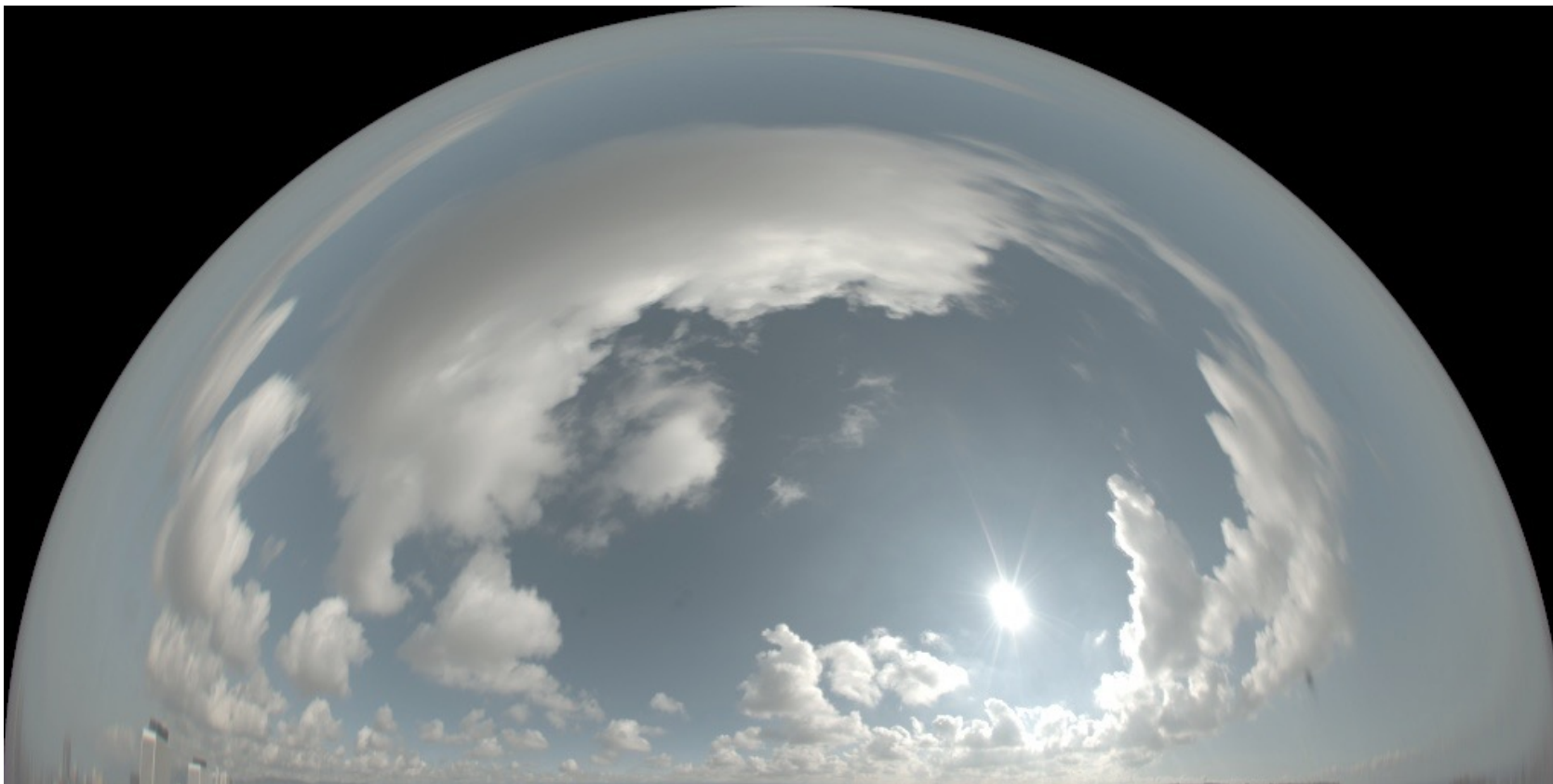


1/1000 sec
f/16



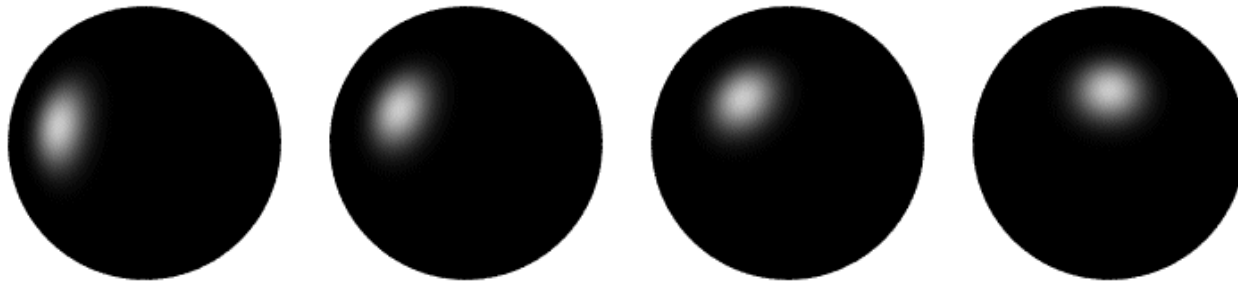
1/8000 sec f/16
only image that does not saturate!

HDRI Sky Probe



Determining Light Directions

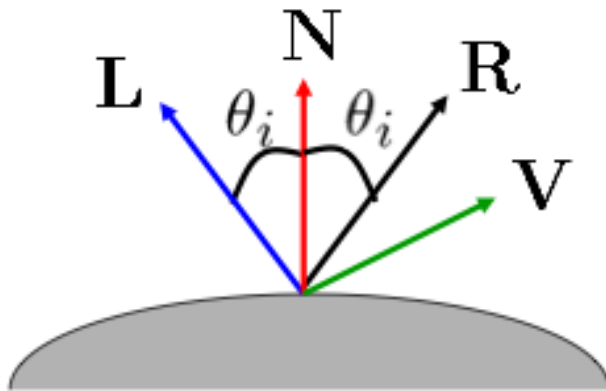
- Trick: Place a mirror ball in the scene.



- The location of the highlight is determined by the light source direction.

Determining Light Directions

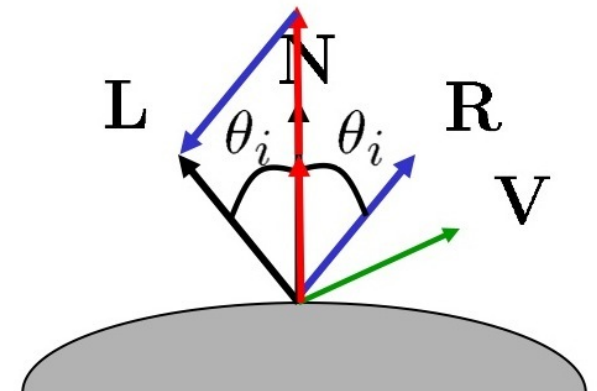
- For a perfect mirror, the light is reflected across N :



$$I_e = \begin{cases} I_i & \text{if } \mathbf{V} = \mathbf{R} \\ 0 & \text{otherwise} \end{cases}$$

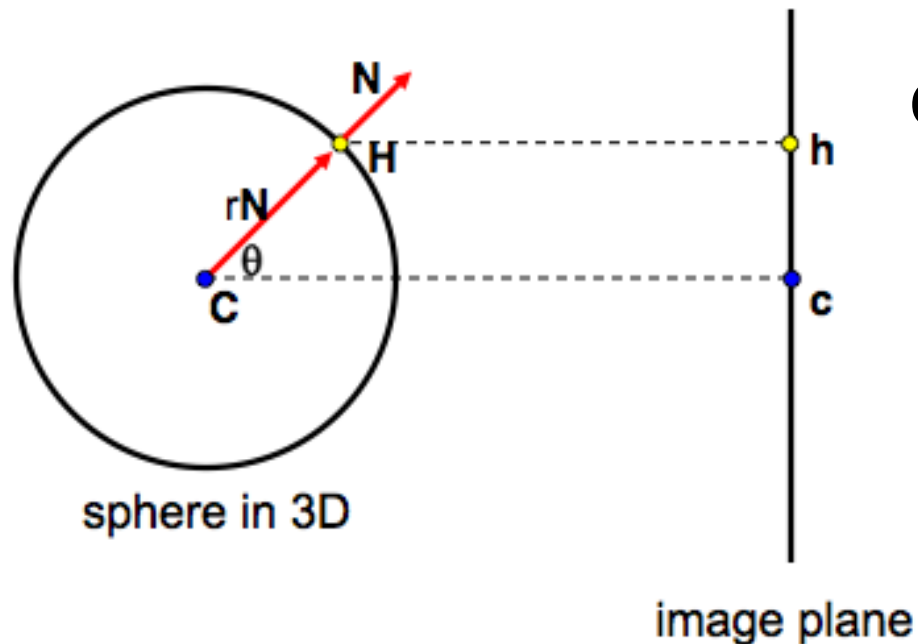
- So the light source direction is given by:

$$L = 2(N \cdot R)N - R$$



Determining Light Directions

- For a sphere with highlight at point H:



Compute N:

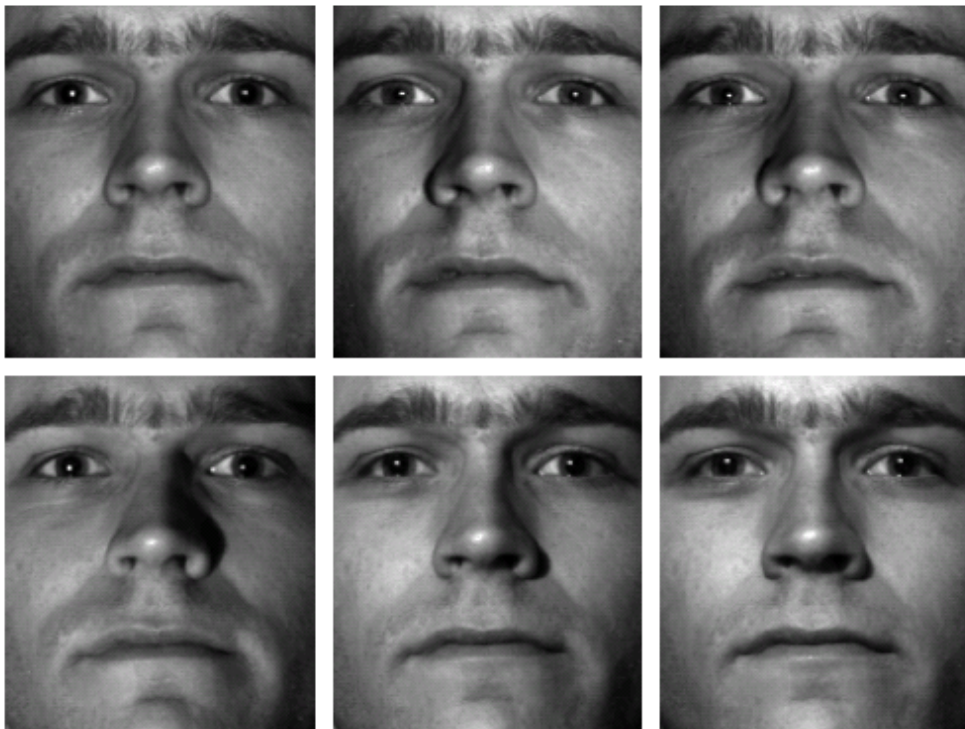
$$N_x = \frac{x_h - x_c}{r}$$

$$N_y = \frac{y_h - y_c}{r}$$

$$N_z = \sqrt{1 - x^2 - y^2}$$

- $R =$ direction of the camera from $C = [0 \ 0 \ 1]^T$
 $L = 2(N \cdot R)N - R$

Results

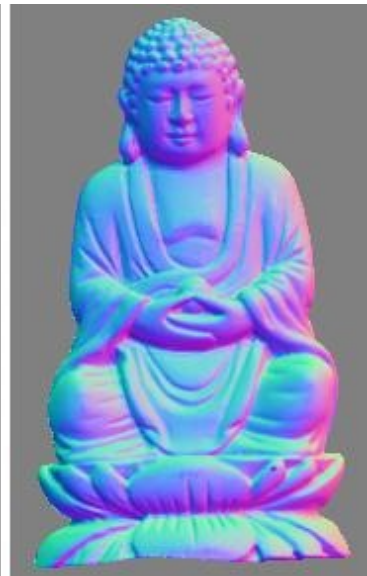


from Athos Georghiades

Results



Input
(1 of 12)



Normals (RGB
colormap)



Normals (vectors)



Shaded 3D
rendering



Textured 3D
rendering

Color Images

- Now we have 3 equations for a pixel:

$$I_R = k_{dR} \mathbf{L} \mathbf{N}$$

$$I_G = k_{dG} \mathbf{L} \mathbf{N}$$

$$I_B = k_{dB} \mathbf{L} \mathbf{N}$$

- Simple approach: solve for \mathbf{N} using grayscale or a single channel
- Then fix \mathbf{N} and solve for each channel's k_d

$$k_d = \frac{\sum_i I_i L_i N^T}{\sum_i (L_i N^T)^2}$$

Color Images

- Then fix N and solve for each channel' s : k_d

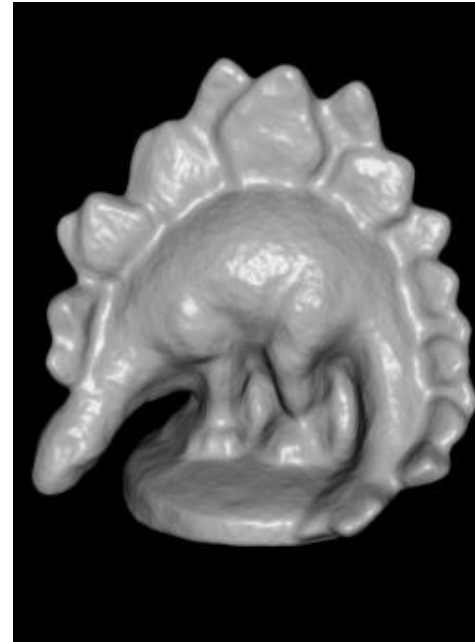
$$Q = \sum_i (I_i - k_d L_i N^T)^2$$

$$\frac{\delta Q}{\delta k_d} = \sum_i -2(I_i - k_d L_i N^T) L_i N^T = 0$$

$$k_d = \frac{\sum_i I_i L_i N^T}{\sum_i (L_i N^T)^2}$$

For (unfair) Comparison

- Multi-view stereo results on a similar object
- 47+ hrs compute time



State-of-the-art
MVS result



Ground truth

Taking Stock: Assumptions

| Lighting | Materials | Geometry | Camera |
|-----------------------------|-----------------------------|------------------------|--------------|
| directional | diffuse | convex / no shadows | linear |
| known direction | no inter- reflections | | orthographic |
| > 2 nonplanar directions | no subsurface scattering | | |

Questions?

Unknown Lighting

- What we've seen so far: [Woodham 1980]
- Next up: Unknown light directions [Hayakawa 1994]

Unknown Lighting

Surface normals Light directions

$$I = kN \cdot \ell L$$

Diffuse
albedo

Light
intensity

Unknown Lighting

Surface normals, scaled
by albedo

Light directions, scaled
by intensity

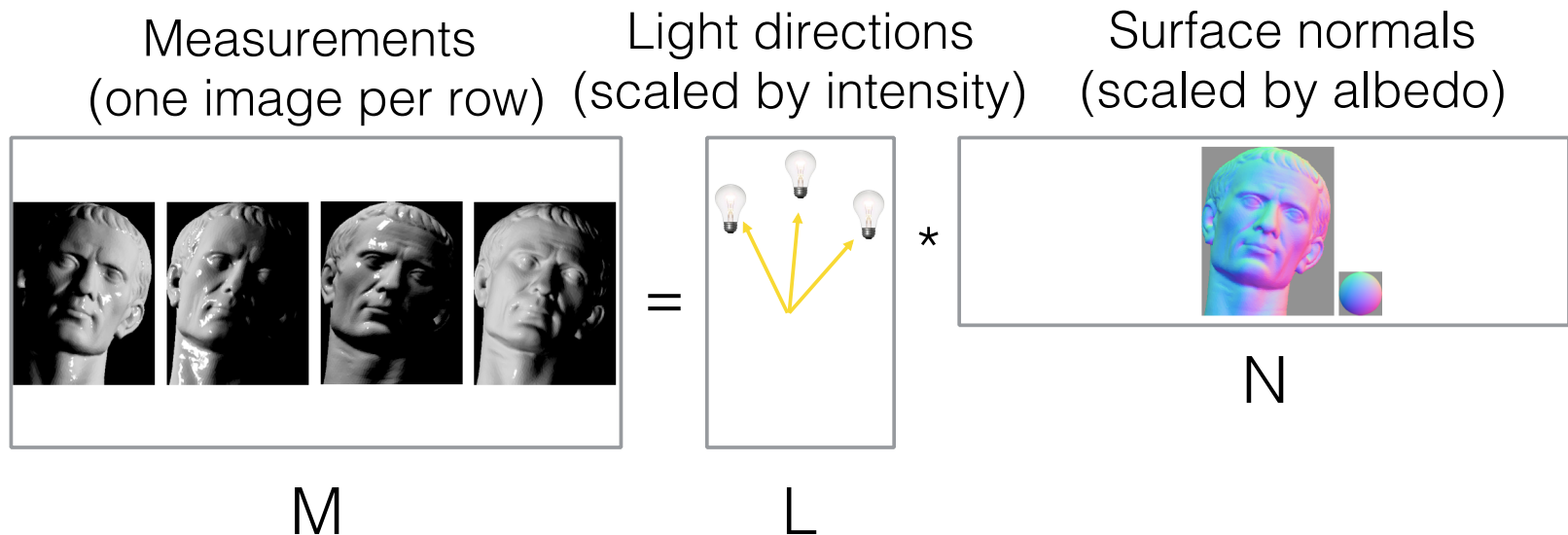
$$I = N \cdot L$$


Unknown Lighting

Same as before, just transposed:

$$\begin{array}{c} n = \# \text{ images} \\ \boxed{M} \end{array} = \begin{array}{c} p = \# \text{ pixels} \\ \boxed{L} \end{array} * \boxed{N}$$

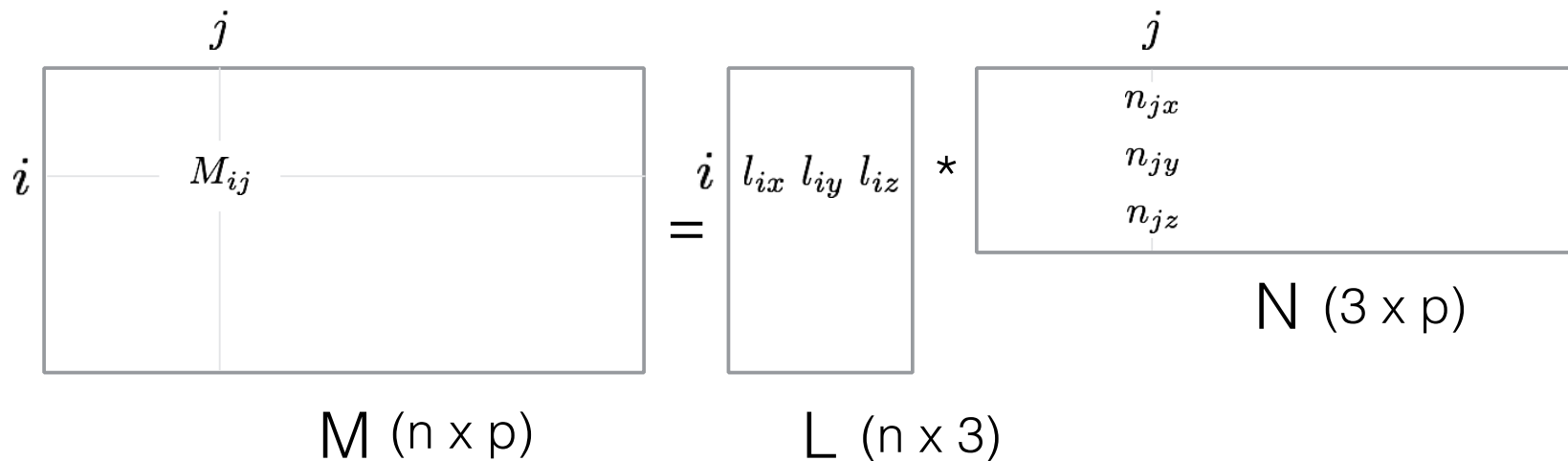
Unknown Lighting



Both L and N are now unknown!
This is a matrix factorization problem.

Unknown Lighting

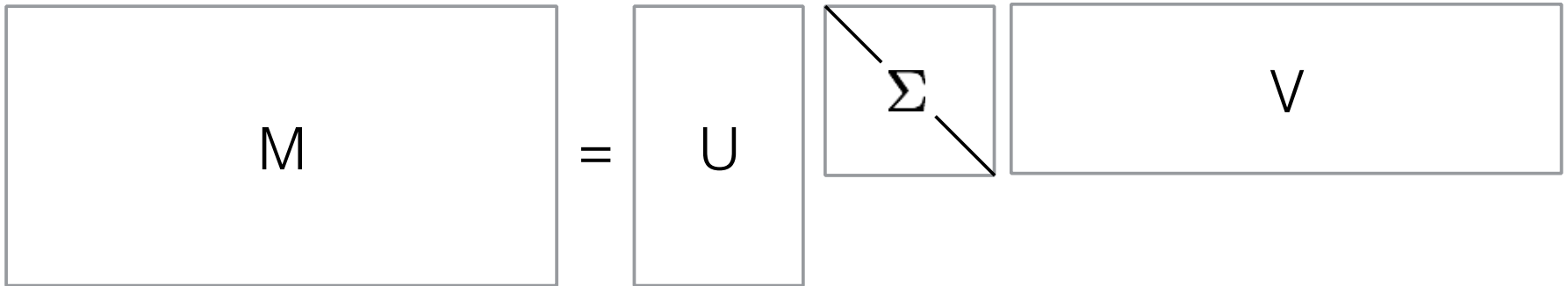
$$M_{ij} = L_i \cdot N_j$$



There's hope: We know that M is rank 3

Unknown Lighting

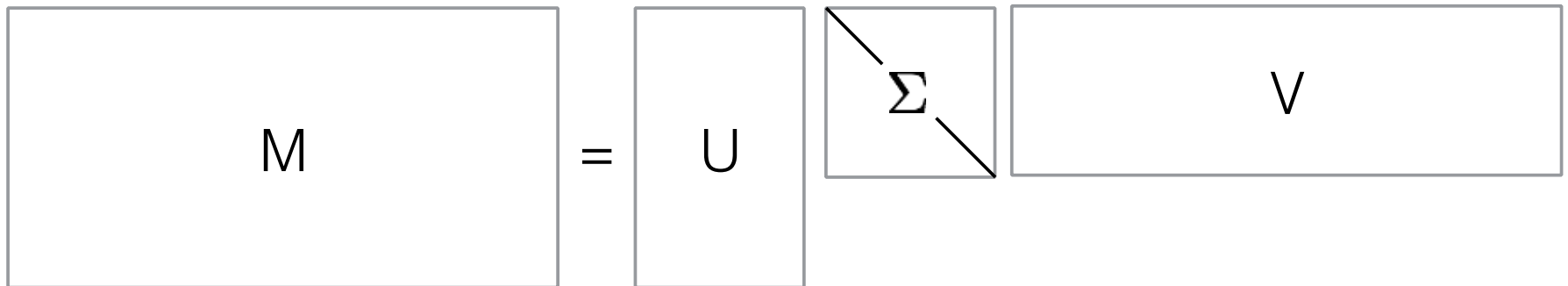
Use the SVD to decompose M:

$$\boxed{M} = \boxed{U} \boxed{\Sigma} \boxed{V}$$


SVD gives the best rank-3 approximation of a matrix.

Unknown Lighting

Use the SVD to decompose M:

$$\boxed{M} = \boxed{U} \boxed{\Sigma} \boxed{V}$$


What do we do with Σ ?

Unknown Lighting

Use the SVD to decompose M:

$$M = U\sqrt{\Sigma}V$$

What do we do with Σ ?

Unknown Lighting

Use the SVD to decompose M:

$$\boxed{M} = \boxed{U\sqrt{\Sigma}} \boxed{\sqrt{\Sigma}V}$$

Can we just do that?

Unknown Lighting

Use the SVD to decompose M:

$$M = U\sqrt{\Sigma} \begin{matrix} A \\ A^{-1} \\ \sqrt{\Sigma}V \end{matrix}$$

$$\hat{L} = U\sqrt{\Sigma}, \hat{S} = \sqrt{\Sigma}V$$

Can we just do that? ...almost.

The decomposition is non-unique up to an invertible 3x3 A.

Unknown Lighting

Use the SVD to decompose M:

$$M = U\sqrt{\Sigma} A A^{-1} \sqrt{\Sigma} V$$

$$L = U\sqrt{\Sigma} A \quad S = A^{-1}\sqrt{\Sigma} V$$

Unknown Lighting

Use the SVD to decompose M:

$$M = U\sqrt{\Sigma} A A^{-1} \sqrt{\Sigma} V$$

$$L = U\sqrt{\Sigma} A \quad S = A^{-1}\sqrt{\Sigma} V$$

You can find A if you know

- 6 points with the same reflectance, or
- 6 lights with the same intensity.

Why 6 points?

- Let $C = A^{-1}$, to match the notation of paper

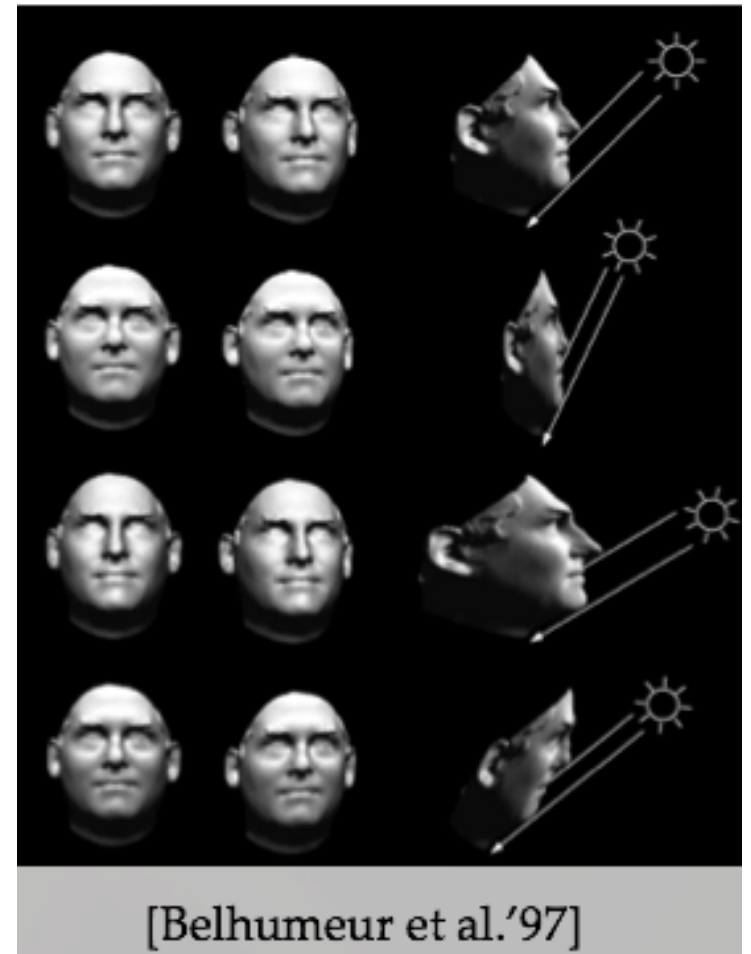
$$\hat{S} = \sqrt{\Sigma}V \quad \begin{aligned} \hat{s}_k^T C C^T \hat{s}_k &= \text{constant} \\ \hat{s}_k^T B \hat{s}_k &= \text{constant} \end{aligned}$$

- B is symmetric, hence 6

$$B = W\Pi W^T, C = W\sqrt{\Pi}$$

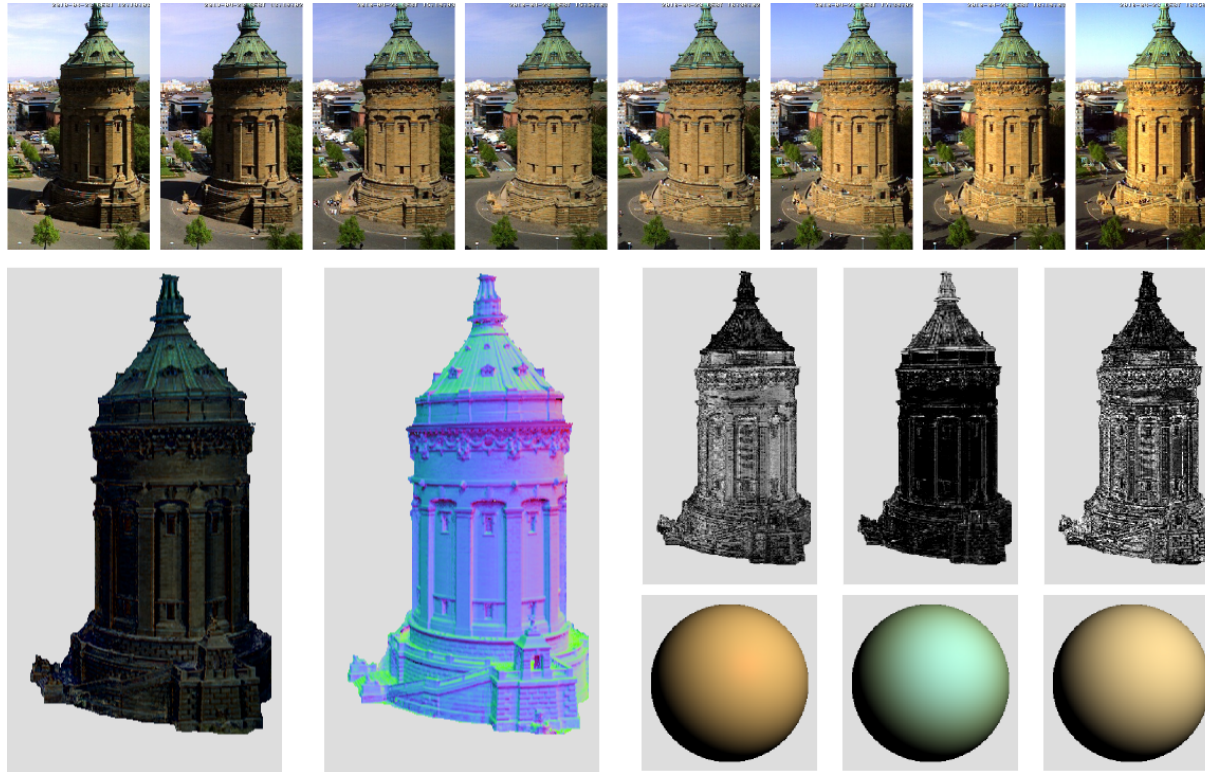
Unknown Lighting: Ambiguities

- Multiple combinations of lighting and geometry can produce the same sets of images.
- Add assumptions or prior knowledge about geometry or lighting, etc. to limit the ambiguity.



Since 1994...

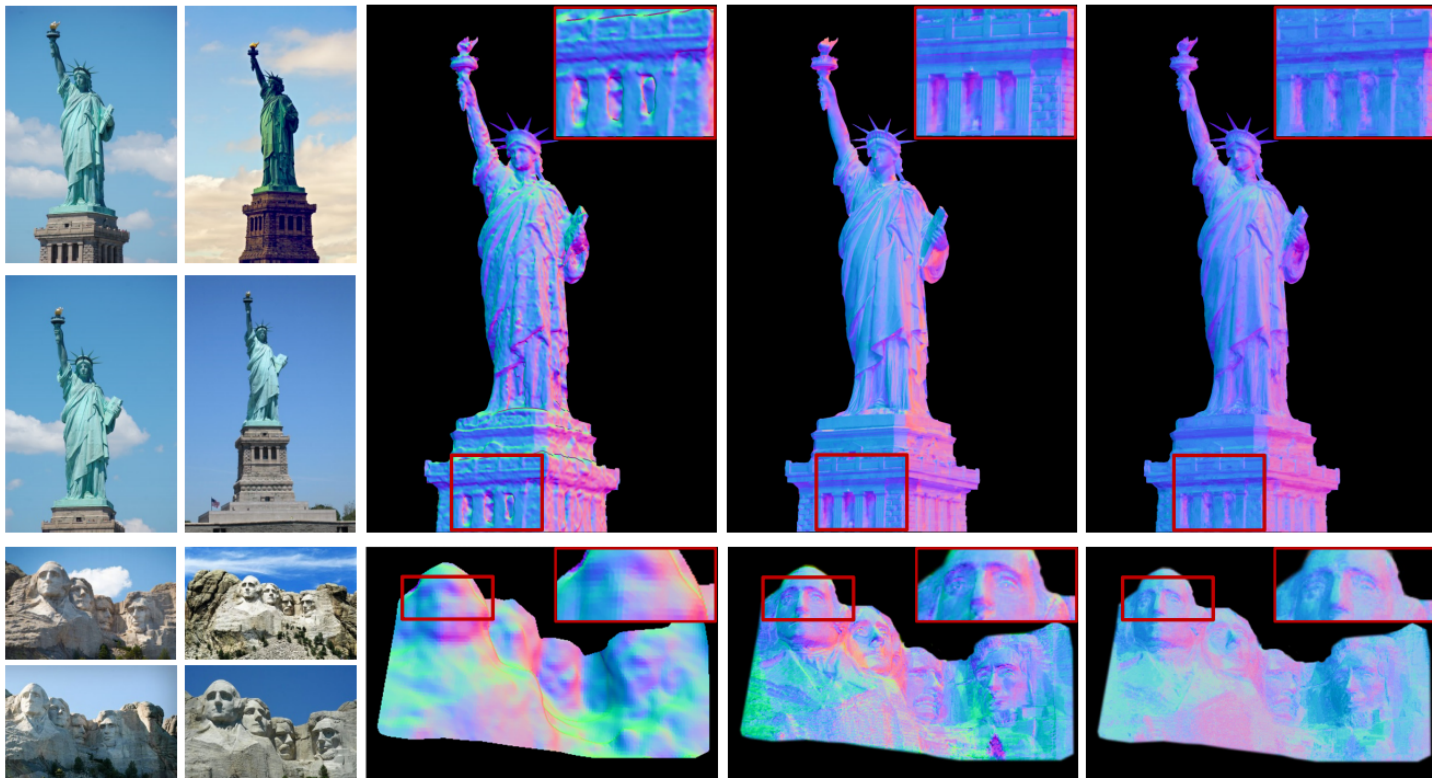
- Workarounds for many of the restrictive assumptions.
- Webcam photometric stereo:



Ackermann et al. 2012

Since 1994...

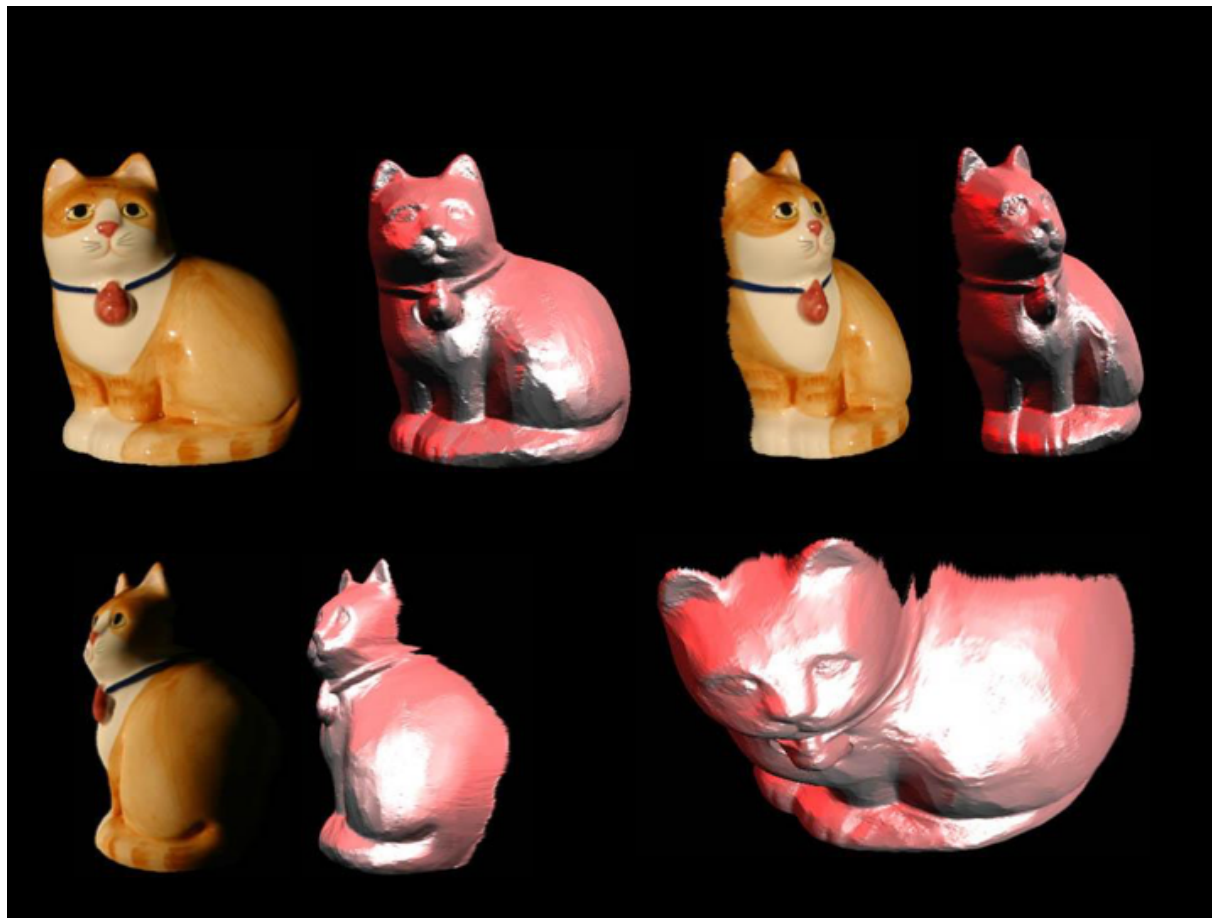
- Photometric stereo from unstructured photo collections (different cameras and viewpoints):



Shi et al, 2014

Since 1994...

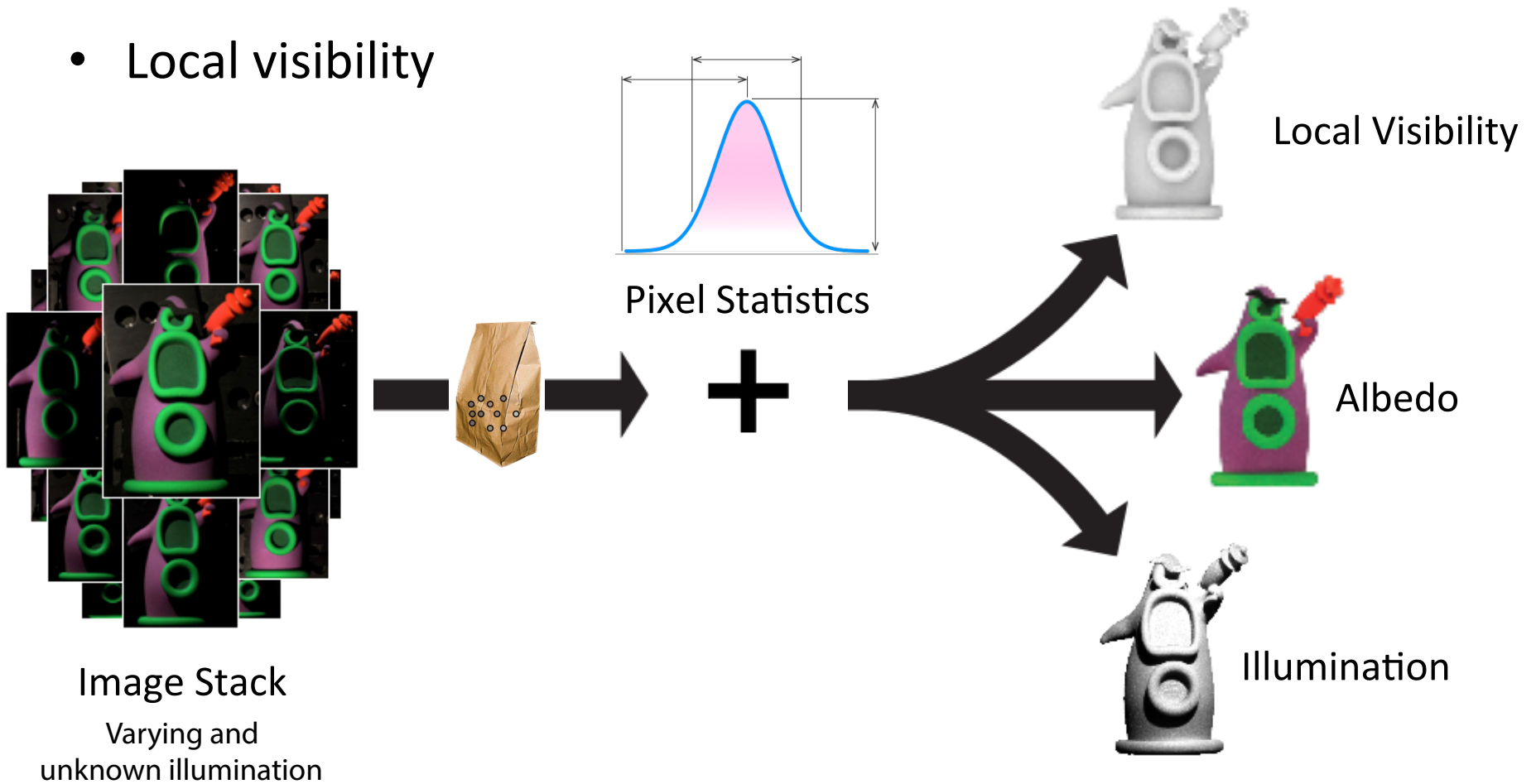
- Non-Lambertian (shiny) materials:



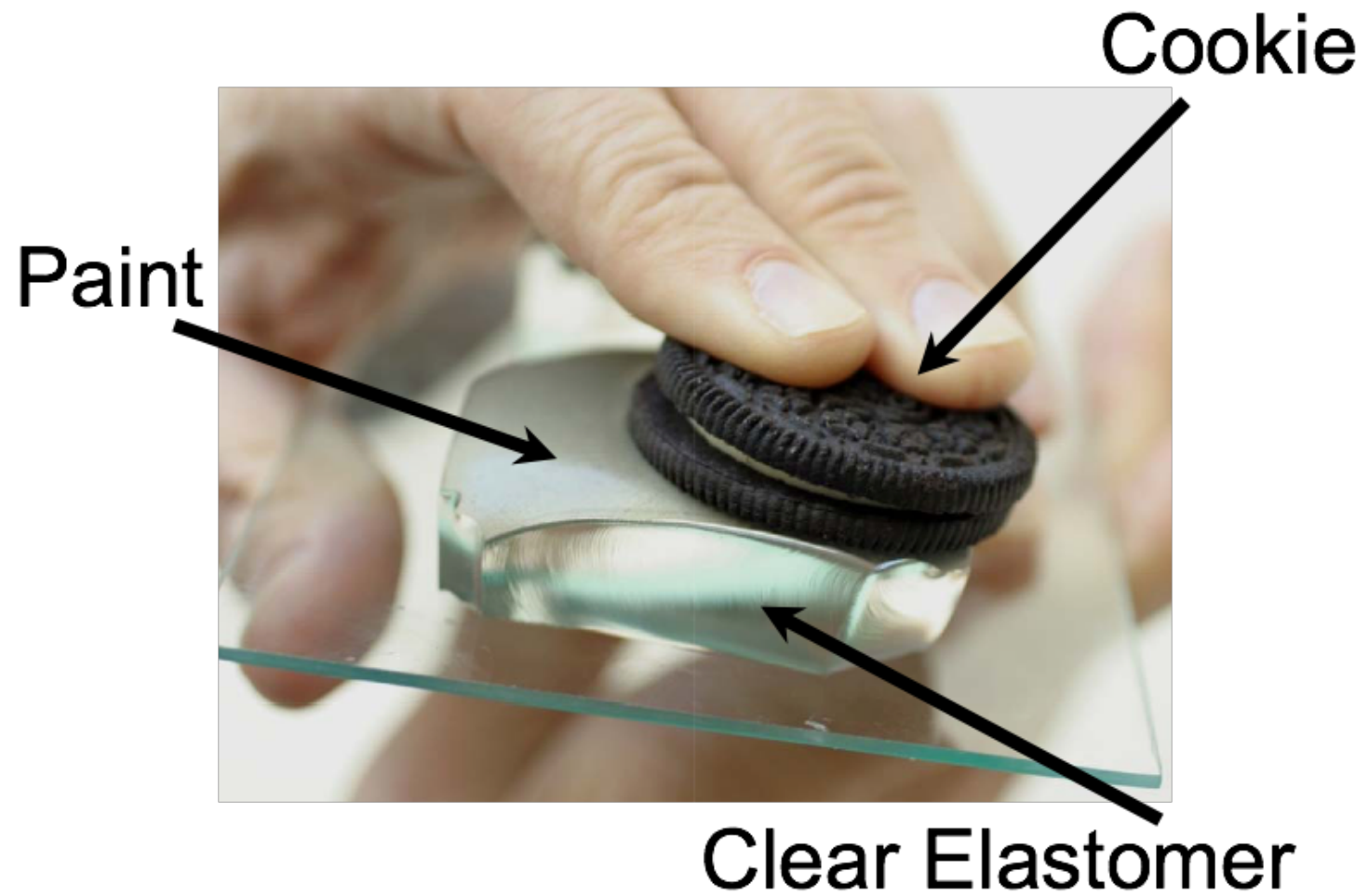
Hertzmann and Seitz, 2005

Since 1994...

- Local visibility



Hauagge, Wehrwein, Bala, Snavely, 2013



Johnson and Adelson, 2009



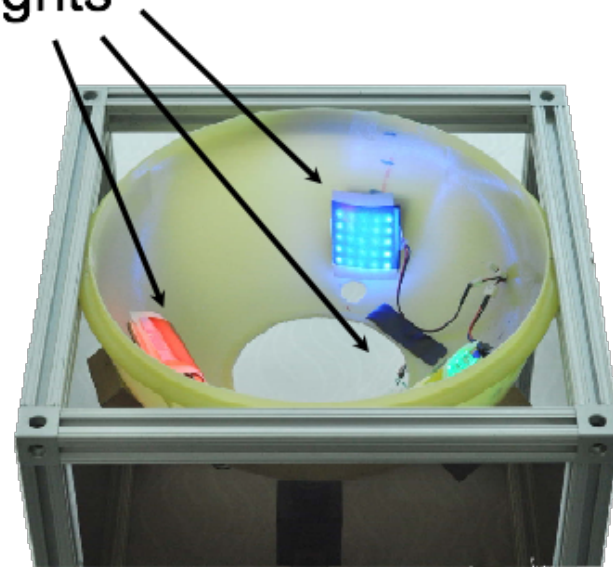


Lights, camera, action

Sensor

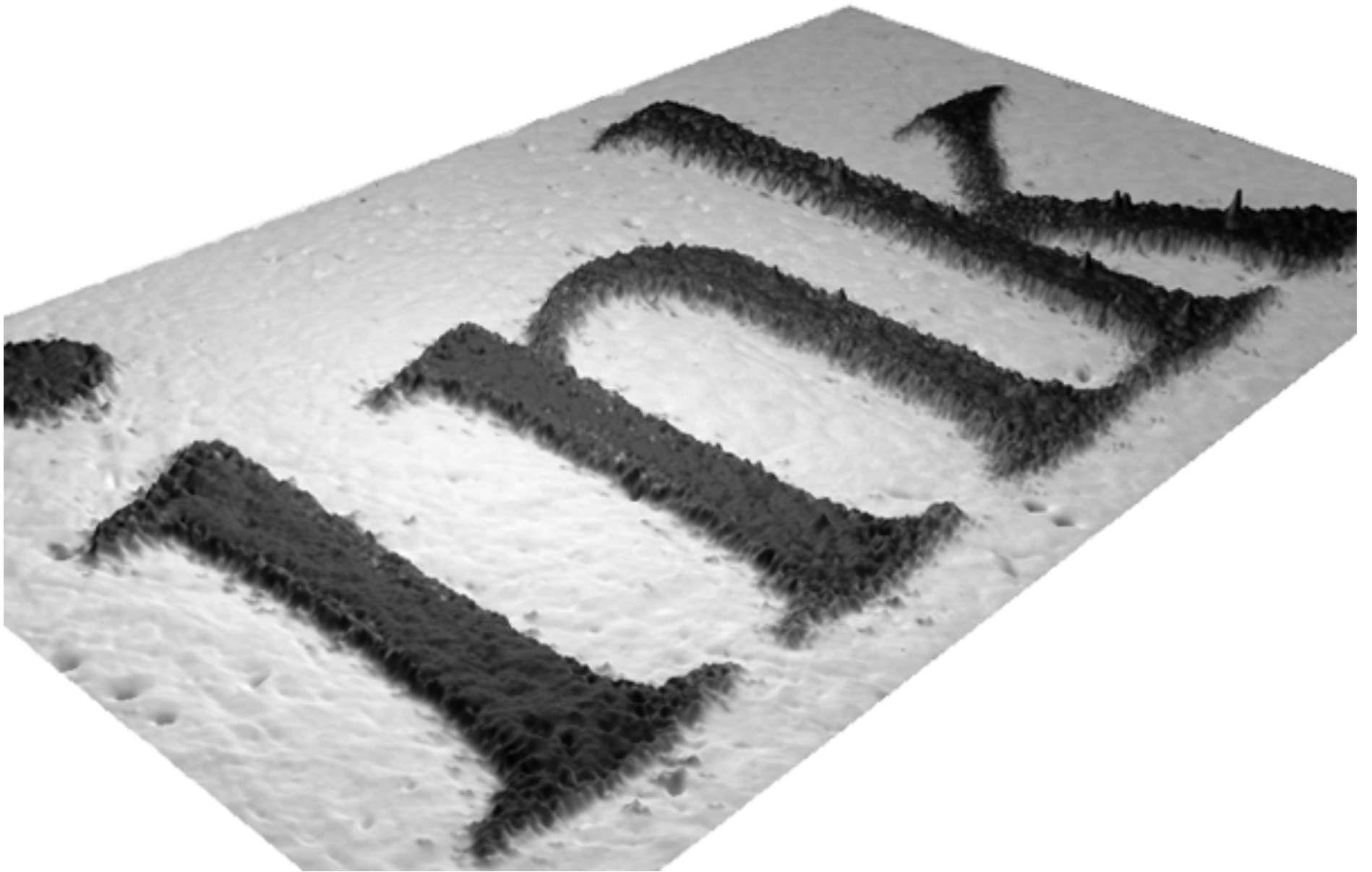


Lights



Camera





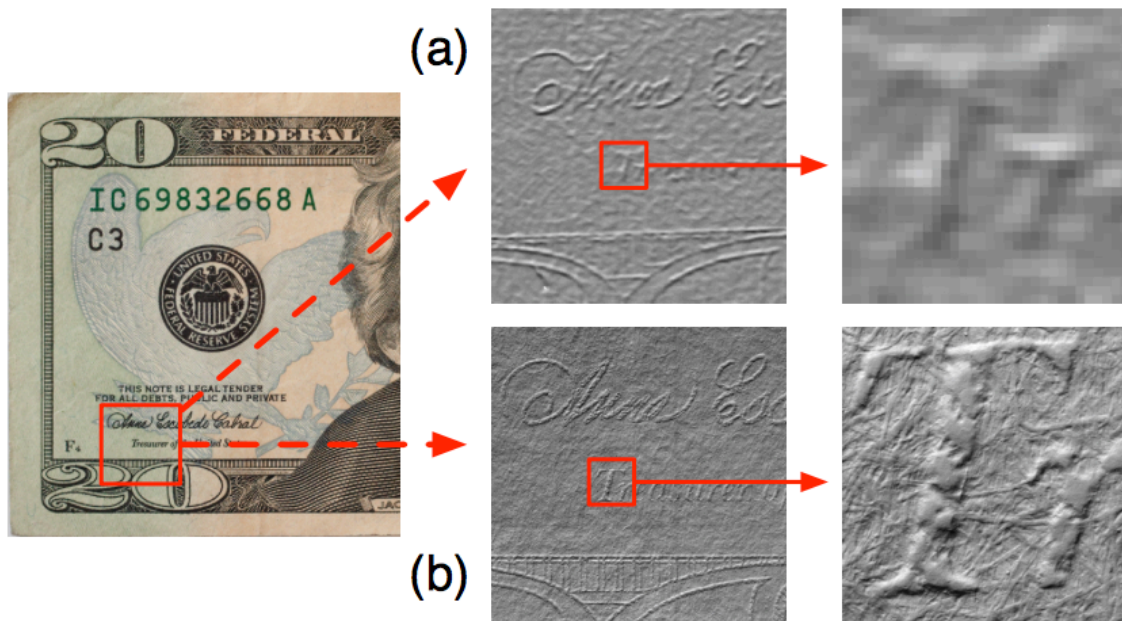
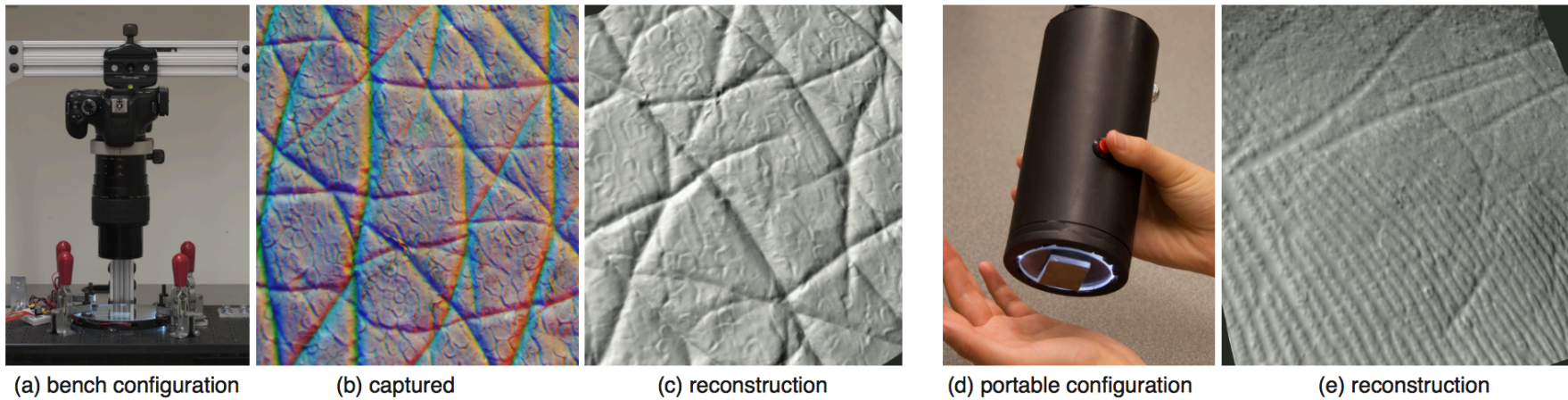


Figure 7: Comparison with the high-resolution result from the original retrographic sensor. (a) Rendering of the high-resolution \$20 bill example from the original retrographic sensor with a close-up view. (b) Rendering of the captured geometry using our method.

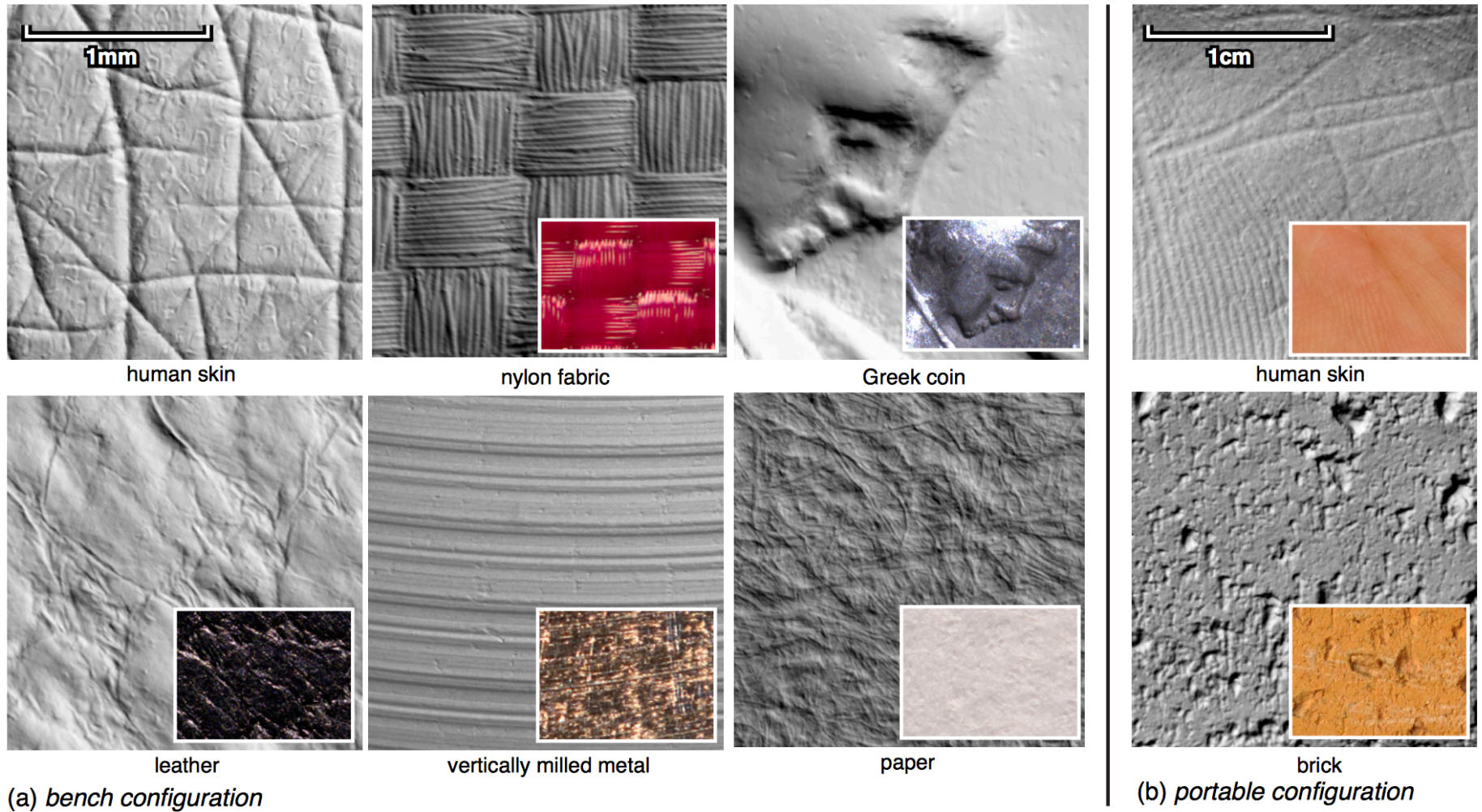


Figure 9: Example geometry measured with the bench and portable configurations. Outer image: rendering under direct lighting. Inset: macro photograph of original sample. Scale shown in upper left. Color images are shown for context and are to similar, but not exact scale.